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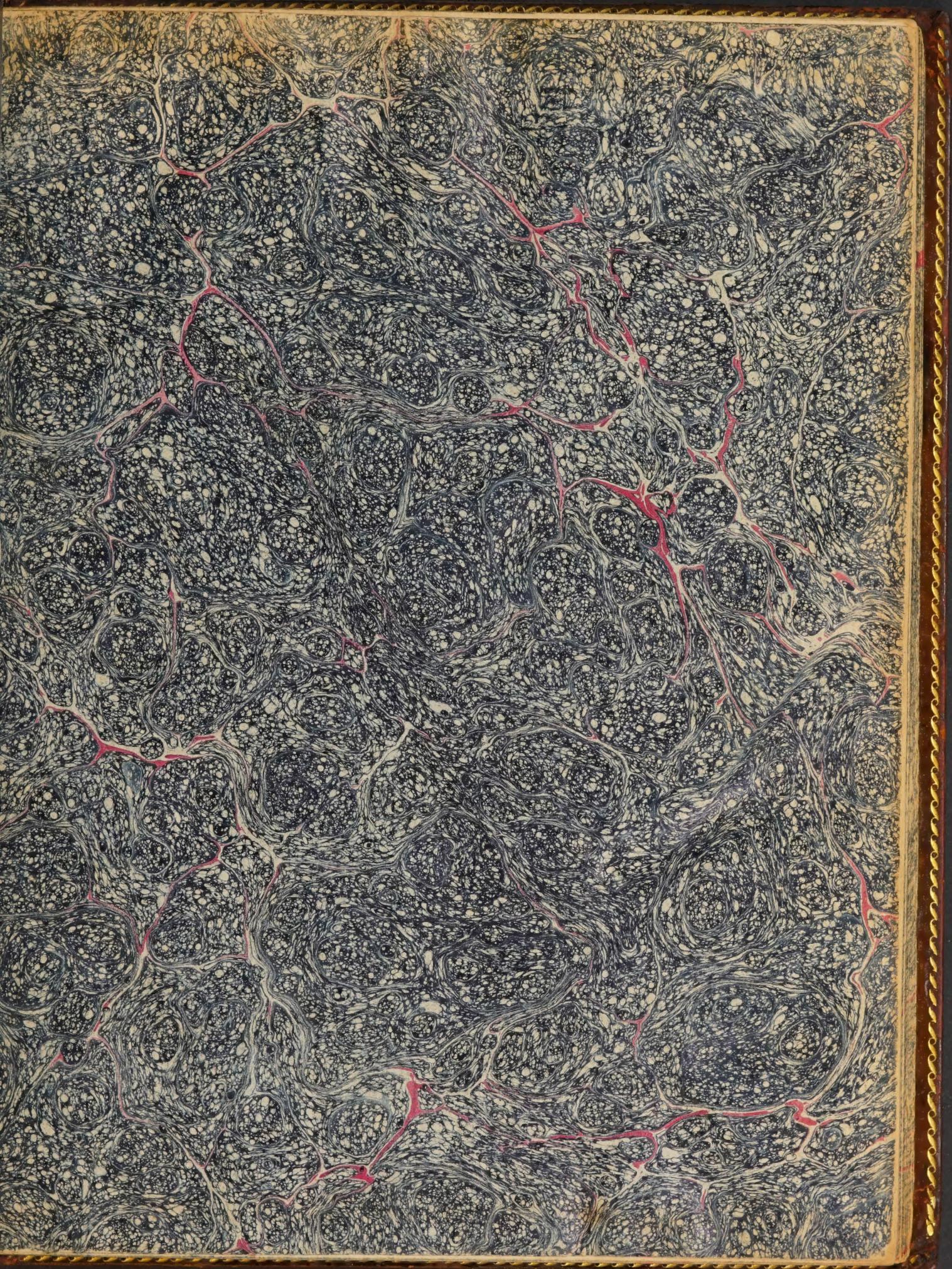
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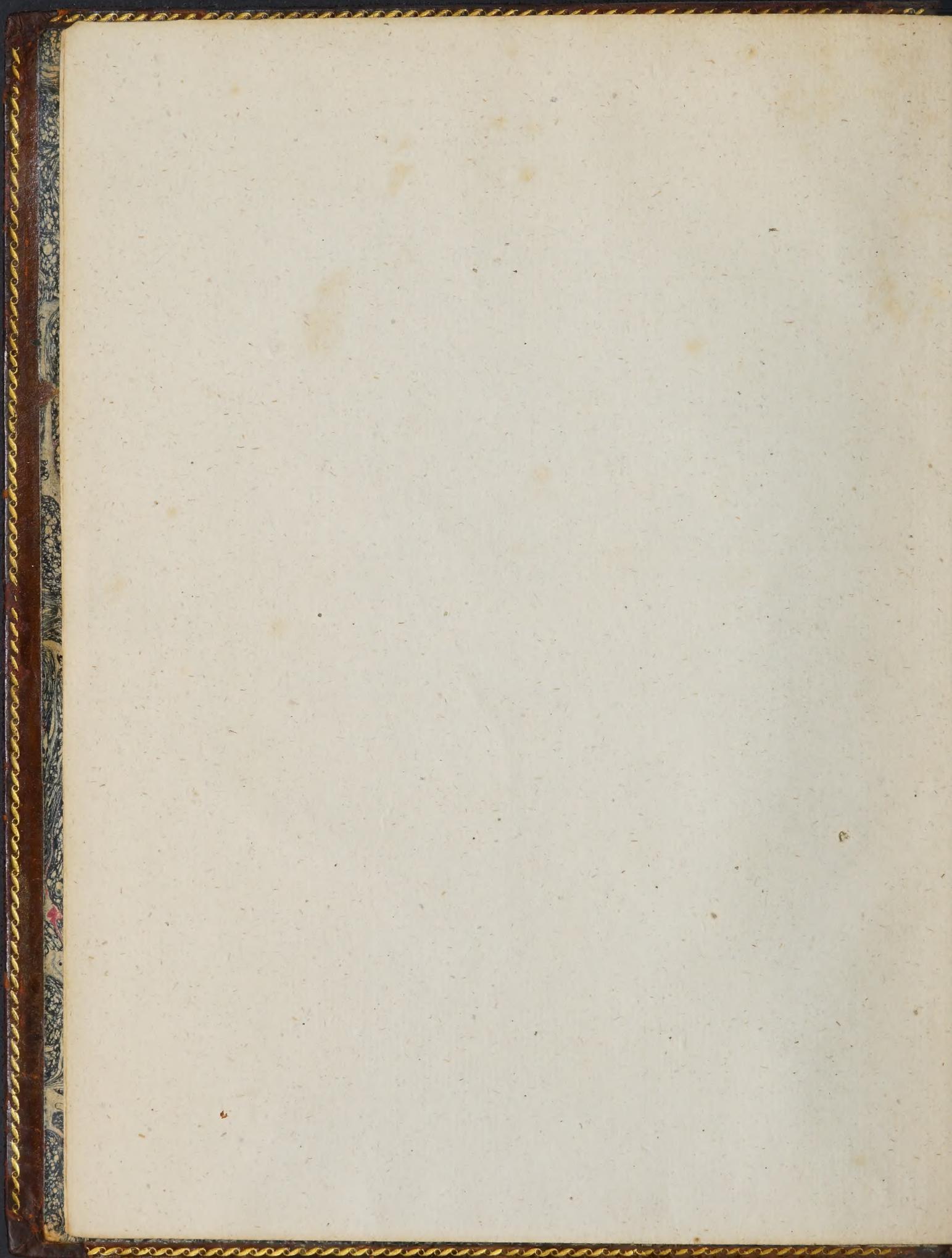
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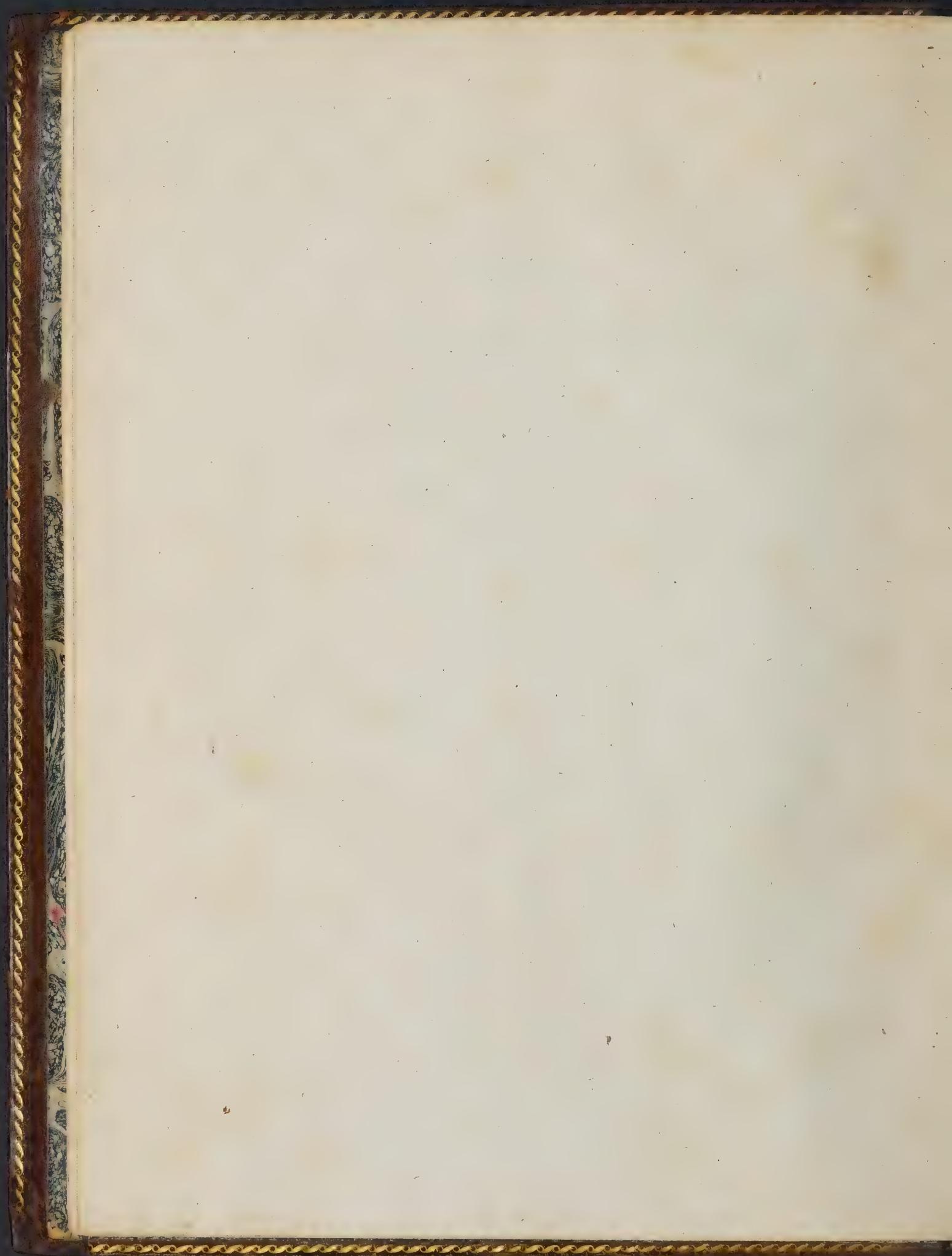
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O F  
M A R I T I M SURVEYING,  
I N T W O P A R T S:  
W I T H A  
P R E F A T O R Y E S S A Y  
O N  
D R A U G H T S A N D S U R V E Y S.

B Y  
M U R D O C H M A C K E N Z I E, Senior;  
Late MARITIM SURVEYOR in his MAJESTY's Service.

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## PREFATORY ESSAY

ON

### DRAUGHTS and SURVEYS of the SEA-COAST.

NO Branch of practical Geometry has been so little considered by Men of Science as Maritim Surveying. This Subject has never been particularly treated of by any Author, nor taught by a Master; nor have Surveyors given any Account of their Operations and Procedure in such Surveys. To this Reserve of Writers, and Silence of Practitioners, it may be ascribed, that an Art of such great Importance in Navigation has hitherto received so little Improvement; that in Practice little or no Distinction is made between Land-surveying and Coast-surveying, though they differ essentially from each other in their Nature and Circumstances; that so few Draughts are found to answer the End proposed by them; and that the Merits of the Draughts to which Seamen trust their Lives and Fortunes, are seldom judged of by any other Rules than the Recency of the Publication, the Neatness of the Engraving, the Authority under which the Survey was executed; or, sometimes, by the Rank and Reputation of the Persons to whom they are inscribed. Had Coast-surveying been considered with the Attention it deserves, the various Sorts of Draughts which are often indiscriminately made use of at Sea, would, before now, have been characteristically dis-

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tinguished from one another; the Particulars of which a complete Draught consists would have been sufficiently known; the several Methods of surveying compared together, and the most perfect Manner pointed out and explained. As yet, whoever undertakes a Survey of the Sea-coast is under the Necessity of devising a Plan of Procedure for himself, and performing most of the Operations by his own Ingenuity and Skill in the Principles, without the Advantage of the Speculations and Experience of others to assist him. This Want of previous Instructions, and the Difficulty and Trouble of frequent Investigation, seem to be the chief Reasons that Coast-surveyors commonly follow that Method which their antecedent Business or Applications suggest, rather than that which a thorough Knowledge of the Theory, and strict Attention to the Nature and Circumstances of the Coast would have directed. The Navigator, or experienced Seaman, constructs his Charts chiefly from Sea-journals of Courses and Distances: the Engineer and Land-measurer, by a continued Mensuration of Sides and Angles; or by a connected Chain of Triangles: Some measure their Distances with a Wheel, some with a Chain, and others with a Rope: one takes his Angles with a Needle, one with a Theodolite, and a third with a common Compass: the Way in which one proceeds in a Survey necessarily restricts his Performance within narrow Bounds: another takes a Method, by which he surveys more Coast in one Year, than he could travel over in three; surveying a little himself, and copying a great deal from others, without distinguishing the certain from the uncertain.

Wherever the Safety of Shipping is concerned, the Public has a Right to some Satisfaction with respect to the Nature and Grounds of the Publication. It is not sufficient to say, in the Title, that it is an actual Survey; or, a new and accurate Survey: it ought to be accompanied with, at least, a short Account of the fundamental Operations, and Manner in which the Survey was conducted. This would

would enable Seamen to form some Notion of the Merits of the Draughts they were trusting their Lives to, and to judge in what Circumstances they might be depended on, and in what Circumstances a Dependance on them would be imprudent. A faithful Answer to the following Questions would be sufficient for this Purpose, and might be easily subjoined to the Title of any Draught.

1. *How long was the fundamental Base-line? Where, and how, was it measured?*

2. *In what Manner was the Survey carried on from that Base-line? And with what Instruments were the Angles taken?*

Or, if the different Sorts of Draughts were classed under different Heads, with Epithets, or Appellations, significant of the Manner in which each was done, and published with that Epithet, or under that Appellation, only; this would in a great Measure answer the same End, and point out readily the Nature and Peculiarities of the several Draughts. Such a Capitulary is attempted in the following Enumeration of Draughts and Surveys; and will serve to explain what is here proposed.

1. A GENERAL, or NAUTICAL, CHART, is commonly constructed and drawn by some experienced Seaman from his own Observations and Journals of the Courses and Distances from Headland to Head-land, and Point to Point; with the intermediate Spaces filled up from such Charts or Maps as are at Hand, or are in most Repute. Sometimes on the Margin of the Chart are added larger Sketches of some of the most noted Harbours he is acquainted with, drawn from his own Memory, or copied from others. These Charts are always done by a very small Scale, so as to comprehend a large Extent of Land and Sea. Rocks and Shoals are seldom, or or never, particularly examined, but laid down according to their ap-

parent Distance from a noted Head-land, or by the Ship's Reckoning, which is looked upon as a sufficient Direction for avoiding them.

In these Charts, as all the Distances are either taken by the Eye, the Log-line, or the Ship's Traverse, without any certain Allowance for the Influence of Tides and Currents; and the Courses, for the most part, deduced from the Reckoning, or sometimes taken at Sea by a Compass; great Exactness is not to be expected: and as the Scale is very small, no striking Resemblance of any Part of the Coast can be expressed in the Draught, to point out with Certainty where Shoals or Harbours lie. Notwithstanding these Disadvantages, such general Charts are necessary to Navigation: and some take the Charge of conducting Ships who have no better Draughts to rely on: but then, if they are not particularly acquainted with the Places they are bound to, they ought always to carry a Pilot, who knows them; and never to trust to such Charts farther than for general Courses at free Sea.

When these general Draughts are corrected by Latitudes and Longitudes carefully observed; and when the Publisher is so just to himself and others, as to distinguish such Courses, Distances, and Tracts of Navigation as are delineated from his own repeated Observations and Experience, *from* those which he has copied, or taken on Trust from others; then they become very serviceable in sailing from Land to Land, or along an open Coast; but near the Shore, where the Danger is greatest; and in bad Weather, when a Ship is under the Necessity of running for a Harbour, or sheltering in a Bay, these general Charts are of little or no Service. In such Circumstances, larger and more particular Draughts, or a skilful Pilot, are indispensably necessary.

Charts of this Sort will not be brought to their utmost Perfection, till skilful Persons are sent on purpose, with good Instruments, to observe, on Land, or near it, the Latitudes and Longitudes of the most remarkable Points and Promontories which Ships have most frequent Occasion to sail along, or to *make*, when they approach the Land. If this shall ever be done, distant Views of the Land ought also to be taken; and each Observation, with its several Circumstances, should be preserved and published along with the Chart constructed from them. A few Years, perhaps, would be sufficient to accomplish this on both Sides of the Atlantic Ocean; or of the German, or North Sea: and would produce an accurate and lasting Chart of these Parts, which would never after need Correction. When a proper Manner of Surveying is more generally understood by seafaring Gentlemen, the intermediate Parts might be surveyed in a particular and complete Manner; which (excepting some Banks and Channels that are of a variable Nature) would be as permanent as the other. If, by the Assiduity of Theorists, or the Ingenuity of Artists, that important Problem of finding a Ship's Longitude at Sea, shall become easy to ordinary Seamen, such a Chart will be necessary to render that Discovery fully serviceable to Navigation.

2. ABBREVIATED COLLECTIONS OF DRAUGHTS, as they comprehend a great deal of Coast in a small Compass, are of the same Nature with the nautical Chart last mentioned, and can serve for no other Purposes in Navigation but to direct Ship's Courses from Land to Land, or from Promontory to Promontory along a Coast. These are commonly the Chamber-performances of Map-sellers and Drawers, who perhaps never saw any of the Places they delineate, and, for that Reason, cannot be supposed capable of making a proper Choice of the Draughts they abridge. Besides, in contracting a Diversity of Scales and Draughts into one, by the utmost Care they can hardly avoid introducing new Errors, or Inaccuracies, over and

above what might have been in the Originals before. By the Diminution of the Figure, the Distinctness that was in the Originals will certainly be lost in the Abbreviation. These miniature Collections often do more Prejudice than Service to Navigation: for, on account of the comparative Lowness of the Price, and the Names and specious Authorities annexed to the Title, many are inconsiderately led to purchase and preserve them; while the Originals, that would have been of much more Service, are as inconsiderately neglected and lost.

3. A MEMORIAL SKETCH is, a Delineation of a Harbour, or any Part of a Coast, from the Memory only; without Notes, or immediate Sight.

When such a Sketch is made by one who has been often in the Place, or who has viewed it with particular Attention, it may convey a general Idea of a Bay, Harbour, or Island, fit to gratify Curiosity, or enlarge Geography, by shewing that some such Places are there: but as every Part of the Sketch was at first by Guess, and that Guess delineated by Recollection only, the whole must be very uncertain, and of little Service to Shipping; farther than to admonish Sailors to keep a Look-out when they sail that Way. It is likely that most of the Islands in the *East Indies*, *Africa*, the *South-Sea*, and on the Coast of *Norway*, have been drawn and inserted in our Draughts after this Manner.

4. An EYE-SKETCH is, a Delineation of any Harbour, or Part of a Coast, done by the Eye at one Station, without measuring Distances; and drawn according to the apparent Shape and Dimensions of the Land.

Though here neither Bearings, Distances, the Figure of the Land, nor the Position of Rocks or Shoals, can be exact; yet a

good Drawer may make such a Sketch convey a sufficient Notion of a small Bay, Harbour, or Island, to be serviceable on some Occasions; and in moderate Weather, with the Help of the Lead, to direct a Vessel to the safest Channel and ordinary Anchorage. A great deal of the Benefit of such a Draught depends on chusing an advantageous Station to take it from, so that the principal Parts of the Harbour, or Bay, may not be too far from the Eye; for the remotest Parts will always be most unlike. It is likewise of Advantage to chuse the Station near the Sea, on that Side of the Bay, or Harbour, which is next the common Channel; because remarkable Parts on that Side of the Bay and Channel may then be delineated according to their true Bearings, by a Compass. Such a Sketch is soon made by one moderately skilled in Drawing: and if sea-faring Gentlemen, when they are in Places unknown, or not surveyed, would be at Pains to take such Sketches, it would be of some Service both to Navigation and Geography.

5. An AMBULATORY DRAUGHT, is made by walking along the Shore, taking the Bearings from Point to Point with a Compass, estimating their Distances by the Eye, and sketching the Figure of the Coast between them. If to this, remarkable Objects on Land are added, and such Rocks, Shoals, and Ledges as are visible at low Water, or by Breakers, with the Depths in the principal Channel, and in the Anchorage, the Draught will be a distinct Representation of any single Bay, or Harbour, and prove serviceable to Vessels on several Occasions. But if it is continued for some Leagues on an irregular Coast, the erroneous Distances will affect the Bearings, and also the Figure and Dimensions of the whole Draught: gross Errors in the more considerable Parts must then be corrected by making the less considerable Parts more erroneous; till at length the whole will become too evidently distorted and out of Proportion to admit of farther Alterations; and then of Necessity

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the Survey must be discontinued. It is likely that most of the particular Draughts used by Seamen have been made in this Manner.

When such Draughts are confined to small Parts of a Coast, or to single Bays, or Harbours, the Errors in them are less sensible; and they become serviceable by exhibiting plainly what Kind of Bays and Harbours are to be met with on that Coast: and, when Rocks, Shoals, and Soundings are inserted, they will be found a pretty good Direction for Vessels that may happen to fall in with them, provided at the same Time they are sure what Harbour it is which they have fallen in with: but to Strangers, and such as are not well acquainted with the Land before, separate Draughts of Harbours will not be found of great Use, especially in bad Weather, when a Direction is most needed; because of the Difficulty of finding where a particular Harbour lies, and the Danger of a Mistake in running for it. A Ship must be off the Mouth of the Harbour before such a Draught can be of Use; and the Pilot that is necessary to bring her thither, may likewise carry her in.

When one makes a Draught of a Bay, or Harbour in this Manner, at the same Time that he goes along the first Side delineating it, the magnetic Direction of the several Points and Heads on the other Side should be drawn out on the Paper, and their Positions limited to that Direction, which will render the Bearings right, however erroneous the Distances may be: if this is neglected, both the Distances and opposite Bearings will be false, and the Draught unlike in every Respect.

None of the three last-mentioned Performances ought to go under the Name of Surveys, because there is no Mensuration of Distances in them, nor any Thing determined in a geometrical Manner.

6. A DISJUNCT SURVEY is, when the Harbours, Bays, or Islands in any Country, are each surveyed separately in a geometrical Manner.

Manner. Or, when one, or more, straight Lines are measured on a Level Plane in each Harbour, Bay, or Island; the Angles taken with a good Instrument, and the Distances of the several Places determined from thence trigonometrically.

In this Survey, all the Rocks, Shoals and Channels are supposed to be carefully examined, and distinctly delineated, by a large Scale; and the Buoys, Beacons, Land-marks, remarkable Hills, Groves of Trees, Churches and Buildings, necessary for directing Vessels on any Occasion, inserted in their proper Places and miniature Forms; the various Diversities of the Sea-coast, and of the Rocks, Shoals and Sand-banks along it, represented so as these Varieties may be readily known by inspecting the Draught. Such a Draught will be a minute and exact Representation of each Place, as far as relates to Shipping; and is, or ought to be, the most complete of any, within the Limits of the Bay, Harbour, or Island surveyed. There is, however, one material Defect not to be avoided in Draughts of this Sort; which is, that Rocks, Shoals and Sand-banks that lie off the Harbour, or Coast, can seldom be laid down either in their just Dimensions, or in their proper Places; nor can intersecting Bearings be given for finding them out; nor Land-marks for avoiding them on all Sides. The Reason is, each Draught comprehends so small an Extent of Coast, that Objects properly situated for these Purposes cannot be included in it. *Lewis Morris*, in his Draughts of the Harbours of *Wales*, and Captain *Collins*, in his more distinct Draughts of Harbours in *Britain* and *Ireland*, very seldom give any Description of the Rocks and Shoals they represent, nor Directions for avoiding them on all Sides: and these they describe are so superficially done, that it is not without much Trouble and Searching that the Shoal can be found at all: the Nature of their Surveys would not admit of being so particular as that requires. Detached Draughts of small Extent, though never so exactly done, will not be found of much more Service to Navigation than the ambulatory Draughts before-mentioned:

for neither of them are of Use till a Vessel is in the Mouth of the Harbour; and then, the apparent Similitude between the ambulatory Draught and the Land, will direct her to the proper Anchorage as well as an exact Similarity in the geometrical Draught; for in sailing along, the Eye cannot perceive the Difference: No one that was not acquainted with the Coast, ever ventured to run for a Harbour by the detached Draughts that are to be seen on the Margins of some Maps and Charts; nor by those done by *Collins*, or *Morris*, though there is a general Chart along with them. However, disjunct Draughts, when well executed, give great Satisfaction, by making the Nature and Capacity of Bays and Harbours exactly and distinctly known, and are good Remembrancers to those who have been in these Harbours, and are so well acquainted as to know the Land at some Distance: but a Stranger to the Coast can receive very little Benefit by them, except that they may prove a Check on unskilful Pilots.

7. A LINEANGULAR SURVEY is, when the Coast is measured all along with a Chain, or Wheel, and the Angles taken at each Point and Turn of the Land with a Theodolite, or magnetic Needle

Some few Bays, that happen to be bordered with Sand or Marsh, may be suited to such Operations; and then, if great Care is taken a pretty exact Survey may be made. But, it is only a small Part of any Coast that is level and smooth enough to admit of a continued Mensuration with any tolerable Degree of Exactness. And, when it is considered that every Error, or Inaccuracy, in any one measured Distance, occasions a proportional Error in all the other subsequent Distances; that every Inaccuracy in the Angles affects all the other subsequent Angles and Sides; that, among such a Number of Angles, Inaccuracies, perhaps Mistakes, are hardly to be avoided; and when they happen, harder to be found out, so as to be properly corrected; when these Things are considered, it will not be reckoned

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a Paradox to affirm, that an irregular Coast may be surveyed by the Eye and a Compass, as exactly as in the lineangular Manner. A Surveyor will easily be satisfied of this, if he has ever tried to survey an irregular Field, or small Island, by measuring all the Sides, and observing the Angles between them: by his Field-book their Circumferences will never meet on the Paper. *That Reproach* of Surveyors is not to be wiped off by such a Procedure. Surveys taken in this Way are tedious, troublesome, and require many Hands constantly employed: the Distances by the Draught, and by observed Latitudes, differ widely from each other; and the whole will come out larger than the Truth; except Errors in the Angles shall happen to counteract the Errors in the Sides; and in that Case, particular Parts must be very much out of Proportion. If the Distances are measured with a Wheel, instead of a Chain; or the Angles taken with a Needle, instead of a Theodolite; then the whole, and particular Parts, will be still more erroneous. Lineangular Surveys that are continued beyond the Limits of a small Bay or Harbour, are in no Case to be depended on: and very few even of these, can be exactly done in that Manner. This, however, seems to be the most common Method of geometrical Surveying.

8. A TRIGONOCATENARY SURVEY is when one long Base-line is carefully measured on a level Plane, the Angles taken with a good Instrument, and from that Base-line a connected Series of Triangles carried on along the whole Coast, by which the Distances of the several Heads, Points and Flexures of the Shore are determined trigonometrically.

By this Procedure, one Source of Error in the former Method is avoided, viz. that which arises from a continued Mensuration: but the other, perhaps no less considerable, which is found, by constant Experience, to flow from a long Chain of Triangles all depending on one another, still remains to affect the mutual Bearings and Dis-

tances of the Places. This Manner of surveying is liable to much the same Inaccuracies and Imperfections as the lineangular, only not to so great a Degree. When Objects, sharp enough for forming the angular Points with Precision, are rarely to be met with (which is the Case on every Coast) Errors in Angles must necessarily arise: one Error causes another, that a third, and so on; each increasing the Inaccuracies of the subsequent Triangles: and, after they have insensibly run through a few trigonal Links of the Chain, their Source becomes indiscernible; and, at length, the Error of the whole too great to be corrected in the material Parts, without making very gross Errors in other Parts: the Survey must then of Necessity be either broke off, or continued with very striking Imperfections.

If a Coast happens to be so circumstanced, that level Planes may be met with on it at proper Distances from each other, and new Base-lines are measured on these Planes; and if Poles, or Signals, are set up forming Triangles all along the Coast; then, with great Care, the Angles may be taken, and the Survey continued with sufficient Exactness. But, such a Procedure would be extremely tedious: a Life-time would hardly be sufficient to survey the Coast of one Kingdom in this Manner, with the Care it requires. But as Planes of a sufficient Length and Smoothness are seldom to be found on any Coast, whoever carries on an extensive Survey this Way, will be often under the Necessity of making random Alterations in the Draught, at least at every new Base-line; that the several Parts may connect with one another, and that the whole may agree better with observed Latitudes. And, after the whole is put together one Way or another, the original Scale must be changed, and a different one substituted in its Place, that is better adapted to the Latitudes, and to the Dimensions of the whole: that is, a Survey taken in this Way must be finished, before the Scale that suits it is known.

9. An OROMETRIC SURVEY is, when one long Base-line is exactly measured, and the Distance of the Summits of two, or more, high Mountains in the Neighbourhood found from thence trigonometrically: then, by Angles taken on these Summits, the Distances of all the Points, Heads, Turns and other visible Objects on the Coast are ascertained; and from them the Positions of other intermediate Objects and Places. From these Mountains likewise, the Distances of other Mountains, farther along, are found; which serve in the same Manner as the first, for determining Distances along that Part of the Coast which is adjacent to them.

By this Way of surveying a Coast, when the Foundation is carefully laid, the Errors which all the fore-mentioned Methods are liable to, are in a great Measure guarded against: for, the Error, or Inaccuracy, of any one Distance, or Angle, is not communicated to the rest, but confined to that Distance, or Angle alone, or perhaps to an adjacent Point or two besides, which are not of great Consequence in the Draught, nor affect the Positions of other Parts. Therefore, in a mountainous Country, where the fundamental Triangles are few, and the Nature of the Coast otherwise favourable, a Survey may be continued a great Length, without any Error considerable enough to affect the Extent of the whole together. But then Shores are generally so low, and of such a Nature, that particular Parts can seldom be perceived from the Tops of two remote and high Mountains; even large Buildings are not often to be seen from thence: and though Points and Promontories, in Gross, may be perceived from both, yet it is scarce possible to distinguish the same Part of them from each Mountain. This will render the Angles, and of Consequence, the Distances inaccurate; and cause the observed Points and Heads in the Draught to extend more, or less, from the rest of the Coast than they ought: and other Distances and Bearings that depend on them will participate of their Error. However, if a moderate Degree of Care is taken, these Errors will

not

not be considerable, and the Draught, in general, may agree sufficiently with observed Latitudes and Longitudes. Though by this Method an extensive Survey may be carried on with more Certainty and Exactness than by any of those before-mentioned, yet there is Ground to believe that it has seldom been put in Practice.

10. A STASIMETRIC SURVEY is, when the mutual Distances of three, or more, proper Objects are carefully measured; and by Means of these Objects, the Position and Distance of all Stations along the Coast determined trigonometrically, each at its respective Station alone, independent of one another.

By this Manner of Surveying, the Errors which all the other Methods are liable to are avoided: for, the Distances along the Coast depending intirely on the Distances of three Objects chosen at Pleasure, the most proper may be selected for that Purpose: or, if proper Objects are not to be had, others may be adapted to it by Signals, and their mutual Distances previously determined to any Degree of Exactness. And if, by Oversight, an Error shall happen in any Part, it is confined to that Part alone, without affecting other Parts; and may be corrected at any Time independent of the rest. This Way of carrying on a geometrical Survey, whether inland or maritim, will appear evidently to be much more perfect than those commonly practised; and to surpass them, not only in Accuracy, but in Expedition; and is the only Method from which an exact extensive Survey is to be expected. Other Methods of surveying are defective either in their Foundation, or their Procedure, or in both: and their Errors, generally, slip insensibly into all the Parts of a Survey, increasing as it goes on. Thence it is, that so few Surveys have been continued beyond the Extent of a large Bay, or River; that the Errors of former Draughts are not to be corrected without an intire new Survey; that some Places have been surveyed again and again, yet in Proceſs of Time are found to require it anew; and that so few Draughts are to be met with that answer the End

for

## ON DRAUGHTS AND SURVEYS. xxiii

for which they are intended, viz. to direct Ships into safe Anchorage without a Pilot. The Design of the following Sheets is, to explain and recommend to more general Practice, the stasimetric Manner of surveying; to render future Draughts of the Sea-coast more complete and serviceable than what have been hitherto offered to the Public; and to direct sea-faring Gentlemen, who have gone through a Course of Navigation, how to make useful Draughts of such Bays or Harbours as they may occasionally be in; and by Means of such Instruments *only* as Seamen are commonly provided with.

USEFUL

## U S E F U L   N O T E S.

	Inches	Feet
The English Foot is	12	3
The Paris Foot	12,788	6
The Amsterdam Foot	11,172	
The Danish great Foot	12,465	
The Swedish Foot	11,692	

	Yards
An English Statute Mile is	1760
A Geometrical Mile	2038,6
A League, 3 Geometrical Miles	6115,8
A French League, called Lieu legale	3996,46
A Sea League of France	6394,375
The Dutch League of 18000 Rhinland Feet	6186
The common German Mile of 22800 Rhinland Feet	7829,666
The League of Sweden 30000 Rhinland Feet	10310

	English Miles
A Degree of the Meridian in Britain	69,5
A Degree at the Equator	67,75
A Degree at the Pole	70,1
The Chord of one Degree of the Meridian	69,150
The Chord of half a Degree	34,960
The Diameter of the Earth	7910, or, 41798117 Feet

The Sun's Semidiameter about the Equinoxes is	16'
About the Middle of Summer is	15', 45"
About the Middle of Winter is	16', 15"
The Sun's greatest Declination is	23°, 28', 9"

### Geometrical Miles in a Degree of Longitude.

Deg. of Lat. Miles and Parts.	60	59,77	59,09	57,95	56,28	54,38	51,96	49,15	45,96
Deg. of Lat. Miles and Parts.	45	50	55	60	65	70	75	80	85
	42,43	38,57	34,41	30,00	25,36	20,52	15,53	10,42	5,23

M A R I T I M

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# MARITIM SURVEYING.

## P A R T I.

*Containing the geometrical Principles and other Pre-requisites of Surveying.*

CHAP. I. Of the Principles and Pre-requisites in general. 2. Trigonometrical Solutions. 3. Longimetrical Operations and Problems. 4. Examination of Instruments. 5. Meridional Problems, and the Variation of the Needle.

## C H A P. I.

*Of the geometrical Principles, and other Pre-requisites of Surveying, in general.*

BEFORE one undertakes a geometrical Survey of the Sea-coast, it will be found of Advantage to have learned Drawing; at least so much of it as to be capable of shading with Indian-ink, or etching with a Pen, Hills, Rocks, Cliffs, miniature Buildings, and the Coast-line of a Map; of sketching, from Nature, the Out-lines of them; and of taking distant Views of the Land, from Sea. This can only be attained by long Practice, together with a few Instructions from a Drawing-master. A Surveyor who is not capable of sketching readily, and shading with some Neatness, ought to have a Drawer along with him for that Purpose.

He should also be acquainted with the Elements of Geometry and Astronomy; at least with as much as is taught in a common Course of Navigation; so as to be ready in protracting Lines and Angles, and calculating the Cases

of plain and spherical Trigonometry by logarithmic Tables, when the Proportions are stated; and applying the former to the Mensuration of Distances, and the latter to that of spherical Arcs and Angles; of taking the Altitude of the Sun, or a Star, with a Quadrant on Sea and Land, and from thence finding the Latitude of a Place; and of knowing by the Eye a few of the fixt Stars, especially such as are near the elevated Pole.

He should also make himself acquainted with the Structure of the principal instruments used in surveying; such as the Theodolite, the astronomical Quadrant, and Hadley's Octant, or Sextant; and understand how to examine, adjust, and observe with them. If he is defective in any of the above Particulars, his Operations and Observations will be tedious and inaccurate, or his Draughts void of Neatness and Elegance.

## C H A P. II.

### *Trigonometrical Solutions.*

THE Cases in plain Trigonometry that occur in surveying, are the four following.

1. Having two Angles of a Triangle, and a Side opposite to one of them, to find another Side; use the following Proportion.

*As the Sine of the Angle opposite to the Side given,  
Is to the Sine of the Angle opposite to the Side required;  
So is the Side given,  
To the Side required.*

2. Having two Sides and an Angle opposite to one of them, to find another Angle; use the following Proportion.

*As the Side opposite to the Angle given,  
Is to the Side opposite to the Angle required;  
So is the Sine of the Angle given,  
To the Sine of the Angle requir'd.*

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In the above Cases, it will assist the Memory to observe, that when a Side is wanted, the Proportion must begin with an Angle; and when an Angle is wanted, it must begin with a Side.

3. Having two Sides of a Triangle, and the included Angle, to find the other Angles, and Side; use the following Proportion.

*As the Sum of the two given Sides,*

*Is to their Difference,*

*So is the Tangent of half the Sum of the two unknown Angles,*

*To the Tangent of half their Difference.*

Then half their Difference (thus found) added to half their Sum, will be the greatest of the two Angles sought; which is the Angle opposite to the greatest Side. If the third Side is wanted, it may be found by Solution I. for two Angles are known, and a Side opposite to one of them.

4. The three Sides of a Triangle being given, from thence to find the Angles; proceed by the following Steps and Proportions.

### PLATE I. FIG. I.

1. Let  $A B C$  be the Triangle: make the longest Side  $A B$  the Base: from  $C$ , the Angle opposite to the Base, let fall the Perpendicular  $C D$ , on  $A B$ ; which will divide the Base into two Segments  $A D$  and  $D B$ . Then,

2. Find the Difference of  $A D$  and  $D B$ , by the following Proportion.

*As the Base  $A B$ ,*

*Is to the Sum of the other Sides ( $A C + B C$ )*

*So is the Difference of the Sides ( $A C - B C$ )*

*To the Difference of the Segments of the Base ( $A D - D B$ .)*

Half the Difference of these Segments (thus found) added to the half of  $A B$ , gives the greatest Segment  $A D$ ; or subtracted, leaves the least  $D B$ . Then,

3. In the Triangle  $A D C$ , by Sol. 2. find the Angle  $A C D$ : for the two Sides  $A D$  and  $A C$  are known, and the right Angle at  $D$  is opposite to one of them. The Complement of  $A C D$  is the Angle  $A$ . Then,

4. In the Triangle A B C, you have two Sides A B and B C, and the Angle at A opposite to one of them; from thence (by Sol. 2.) find the Angles C and B required.

## C H A P. III.

*Longimetricāl Operations and Problems.*

## S E C. I. O P E R A T I O N I.

*How to proceed in measuring a straight Line.*

BEFORE any Distance can be measured trigonometrically, the Length of one straight Line, or Distance, must be first found mechanically, by measuring it on a Plane with a Pole or Chain. And as this Mensuration is more or less exact, so will all the other Distances be that are deduced from it. The utmost care therefore ought to be taken in measuring this first Distance, or fundamental Base-line, exactly; and all Sources of Error in the Operation guarded against, as much as possible. The following Directions will be found sufficiently exact for any Survey.

1. Prepare a Measuring-pole 20 or 30 Feet long, precisely measured, square at one End, and a firm oblong Ring at the other; or a Chain 60 Feet long: three Poles, each 10 Feet long, with white, or dark-coloured Flags, to be tied to their Tops, according as they may appear on the Land, or between you and the Sky: 12 Poles, each 5 or 6 Feet long, with a small Flag on every fourth Pole: four iron Reels and Stakes, such as Gardners use, with about 200 Yards of small Cord to each Reel: 20 cylindrical Pins for counting the Number of Poles, or Chains, as you are measuring, each about 8 Inches long, and sharpened at the Point for piercing the Ground easily: two wide-mouth'd Canvas-pockets to carry the Pins in, with Strings for tying one of them round the Middle of each of the Measurers.

2. Pitch on a level Plane, as long as can be had; such as a smooth Sand, a Salt-marsh, or (which is by far the best) a frozen Lake.

In this Plane, chuse such a Direction for the Line to be measured, as that the principal Objects whose Distances are required, may be seen from both its

its Extremities ; and that the greatest Angle any of these Objects make with it, may not be too oblique (or not above 140 Degrees :) and, if convenient otherwise, let one End of this Line run directly on some remarkable Part of a remote Object ; so that, at the other End, its Direction may be readily found, at any time, by that Object.

4. At each End of the Line to be measured, set up a long Pole, with a white Flag flying at the Top of that which appears on the Land ; or a dark Flag on that which appears between you and the Sky, or on white Sand ; so that they may be easily seen from each other : about the Middle between these two set up the other long Pole, exactly in a Line with them : and in a right Line between these three set up shorter Poles all along, about 200 Yards asunder, with a Flag at the Top of every fourth or fifth Pole.

5. Stretch a Line, or Cord, from Pole to Pole, touching the same Side of each Pole as it passes them.

6. Immediately before you begin measuring along these Lines, try the Length of your Measuring-pole, or Chain, with a Six-foot mahogany Rod divided into Feet ; one of the Feet into Inches and Tents of an Inch ; the Thickness of the Wire of the Ring at the End of the Measuring-pole is not to be included in its Length : Count also the Pins, put them in the Canvas-pocket, and give them all to the Leader, or foremost Man, that carries the Pole.

7. Lay the Measuring-pole exactly along the Cords, with the Ring at the Part you begin the Measurement from, and its Square-end foremost ; at the Square-end, just touching it in the Middle, the Leader is to set down a Pin, and to go on with the Pole until the Follower (or he who carries the Ring-end) comes up to that Pin, and puts the Ring round it ; then the Leader laying the Pole exactly along the Cord, must thrust down another Pin at the Square-end. When that is done, he calls to the Follower, who must take up the Pin in the Ring, put it into the Canvas-pocket which is tied round his Middle, and go on to the next Pin, putting the Ring round it ; and when the Leader has put down another Pin, and called, then the Follower must take his Pin up, put it in the Pocket, and go on to the next,

as before ; proceeding in this manner till all the Pins are put down by one, and taken up by the other. Then leaving the Measuring-pole exactly placed along the Cord, without a Pin at any of its Ends, let the Pins be counted over, to be sure that none are wanting, and the Number marked down on Paper, and all of them given back to the Leader, to begin anew, by placing a Pin at his End of the Pole, as it lies along the Line, which the Follower is to put his Ring round, and take up as before, and so go on till the whole Distance proposed is measured.

The Use of the Ring at one End of the Pole, instead of laying one Pole at the End of another, is, to confine the Measurers to more Exactness ; which will be found a necessary Precaution for a Surveyor.

A Chain is not so exact for measuring with, when great Precision is required, as a Pole ; but is much more conveniently carried ; and, when Care is taken to lay it straight, and to prevent the Links from riding on one another, it will be found sufficiently exact on most Occasions.

8. If the Sand measured has a sensible and gradual Declivity, as from High-water Mark to Low-water, then the Length measured may be reduced to the horizontal Distance (which is the proper Distance) by making the perpendicular Rife of the Tide, one Side of a right-angled Triangle, the Distance measured along the Sand the Hypotenuse, and from thence finding the other Side trigonometrically, or by Protraction on Paper ; which will be the true Length of the Base-line. If the Plane measured is on the dry Land, and there is a sensible Declivity there, the Height of the Descent must be taken by a Spirit-level, or by a Quadrant, and that made the perpendicular Side of the Triangle.

If, in a Bay, one straight Line of a sufficient Length cannot be measured, let two, or three Lines, forming Angles with each other, like the Sides of a Polygon, be measured on the Sand along the Circuit of the Bay : these Angles carefully taken with a Theodolite, and exactly protracted, or calculated, will give the straight Distance between the two farthest Extremities of the first and last Line. In taking the Angles, let the Center of the Theodolite be placed exactly over the Point where the two Sides meet, by a Plum-

met, or by dropping a Stone down; and let Poles be set straight up in the Ends of the two Sides that form the Angle: for a small Error in the Angle, especially in the first, will often make a considerable Error in the Distance deduced from it.

Some Surveyors measure their Distances by a *Wheel*. But as a Wheel has a small serpentine Motion from Side to Side, occasioned by the Steps of the Person that leads it; and besides, traces all the little Inequalities of the Ground, which, however smooth to Appearance, no Plane (if it is not smooth Ice) is quite free from; this way of measuring cannot be reckoned sufficiently exact for a fundamental Base-line. Much less is any Distance measured by a Line, or Rope, to be esteemed sufficiently exact: for that will lengthen more, or less, by the Strength applied in stretching it along the Ground; and likewise lengthens, or shortens, by the various Degrees of Heat, or Moisture in the Air, or on the Plane measured.

## OPERATION II.

### *How to measure a straight Line on the Surface of the Sea.*

First, prepare a Measuring-line of strong Cord, two or three hundred Yards in Length, with small Pieces of Cork of equal Thickness made fast to it at small Distances, all along, like a Fishing-net, so that it may float straight on the Surface of the Water: if the Line has been well stretched, or much used before, it is the better: also prepare two Ropes somewhat longer than the greatest Depth of the Water to be measured, with a Pig of Lead or Iron Ballast (which we shall call an Anchor) 50 or 60 Pounds-weight, tied by the Middle to one End of each Rope, that when it is at the Bottom it may be able to anchor a Boat, and bear to be stretched straight without shifting the Place of the Anchor. Let the Measuring-line be thoroughly wet immediately before you begin to use it, and then stretched on the Water close by the Shore, and its Length measured there with a Pole. Then, in the Direction intended to be measured, take two remarkable sharp Objects on the Land in a Line, one near the Shore, the other as far up in the Country as you can: if such are not to be had, place Buoys on the Water at proper Distances in that Direction.

2. Take the Objects, or Buoys, in a Line, and holding one End of the Measuring-line fast on the Shore, carry out the other in a Boat, in that Direction, till it is stretched straight at its full Length by one Man in the Boat, and exactly at the End of the Line let another Man drop the Anchor, which will mark one Length of it. There keep the Boat, and the End of the Measuring-line, close to the Anchor-rope, drawn right up and down, till another Boat takes in the other End which was on the Shore, and rows farther on, and lays it straight in the Direction of the Land-marks, or Buoys, and there drops another Anchor, which will mark the second Length of the Measuring-line. Go on thus till the whole proposed Distance is measured; and immediately after let the Measuring-line be again measured with a Pole on the Water near the Shore, as at first, and if the Lengths differ, take the mean between them for the true Length. It is obvious, that to measure with any Exactness this way, the Sea must not only be smooth, but void of a Swell, and of all Stream of Tide; either of which will hinder the Line to lie straight. This Method of measuring a straight Line may be convenient on some Occasions; and if Care is taken to keep the Anchor-rope right up and down when the Measuring-line is applied to it, will be found sufficiently exact for many Purposes, but not for a fundamental Baseline from which other Distances are to be deduced.

There is another way of measuring a straight Line, mechanically, on the Sea, which is so well known to Seamen, that it is needless to describe it particularly here: and that is, by heaving the Log over a Ship's Stern while she is under sail, and observing how many Knots of the Log-line runs out in half a Minute; for the Line is so divided that the Ship will run (or, is supposed to run) so many miles in an Hour, in a straight Course; and twice as much in two Hours, and so on. But this Conclusion is founded on three Suppositions, neither of which are certain, viz. That the Log remains in the same Place during the whole half Minute that the Line is running out from the Ship's Stern; that the Ship continues to sail with the same Velocity, and also in the same Direction, during an Hour, or two, that she did during the half Minute: the contrary of which is more likely in most Cases. For the Log-line may shrink, or stretch, while it is running out; or may drag after the Vessel by the Weight of the Line, or by not running easily and readily off the Reel; the Swell of the Sea may alter the Place of the Log; and

Currents or Streams of Tide, stronger or weaker below the Surface than at it, an unsteady Helm, Lee-way, and varying Winds may change the Direction, or Celerity of the Ship's Motion; for neither of which can any certain Allowance be made. This way, therefore, of measuring a straight Line, or Distance, is not to be depended on as exact: but is mentioned here, because Rocks, Shoals, or Islands sometimes lie so far from the Coast, that there is no other way of forming any Notion of their Distance. If any such Distance is to be measured after this manner, let the Log-line be thoroughly wet when it is measured; let the Length between each Knot be 51 Feet, which is the Hundred-and-twentieth Part of a geometrical Mile, as half a Minute is  $\frac{1}{120}$ th Part of an Hour; chuse Neap-Tide, as much slack Water as can be got, and a moderate Breeze of following Wind; let the Line be run off the Reel so as never to be stretched quite straight; and if the Half-minute is measured by a Watch that shews Seconds, rather than by a Glass, it will generally be more exact. Perhaps one Second should be allowed for the Loss of Time in calling out at the Beginning, and stopping it at the End of the Time; except the Person who holds the Watch can contrive to observe the going out of the red Rag at the Beginning, and also to stop the Line himself at the End of the Time; which does not seem a difficult Matter.

## OPERATION III.

*To find the Distance of two Places, by the Flash and Report of a Gun.*

Stand in one of the Places with a Stop-watch in your Hand; the Watch stopt with the Second-hand at 60, and your Finger at the Stopper: let a Gun be fired at the other Place, at an appointed time. When you see the Flash, or Smoke, that instant set the Watch agoing; and immediately as you hear the Report, stop it; and it will shew the Number of Seconds of Time between the Flash and Report. By that Number multiply 1142, and the Product will be the Distance of the two Places in Feet: for Sound is found to move 1142 Feet in a Second.

This Operation should be repeated two or three Times, for more Certainty; and the exactest of them, if it is certainly known, fixed on for the true Time:

C

or

or a Mean of all, if no Distinction can be made. It should also be performed when there is little Wind ; for a Breeze of Wind, with or against the sound, is found to quicken or retard it a little. It is likely that some small Variation may also arise from different Degrees of Density, and Elasticity in the Air ; the former retarding, and the latter accelerating Sound. But this has not yet been ascertained by Experiments.

By this way of measuring Distances, such as are not more than two Miles cannot be reckoned exact ; because of the Difficulty of setting the Watch agoing, and stopping it at the precise Times ; which may cause an Error of, at least, one Quarter of a Second or 90 Yards, which is too much in so short a Distance. If the Distance is four Miles, or more, it may be reckoned sufficiently exact for ordinary Purposes. I incline to think, if half a Second is allowed for the Time lost in stopping the Watch and setting it agoing, that, after a little Practice, one may measure two Miles pretty exactly by this Method ; provided the State of the Air is such as it was when the Experiments for determining the Velocity of Sound were made.

It is to be wished that those Experiments were repeated over again by some careful Person, who would attend to the several Circumstances of the Air and Wind that might affect the Celerity of Sound, and publish the Effect of each, for the Conveniency of Surveyors, and the farther Satisfaction of the Curious.

## S E C T I O N II.

### *Longimetrical Problems.*

#### P R O B. I.

*Having the Length of a Base-line X Y, to find the Distance of two, or more remote Objects A, B and C, from it, and from each other.*

#### P L A T E I. FIG. II.

#### B Y P R O T R A C T I O N.

With a Theodolite, or Hadley's Quadrant, at one End of the Base-line X, take the Angles Y X A, Y X B, Y X C ; at the other End Y, take the Angles

Angles  $YXA$ ,  $YXB$ ,  $YXC$ , and write them down in a Book. Then draw a Line on the Paper of a sufficient Length, and in a convenient Position: pitch on any convenient Part of that Line for the Point  $X$ ; and taking the Length of  $XY$  from a Scale of equal Parts, set it off from  $X$  to  $Y$ : at  $X$  draw the observed Angle  $YXA$ ; and at the other End  $Y$ , draw the observed Angle  $XYA$ , and the Intersection of these two Lines will be the Point  $A$ ; and  $XA$ , measured on the Scale from which  $XY$  was taken, will be the Distance of  $A$  from  $X$ . In the same manner the Points  $B$ , or  $C$ , may be determined; and the Length of  $AB$ , or  $BC$ , measured on the Scale, will give the Distance of  $A$  from  $B$ , or  $B$  from  $C$ : and so for any other Points whose Distances are required, whether they lie on one Side of the Base-line or the other.

Let it be observed, that if  $XY$  is drawn on the Paper according to its magnetic Direction, then the Points  $A$ ,  $B$ , and  $C$ , will likewise be in their magnetic Directions from  $X$  and  $Y$ , and from each other.

To render protracted Distances more exact, the Length of the Base-line should not be less than one seventh, or one eighth Part, of the Distance of the farthest Object; none of the Angles should be too oblique, or above  $140$  Degrees; the Protractor should be at least  $12$  Inches Diameter, and each Degree subdivided into four Parts; or more minutely by an Index with a vernier Scale on it.

#### BY CALCULATION.

In the Triangle  $XYA$  there are given the measured Side  $XY$ , and the observed Angles  $YXA$  and  $XYA$ , and therefore the Angle  $YAX$ , their Compliment to  $180^\circ$ ; that is, two Angles and a Side opposite to one of them: therefore, (by Trigon. Sol. 1. Page 2.) the Side  $XA$  may be found. In the same manner, in the Triangle  $XYB$ , may  $XB$  or  $YB$  be found. Then, in the Triangle  $AXB$ , there are given the two Sides  $XA$  and  $XB$ , and the included Angle (the Difference of the observed Angles  $YXA$  and  $YXB$ ); hence  $AB$  may be found, by Trigon. Sol. 3. Page 3.

## PROB. II.

*To find the Distance of two Places, when the Latitudes, their Bearing by the magnetic Needle, and the Variation of the Needle are known.*

## PLATE I. FIG. III.

## BY PROTRACTION.

1. Draw the Line X A right up and down on the Paper, for the magnetic Meridian.

2. Pitch on a convenient Point X in that Line, to represent one of the Places; and at the Point X make the Angle A X B equal to the Variation, and on the East-side of X A, if the Variation is West; but on the West-side of it, if the Variation is East; and X B will be the true Meridian.

3. At X make the Angle A X Y equal to the magnetic Bearing of the other Place from X.

4. Take from a Scale of equal Parts the Difference of Latitude of the two Places in geometrical Miles, and set it off on the true Meridian from X to B.

5. At B raise a Perpendicular B Y cutting X Y in Y; and the Point Y will be the other Place; and X Y, measured on the Scale, will be the Distance of the two Places in geometrical Miles.

If their Distance is wanted in English Miles, say, as 60 is to 69, 5 so is their Distance in geometrical Miles, to their Distance in English Miles: for 69 and a half English Miles make 60 geometrical Miles.

## BY CALCULATION.

In the right-angled Triangle Y B X, there is given the Angle Y X B, in this Case, equal to the Bearing of Y and Variation of the Needle together;

and therefore the Angle at Y is also known; and the Side X B opposite to one of them: hence (by Trigon. Sol. 1.) X Y is found.

In this Problem, the greater the Difference of Latitude of the two Places is, and the nearer they are to the same Meridian, their Distance will come out the more exact. For, an Error of half a Minute can hardly be avoided in the Latitudes, though taken with the best portable Quadrant; nor an Error of 15 Minutes, sometimes, in the Bearing of Y from X, though taken with the best Needle: more especially as the Polarity of Needles is found to vary, in the same Place, above 20 Minutes in a Day, according to the State of the Atmosphere, as is thought. To be sufficiently exact, therefore, the Difference of Latitude of the Places should be no less than 15 Miles, and the Angle B X Y no greater than 45 Degrees.

The Accuracy of this Problem depends chiefly on the Accuracy of the two Latitudes: for supposing the Needle to be twenty Minutes wrong, it will not occasion an Error of a Quarter of a Mile in a Distance of thirty Miles, if the Intersection of the Bearings is not oblique: and therefore, if the Quadrant with which the Latitudes are found, and the Observations likewise, are not known to be good, the Distance deduced from them is not to be depended on as exact. In calculating the Latitudes by the Sun's Meridian Altitude, remember to make the proper Allowance for the Sun's Semidiameter, (at a Medium 16', as in the Table, Page 17,) the Refraction, Variation of the Sun's Declination for the Time between your Meridian and the Meridian of the Tables; and if the Altitude is taken with Hadley's Quadrant, by the visible Horizon, make Allowance also for the Depression of the Horizon, or Altitude of the Eye above the Surface of the Sea; according to the Table in Page 18.

Here it may be observed, that the Distance between X and Y, thus found, is the Length of an Arc of a great Circle of the Earth between these two Stations, and not the straight Line, or shortest Distance. The Chord of that Arc is the true Distance to be used in Surveying. Therefore instead of X Y measured by two Latitudes, take its Chord, which is twice the Sine of half the Arc. The Proportions between the following Arcs and their Chords were found by the Tables of natural Sines by saying, As

the

the Radius of the Tables, is to the natural Sine of half the given Arc, so is the Radius of the Earth to a fourth Number; which doubled, is the Chord required. If X Y therefore is found by two Latitudes to be  $69\frac{1}{2}$  English Miles, the Chord of that Arc, or shortest Distance between X and Y is only 69 Miles and 150 Yards: if the Arc is  $34\frac{3}{4}$  Miles, its Chord is  $34\frac{1}{2}$  Miles and 80 Yards: if the Arc is 20 Miles, its Chord will be  $19\frac{3}{4}$  Miles and 264 Yards nearly. If the Arc measured is less than twenty Miles, its Chord may be reckoned the same with the Arc.

If, in Practice, it shall happen that Y cannot be seen from X; or that, because of their Distance, both Points cannot be included in the same Draught; in these Cases, to avoid the Trouble and Loss of Time in finding the Latitude of another Place instead of X, pitch on some other Station nearer Y, as S, whose Distance and Direction from X is determined on the Draught, and where Y can be seen: find its Latitude, by letting fall the Perpendicular S b on the true Meridian drawn through X; and b X measured on the Scale and reduced to geometrical Miles, will be the Difference of Latitude between X and S: which added to, or subtracted from, the Latitude of X (as S is northward or southward of it) will be the Latitude of S. Then, having the Latitude of S and of Y, find their Distance, as before directed.

*To find how much the Sun's Declination, at any Place, differs from the Declination in the Table.*

Judge, as near as you can, what the Difference of Longitude is, in Time, between the Place of Observation, and the Place for which the Table of Declination is calculated. Then, subtract the Declination for the Day of Observation, and for the Day next following, from each other, and the Remainder will be the Variation of Declination for one Day, or 24 Hours; and allow a twenty-fourth Part of that Variation for every Hour between your Meridian and the Meridian of the Table of Declination. For Example: suppose the Variation of the Sun's Declination in 24 Hours, is found by the Table to be ten Minutes; and the Longitude of the Place of Observation, from the Place for which the Table of Declination is calculated, is 45 Degrees, or 3 Hours; then state the Proportion thus,

As

As 24 Hours : is to  $10' ::$  so is 3 Hours : to  $1', 15''$ . Which  $1', 15''$  is to be added to, or subtracted from, the Declination in the Table, according to the Circumstances following.

*When the Sun's Declination is increasing* : if you are eastward of the Meridian of the Table, the Remainder must be subtracted from the Declination in the Table: but if you are westward of the Meridian of the Table, it must be added, to give the Sun's Declination at the Place of Observation.

*When the Sun's Declination is decreasing* : if you are eastward of the Meridian of the Table, the Remainder must be added to the Declination in the Table: if you are westward, it must be subtracted from it, to give the Declination at the Place of Observation.

Within three, or four, Days of the Solstices, an Error of 30, or 40 Degrees in the Longitude of the Places, will make no sensible Error in the Sun's Declination: but when the Sun is within 10 Degrees of the Equinoxes, to be sufficiently exact in the Declination, the Longitude should be known to half an Hour, or 8 Degrees.

*A sure and ready Way to find, whether the Sun's Declination, or Zenith Distance is to be added or subtracted, to produce the Latitude.*

Let it be observed, that if the Sun's Place in the Ecliptic is transferred to the Earth, or, marked by a Point on the Meridian of a Place; then, the Sun's Meridian Zenith Distance, obtained by the Quadrant, is equal to the Distance of that Point from the Place of Observation: and therefore, when the Distance of the Sun (or the Point which represents it) from the Equator, is known, the Distance of the Place of Observation from the Equator will likewise be known.

#### PLATE I. FIG. IV.

If therefore a Circle N E P Q is made by Hand, representing a Meridian on the Earth, and a Line E Q, for the Equator, drawn through the Middle of it; then a Point S, representing the Sun's Place, marked by Guess on the Circle;

Circle; but on the North, or South Side of the Equator-line, as the Sun's Declination is North or South; and another Point Z, for the Place of Observation, North from S, if the Sun is South at Mid-day; but South of S, if the Sun is North at Mid-day: then by inspecting that Figure, it will be easy to see whether the Declination, or Zenith Distance, is to be added or subtracted from each other, to give Z E, the Latitude, or Distance of the Place from the Equator. In this Case, it is obvious, that the Declination S E must be added to Z S the Zenith Distance, to give Z E the Distance of the Place from the Equator. Again, let s be the Sun's Place, then Z s is the Zenith Distance, from which s E must be subtracted to give Z E the Distance from the Equator. Again, if s is the Sun's Place, and z the Place of Observation, the Sun being supposed North at Mid-day; then z s is the Zenith Distance, to which s E, the Declination, must be added, to give z E, the Distance of z from the Equator. If z falls between s and E, then the Zenith Distance must be subtracted from the Declination to give the Latitude: as will be manifest by marking the Points properly on the Figure.

## P R O B. III.

*To find the Latitude of a Place which is inaccessible.*

First find the Variation of the Needle; then, by the Direction of the Needle, pitch on a convenient Place which bears due East, or due West, from the inaccessible Place whose Latitude is required; there find the Latitude, and that will be the Latitude of the other likewise: for both are in the same Parallel.

The nearer the Place of Observation is to the other, the more exactly will the Latitudes agree; because they will be less affected by any Inaccuracy in the Bearing by the Needle.

A TABLE

A TABLE of the Refraction of the Heavenly Bodies  
in Altitude.

App. Alt.	Refrac.	App. Alt.	Refrac.	App. Alt.	Refrac.	App. Alt.	Refrac.	App. Alt.	Refrac.
°	'	°	'	°	'	°	'	°	'
"	"	"	"	"	"	"	"	"	"
0. 0	33. 0	3. 30	13. 6	7. 40	6. 45	13. 0	4. 3	27. 0	1. 51
5. 32.	10.	3. 40	12. 40	7. 50	6. 37	13. 20	3. 57	28.	1. 47
10. 31.	22.	3. 50	12. 15	8. 0	6. 29	13. 40	3. 51	29.	1. 42
15. 30.	35.	4. 0	11. 51	8. 10	6. 22	14. 0	3. 45	30.	1. 38
20. 29.	50.	4. 10	11. 29	8. 20	6. 15	14. 20	3. 40	31.	1. 35
30. 28.	22.	4. 20	11. 8	8. 30	6. 8	14. 40	3. 35	32.	1. 31
32. 28.	5.	4. 30	10. 48	8. 40	6. 1	15. 0	3. 30	33.	1. 28
36. 27.	30.	4. 40	10. 29	8. 50	5. 55	15. 30	3. 24	34.	1. 24
40. 27.	0.	4. 50	10. 11.	9. 0	5. 48	16. 0	3. 17	35.	1. 21
50. 25.	42.	5. 0	9. 54.	9. 10	5. 42	16. 30	3. 10	36.	1. 18
I. 0. 24.	29.	5. 10	9. 38	9. 20	5. 36	17. 0	3. 4	37.	1. 16
I. 10. 23.	20.	5. 20	9. 23	9. 30	5. 31	17. 30	2. 59	38.	1. 13
I. 20. 22.	15.	5. 30	9. 8	9. 40	5. 25	18. 0	2. 54	39.	1. 10
I. 30. 21.	15.	5. 40	8. 54	9. 50	5. 20	18. 30	2. 49	40.	1. 8
I. 40. 20.	18.	5. 50	8. 41	10. 0	5. 15	19. 0	2. 44	41.	1. 5
I. 50. 19.	25.	6. 0	8. 28	10. 15	5. 7	19. 30	2. 39	42.	1. 3
2. 0. 18.	35.	6. 10	8. 15	10. 30	5. 0	20. 0	2. 35	43.	1. 1
2. 10. 17.	48.	6. 20	8. 3	10. 45	4. 53	20. 30	2. 31	44.	59.
2. 20. 17.	4.	6. 30	7. 51	11. 0	4. 47	21. 0	2. 27	45.	57.
2. 30. 16.	24.	6. 40	7. 40	11. 15	4. 40	21. 30	2. 24	46.	55.
2. 40. 15.	45.	6. 50	7. 30	11. 30	4. 34	22. 0	2. 20	47.	53.
2. 50. 15.	9.	7. 0	7. 20	11. 45	4. 29	23. 0	2. 14	48.	51.
3. 0. 14.	36.	7. 10	7. 11	12. 0	4. 23	24. 0	2. 7	49.	49.
3. 10. 14.	4.	7. 20	7. 2	12. 20	4. 16	25. 0	2. 2	50.	48.
3. 20. 13.	34.	7. 30	6. 53	12. 40	4. 9	26. 0	1. 56	51.	46.
52.	44. 59.	34. 67.		24. 75.		15. 83.		7.	
53.	43. 60.	33. 68.		23. 76.		14. 84.		6.	
54.	41. 61.	32. 69.		22. 77.		13. 85.		5.	
55.	40. 62.	30. 70.		21. 78.		12. 86.		4.	
56.	38. 63.	29. 71.		19. 79.		11. 87.		3.	
57.	37. 64.	28. 72.		18. 80.		10. 88.		2.	
58.	35. 65.	26. 73.		17. 81.		9. 89.		1.	
	66.	25. 74.		16. 82.		8. 90.		0.	

A T A B L E of the Depression, or Dip, of the Horizon of the Sea.

Height of Depression the Eye a- of the Ho- bove the rizon of SeainFeet. the Sea.	' "	Height of Depression the Eye a- of the Ho- bove the rizon of SeainFeet. the Sea.	' "	Height of Depression the Eye a- of the Ho- bove the rizon of SeainFeet. the Sea.	' "
1	0. 57	12	3. 18	35	5. 39
2	1. 21	14	3. 34	40	6. 2
3	1. 39	16	3. 49	45	6. 24
4	1. 55	18	4. 3	50	6. 44
5	2. 8	20	4. 16	60	7. 23
6	2. 20	22	4. 28	70	7. 59
7	2. 31	24	4. 40	80	8. 32
8	2. 42	26	4. 52	90	9. 3
9	2. 52	28	5. 3	100	9. 33
10	3. 1	30	5. 14		

P R O B. IV.

*Having the Distances of three Points, A, B and C, either in a straight Line, or forming a Triangle; at any Station S, without the Triangle, to find its Distance from these Points.*

P L A T E I. F I G. V.

B Y P R O T R A C T I O N.

At S, with a Theodolite, or Hadley's Sextant, take the Angles A S C and C S B: then, making A B (the Side which subtends the greatest Angle, and is opposite to the middlemost Point) the Base; at B, make the Angle A B D equal to the Angle A S C, and lying on that Side of A B which is farthest from S; and at A, make the Angle B A D equal to C S B, intersecting A D in D; draw a Circle passing through the three Points

Points A, D, B, and it will also pass through S: draw out the Line C D, and it will cut the Circle in S, the Point of Station required.

*Demonstration.* The Angles A S D and A B D are equal, for they stand in the same Segment A S B D (Eucl. B. III. Prop. 21.) A S D is therefore equal to the first observed Angle. In the same manner, B S D and B A D are equal, for they stand in the same Segment D A S B; and D S B is therefore equal to the other observed Angle; and S the Point of Station, and S A, S C, S B, the Distances required.

#### BY CALCULATION.

1. In the Triangle A B C, the Sides being known, the Angles are found by Trig. Sol. 4. Page 3.

2. In the Triangle A B D, the Side A B is given by Supposition; and the Angles D A B and D B A equal, respectively, to the corresponding observed Angles; therefore the Angle at D is also known: from thence (by Trig. Sol. 1.) find the Side A D. Then,

3. In the Triangle C A D, the Sides C A and A D are known, and the included Angle C A D (which is what B A D wants of B A C) from thence (by Trig. Sol. 3.) find A C D. Then,

4. In the Triangle A S C, two Angles A S C and A C S, and the Side A C opposite to one of them, are known; from thence (by Trig. Sol. 1.) A S may be found. In the same manner may S B or S C be found.

If the Point C is on the other Side of A B, towards S, or in the Line A B, it is obvious that S may be found in the same manner; the same Protraction, and Calculation, serving for each of these Cases.

#### FIG. VI.

If it shall happen that the Points C and D fall so near each other, that in protracting, C D is produced with Uncertainty, which will happen

when the Point C lies beyond A B ; and the observed Angle A S C is nearly equal to the Angle A B C , and the other observed Angle C S B is nearly equal to the Angle C A B . In which Case take the Point C on the Side of A B which is next to S , as at K : or, shift your Station nearer to, or farther from C , and find the Position of that Station, and then its Distance from S . Or, on C B , make the Angle C B D equal to the observed Angle A S C , and the Angle B C D equal to the Supplement to  $180^{\circ}$  of both the observed Angles A S B ; then the Intersection of these Sides will be the Point D : draw a Circle passing through the Points B , C , D , and A D produced, will cut that Circle in S , the Point of Station required. For,

B S D C being a quadrilateral Figure inscribed in a Circle, B S D is the Supplement of B C D to  $180^{\circ}$  (by Eucl. B. III. Prop. 22.) and by Construction equal to the two observed Angles : and A S C and D B C being equal, as standing on the same Arc D C ; A S C must be the first observed Angle, and C S B (the Remainder of A S B) the other ; and S therefore the Point of Station.

Here it may be observed, that the Distance of an Object found in this manner by Hadley's Sextant, is the real aerial Distance in a straight Line from the Eye to the Object ; but it is the level Distance that is necessary in surveying. If the Object therefore is on a Hill, or Eminence, that is near, and of considerable Height, it will be greater than is required in surveying, as the Hypotenuse of a right-angled Triangle is greater than the Base. To find the level Distance, or Distance of a Point within the Base of the Hill, right below the Object on the Top of it ; pitch on some remarkable thing that lies below, and in a Line with the Object, and take the Angle with that instead of the Object on the Hill.

## P R O B. V.

Having the Distances of three Points A, B and C, that form a Triangle, at any Station S, in the Direction of one of the Sides, to find its Distance from these Points.

## F I G. VII.

## B Y P R O T R A C T I O N.

Take the Angle A S B with an Instrument; subtract that Angle from C A B, and the Remainder will be the Angle A B S; for C A B is the external Angle of the Triangle A B S, and therefore (by Eucl. B. I. Prop. 32.) is equal to A S B and A B S taken together. Then, at B, on the Side B A, draw the Angle A B S; produce C A, and it will intersect B S in S, the Point of Station; and S A, S C, S B will be the Distances required.

## B Y C A L C U L A T I O N.

In the Triangle A B S, the Angle at S is found by Observation; the Angle B A S is the Supplement of C A B to  $180^{\circ}$ , and therefore the Angle A B S is also known: that is, in the Triangle A B S, all the Angles and a Side are known; from thence the other two Sides S A and S B are found by Trig. Sol. I.

## P R O B. VI.

Having the Distances of three Points A, B and C, forming a Triangle; at any Station S, within the Triangle, to find its Distance from these Points.

## F I G. VIII.

## B Y P R O T R A C T I O N.

Make any one of the Sides, suppose A B, the Base: take with an Instrument the Angles A S C and C S B, which the two other Sides subtend: at

B make the Angle A B D (lying without the Base of the Triangle) equal to the Supplement of A S C to  $180^\circ$ ; and the Angle B A D equal to the Supplement of C S B: these Sides produced, will intersect in D. Draw a Circle passing through the three Points A, D and B; join C D, and it will intersect the Circle in S, the Point of Station; and S A, S C, S B will be the Distances required.

*Demonstration.* The Angles A S D and A B D are equal, for they stand in the same Segment A S B D (Eucl. B. III. Prop. 21.) therefore A S C, the Supplement of A S D, is also the Supplement of A B D ; and, by Construction equal to the first observed Angle. Again, B S D and B A D are equal, for they stand in the same Segment B S A D ; and B S C, the Supplement of B S D, is also the Supplement of B A D ; and therefore, by Construction, equal to the other observed Angle ; and S the Point of Station ; and S A, S B, S C, the Distances required.

## BY CALCULATION.

1. In the Triangle A B D, the Angles B A D and A B D are known, being respectively Supplements of the observed Angles A S C and B S C ; and therefore their Supplement A D B is known ; and the Side A B opposite to one of them : from thence (by Trig. Sol. 1.) find the Side A D. Then,

2. In the Triangle  $C A D$ , the Sides  $A C$  and  $A D$  are known, and the included Angle  $C A D$  (equal to  $C A B$  and  $B A D$  together) from thence find the Angle  $A C D$ , by Trig. Sol. 3.

3. In the Triangle A C S, the Angles at C and S are known, and the Side A C, opposite to one of them; from thence (by Trig. Sol. 1.) find A S and C S.

PLATE I. FIG. IX.

If the Point of Station is in any of the Sides, as at  $S$  in the Side  $A B$ ; take the Angle which one of the other Sides, suppose  $A C$ , subtends there; and on any Part of  $A B$ , as  $D$ , make the Angle  $A D E$  equal to the observed Angle  $A S C$ : through  $C$ , draw the Line  $C S$  parallel to  $D E$  cutting

cutting A B in S, and S will be the Point of Station. For C S and D E being parallel, the Angles A S C and A D E are equal (Eucl. B. I. Prop. 29.) A S C is therefore the observed Angle, and S the Point of Station.

On account of the great Utility of these last Problems in Surveying, I shall shew a general, and more ready Method, of finding, by Protraction, the Point of Station S, in the three principal Cases foregoing; communicated by Capt. John Campbel of his Majesty's Navy.

1. *When the observed Angle A S B is less than 90 Degrees.* From the Points A and B draw the Angles B A m and A B M, lying towards S, each equal to the Complement of the observed Angle A S B, and the Intersection of the Lines A m, B M, will be the Center m. of the Circle A S B. From the Extremities of either of the other Sides of the Triangle A B C, as B C, draw the Angles B C n and C B N, each equal to the Complement of the observed Angle C S B, and the Lines C n, B N, will intersect each other in n, the Center of the Circle C S B; and the Point of Intersection of these two Circles shall be the required Station S. For,

The Angle A S B being less than a Right-Angle, the Line A B will divide the Circle A M S B into two unequal Segments, in the greatest of which (A M S B) S must stand (Eucl. B. III. Prop. 31.) and therefore the Center, m, which is always in the greatest Segment, will be on the same Side of A B that S is. Also, in the Triangle m A B, the Angles A B m and B A M together, are the Supplement of the Angle A m B to two Right-Angles; but A B m and B A m are (by Construction) the Supplement of twice A S B to two Right-Angles; A m B is therefore double the Angle A S B; and standing on the same Arc (by Eucl. B. III. Prop. 30.) S must be in the Circumference of the Circle A S B.

In the same Manner it may be shewn that S is in the Circumference of the Circle C S B, and therefore it must be in their Intersection.

2. *When the observed Angle A S B (Fig. 19.) is more than 90 Degrees.* Then subtract 90 from it, and make the Angles at A and B each equal to the Remainder; and to lie on the Side of A B, which is farthest from S;

and  $m$ , the Point of Intersection of these two Lines will be the Center of a Circle, which will pass through  $A$ ,  $B$  and  $S$ . For, the Angle  $ASB$  being more than  $90^{\circ}$  Degrees, the Line  $AB$  will divide the Circle into two unequal Segments, in the least of which ( $ASB$ ) the Point of Station  $S$  must be; and the Center  $m$ , therefore, on the Side of  $AB$  opposite to  $S$ . Also, the Angle  $mAB = mBA =$  (by Construction) to the difference between  $ASB$  and a Right-Angle: also the Angle  $nCB = nBC = \overline{CSB} - 90^{\circ}$ .

*The Point of Station  $S$  found instrumentally.*

The Point  $S$  may be readily laid down on a Draught, by drawing on a loose transparent Paper indefinite Right-lines  $SA$ ,  $SB$ ,  $SC$ , at Angles equal to those observed; which being placed on the Draught so as each Line may pass over, or coincide with, its respective Object, the angular Point  $S$  will then coincide with the Place of Observation. Or,

Provide a graduated Semicircle of Brass, about 6 Inches in Diameter, having three Radij with chamfered Edges, each about 20 Inches long, (or as long as it may be judged the Distance of the Stations from the three given Objects may require) one of which Radij to be a Continuation of the Diameter that passes through the Beginning of the Degrees on the Semicircle, but immovably fixed to it, the other two moveable round the Center, so as to be set and screwed fast to the Semicircle at any Angle. In the Center let there be a small Socket, or Hole, to admit a Pin for marking the central Point on the Draught. When the sloped Edges of the two moveable Radij are set and screwed fast to the Semicircle, at the respective Degrees and Minutes of the two observed Angles, and the whole Instrument moved on the Draught until the Edges of the three Radij are made to lie along the three stasimetric Points, each touching its respective Point, the Center of the Semicircle will then be in the Point of Station  $S$ ; which may be marked on the Draught, through the Socket, with a Pin. Such an Instrument as this may be called a *Station-pointer*; and would be found convenient for finding the Point of Station readily and accurately, except when the given Objects were near; when the Breadth of the Arch, of the Radij, and of the Brass about the Center of the Semicircle might hinder the Points to be seen, or the Radij to be placed so as to comprehend a very small Angle between them.

PROB.

PROB. VII.

Having the Distance and Position of two Points A and B; at two Stations, S and s, to find the Distance of these Stations from each other, and from the given Points.

PLATE I. FIG. X.

1. At  $S$ , take the Angles  $ASs$  and  $BSs$ ; and at  $s$ , take the Angles  $A_sS$  and  $B_sS$ . Then,
2. Making  $Ss$  any length, suppose 10, take it from a Scale of equal parts, and draw it on Paper: and at  $S$  and  $s$ , protract the respective observed Angles, and the Points  $A$  and  $B$  will be determined in Proportion to the assumed Length of  $Ss$ . Then, taking the protracted Length of  $AB$  from the Scale, say,

As the protracted Length of A B,  
Is to the assumed Length of S s,  
So is the true Length of A B,  
To the true Distance of S and s.

3. The Distance  $Ss$  being thus found, and the Angles at  $S$  and  $s$  observed, the Distances  $SA$ ,  $SB$  and  $sA$ ,  $sB$ , may be taken from the proper Scale, after a new Protraction; or calculated, as in Prob. I.

PROB. VIII.

Having the Distance and magnetic Bearing of two Points, A and B protracted; at any Station S, (not very oblique to A B) to find its Distance from these Points, by the Needle.

PLATE I. FIG. XI.

At S, with a good magnetic Needle, take the Bearings of A and B, in Degrees and Parts of a Degree: then, from these Points draw out their E respective

respective Bearings in the opposite Direction, towards S: that is, if A bears exactly North; draw a Line from the Point A exactly South; if it bears East 10, or 20 Degrees southward, draw the Line West 10, or 20 Degrees northward, and so for any other Bearing. Draw the opposite Bearing of B in the same manner; and S, the Intersection of these two Lines, will be the Point of Station; and S A, S B, the Distances required.

This is an easy and ready way of finding the Distance of any Station, from two Places whose Distance has been accurately determined before; and will be found convenient very often in the Course of a Survey; and sufficiently exact on most Occasions; provided the Places, A and B, are not very remote from the Station, nor the Intersection of their Bearings too oblique. If the Needle is good, a Distance of 20 Miles is not to be reckoned too far, when the Angle subtended by the two Places is not less than 50 Degrees, nor more than 140.

#### C H A P. IV.

##### *Of the Examination of Surveying Instruments.*

THE Accuracy of geometrical Mensuration depends so much on the Accuracy of the Instruments made use of in it, that it is a very necessary Qualification in a Surveyor to be capable of examining and adjusting them. Some Instruments are made in such a Manner, without any Necessity, and as if it were on purpose, that it is very troublesome to find whether all their Parts are exact or not; and for that Reason Surveyors are too apt to take them on Trust, without Examination: but this often creates more Trouble in their Busines, and occasions Errors which they can neither account for, nor correct. If Surveyors were more cautious, Instrument-makers would become more careful.

##### *How to examine the Accuracy of a Theodolite.*

#### PLATE I. FIG. XII.

1. On a Level about a Mile or two long, such as the Margin of a Lake, or the High-water Mark along a well-sheltered Sand, when the Sea has been smooth

smooth a little before, set three upright Pales\* A, B, C, exactly in a Line; and about half a Mile, or a Mile from each other. If it can be done, let the Line run exactly on some remarkable Part of a distant Object.

2. Set the Theodolite horizontal, with its Center right over the Hole B, where the middle Pole stood: set the Index at 0 Degrees, and turn the Theodolite till, through the Telescope, you see the Pole A at the vertical Wire; then screw the Instrument fast in that Position. Next turn the Index to 90, and remark some distant Object exactly at the vertical Wire of the Telescope: if no such Object is to be seen, set up a Pole D in that Direction half a Mile off, or as far as it can be seen.

3. Then set the Index at 0 Deg. again, and turn the Instrument till you see the Pole C, through the Telescope exactly at the vertical Wire: screw the Instrument fast in that Position; turn the Index to 90; look through the Telescope; and if the Object, or Pole, D, is exactly at the vertical Wire, then is the Index right centered, and the graduated Circle exactly divided into four Quadrants; for the Line D B, standing on the right Line A C, makes the Angles on each Side, A B D and C B D, equal. But, if the Object is not seen exactly at the vertical Wire, observe how much it is over, or falls short of it, and half that Difference is the Error of the Division at 90 †.

4. To find if the Graduations of each Quadrant are right: Apply the Beginning of the Divisions of the vernier Scale, to the Beginning of each Degree of the Theodolite; and if the End of the Divisions of the Vernier reaches the respective Degree precisely in each Application, then the Graduations of the Theodolite are right; otherwise one, or more, are wrong. If each Degree is subdivided, these Subdivisions may be examined in the same manner. This is a very easy and nice way of examining the Graduations of any Instrument that has a vernier Scale on it.

\* To place a Pole upright; stand straight up with your Feet close together; put one End of the Pole between your Toes, and take the other up along your Breast and Chin, and in that Posture press it down with both Hands till it stands firm.

† By a proper Application of the above Principles, the Graduation of Instruments, large or small, may be rendered as accurate as the Materials made use of will admit, and the least Degree of Inaccuracy in them readily discoverable; as could be easily shewn, were this a Place for it.

To find if the Wire of the Telescope is exactly vertical: hang up a Plumb-line at some Distance; set the Theodolite level; look through the Telescope to the Plumb-line when at rest, and if the upright Wire moves exactly along the Plumb-line when the Telescope is elevated or depressed, then is the Wire exactly vertical. If the Wire crosses the Plumb-line in any Part, turn it by the proper Screws, or Pins, till you get it exact.

*Cautions in taking Angles with a Theodolite.*

1. Spread the Legs that support it pretty wide, and thrust them firm in the Ground, especially on Sand, or soft Ground; that they may not yield, or sink unequally during the Observations.

2. Set it level; otherwise the Angles will not be precise.

3. Screw the Ball firm in the Socket, that in turning the Index, the Diameter of the Theodolite may not vary from the Object it was directed to; which, without Care in the Observer, or in the Structure of the Instrument, will happen frequently.

4. The Angles should be taken twice over, and read off by different Persons.

5. It would often prevent Mistakes in Angles, and contribute to Dispatch, if a Theodolite was provided with two Telescopes, one above its Plane, and the other below it; the lower one to move round only when the Circumference of the Theodolite was turned; by which Means it would be readily seen through this Telescope when the Diameter shifted from the Direction of the Object it was fixed at. Also if these Telescopes shewed Objects erect, and in their natural Position, instead of being reversed and inverted; and if the Telescopes magnified more than they commonly do. It would likewise be an Advantage if the Vernier was made to give every Minute of a Degree, in place of four or five, as in most Theodolites.

*Of Hadley's Sextant.*

This Instrument may be used with great Advantage in Maritim Surveys, on most Occasions; being more portable, more readily applied to the taking of

of Angles, and generally more accurately and minutely divided than Theodosites are: an Observer is less liable to Mistakes with it; and, which is a very material Advantage, he can take Angles with it at Sea, as well as on Land. There are two Inconveniences attending the Use of it; one is, that small, or obscure Objects, especially if the Ground rises beyond them, cannot be easily perceived in the Glasses: the other is, that no Angle above 120 Degrees can be taken by it at once, and that sometimes there are not proper intervening Objects for dividing a larger Angle into two \*.

A Brass Sextant of nine Inches Radius, though it does not admit of Back-observations, will be found as convenient a Size as any for a Surveyor. It should be provided with a small Telescope, that can be taken off and put on at Pleasure, for observing celestial Bodies; and with a Tube for terrestrial Objects, which is better than a thin Plate of Brass with Holes in it. There should likewise be a Screw at the End of the Index, to bring the two Objects exactly to one another, after they have been first brought near by the Hand; and if a Staff with a Ball and Socket is fitted to it, and a Belt to support the whole, it will be found a great Ease to the Observer. The Telescope and Tube should be so contrived, as to be raised occasionally as high as the Middle of the transparent Part of the Horizon-glass, for observing Objects that are bright enough to be seen by Reflection from it.

Each of the Glasses belonging to the Sextant should have its two Surfaces perfect Planes, and parallel to one another; and the Index and Horizon-glasses should stand perpendicular to the Plane of the Instrument. These Particulars of the Glasses ought to be carefully attended to, and adjusted by the Instrument-maker, before he gives the Sextant out of his Hand: for though Defects in the Glasses may be detected by proper Examination, yet the want of Parallelism and Plainness cannot be corrected without new Glasses; and, to adjust the Inclination, is troublesome to young Practitioners. The first thing however which a Beginner should learn is, the easiest way of holding the Sextant for making Observations, and of bringing the reflected

\* Any Angle may be taken by the Back-observation, by one experienced in this manner of observing; but the Back-observation is too difficult and troublesome to be recommended in general to all Practitioners; and requires a Sextant or Octant of at least 16 Inches Radius, to have sufficient Room for the Back-vanes, which is an inconvenient Size for Land Surveys.

Image and Object near each other by moving the Index. Tho' this requires Practice more than Precept, the following Directions may be found serviceable at first.

*How to hold the Sextant for taking Angles.*

Whether the Sextant is to be held vertically, or horizontally with its Face uppermost, it should be supported chiefly with the Right-hand, and the Index governed by the Left, in the following Manner. For horizontal Angles; hold the Instrument horizontally, with the Face uppermost, and place the Thumb of the Right-hand against the Edge of the curved Part on which the Sight-vane stands, but so as not to touch the Index when it is set at 0, or the Beginning of the Degrees; and extend your Fingers across the Back of the Instrument, so as to lay hold of the opposite Edge of the additional Part on which the Horizon-glass stands, placing the Fore-finger on one Side, the two next Fingers on the other Side of the greatest Swelling, and the Point of the little Finger against the Back of the Instrument. When there is no Staff and Belt, horizontal Angles are easiest taken sitting with the right Ankle on the left Knee, and the right Elbow supported by the right Knee. The Index may be managed by placing the Thumb of the left Hand against the outer Edge of the graduated Arch, and keeping the End of the Index betwixt the two first Fingers: or by taking the End of the Index between the Fore-finger and Thumb, and resting the other Fingers against the middle Bar, or against the Far-side of the Instrument, as a very little Practice will direct.

For taking vertical Angles: hold the Sextant in a vertical Position, with the Thumb and Fingers of the Right-hand, as before, for horizontal Angles; or as the Form of the Swelling at the Horizon-vane enables you to take the surest Hold; rest the Extremity of the Arch against your Breast, and that will enable you to manage the Index more easily with your Left-hand.

When Angles are to be taken with the Instrument inverted, or with the Face below, then it must be held and supported chiefly by the Left-hand, and the Right reserved for governing the Index, in the following Manner. Hold the Sextant horizontal, with its Face downward: place the Thumb of your Left-hand on the Center-plate at the Back of the Index-glass; the

three first Fingers above the Center-part, and the Inside of the little Finger against the Edge of it; and when the Eye is applied to the Sight-vane, press the Edge of that Side against your Forehead: take care that the little Finger, by being too low, does not intercept the Sight of any of the Objects.

In all Observations, the Sextant should be held in a Plane passing through the Eye of the Observer and the two Objects whose angular Distance is to be found.

*How to examine the Glasses of a Sextant.*

To find whether the two Surfaces of any one of the reflecting Glasses are parallel: apply your Eye at one End of it, and observe the Image of some distant Object reflected very obliquely from it; if that Image appears single and well defined about the Edges, it is a Proof that the Surfaces are parallel: on the contrary, if the Edge of the reflected Image appears misted, as if it threw a Shadow from it, or separated like two Edges, it is a Proof that the two Surfaces of the Glass are inclined to each other. If the Image in the Speculum, particularly the Sun, is viewed through a small Telescope, the Examination will be more perfect.

To find whether the Surface of a reflecting Glass is a perfect Plane. Chuse two distant Objects as nearly on a Level as can be had: hold the Sextant horizontal, with your Eye at the Sight-vane, view the Left-hand Object directly through the transparent Part of the Horizon-glass, and move the Index till the reflected Image of the other is seen below it in the silvered Part; make the two Images unite just at the Line of Separation: then turn the Sextant round slowly in its own Plane, so as to make the united Images move along the Line of Separation of the Horizon-glass: if the Images continue united, without receding from each other, or varying their respective Position, the reflecting Surface is then sufficiently plane.

To find if the reflecting Glasses are perpendicular to the Plane of the Sextant. Move the Index near the Middle of the Arch of the Sextant; apply your Eye obliquely near one End of the reflecting Glass, so as to see Part of the Arch of the Sextant by Reflection, and Part of it by direct Vision: if the two Parts of the Arch form one uniform Curve, without

appearing

appearing in the least above or below each other, then is the Glass perpendicular to the Plane of the Sextant. If the Image of the Arch seen by Reflection is higher than the Arch seen directly, then the Glass inclines forward; in which Case, slacken the fore-adjusting Screw of the Glass, and screw in the back one. If that Part of the Arch seen directly is higher than its Image seen by Reflection, then the Glass leans backward, and the back-adjusting Screw must be slackened, and the fore one screwed in. Be careful, in all Cases where one Screw draws, or pushes against another, to slacken one of them before you screw up the other.

To find if the two Surfaces of a red or darkening Glass, are parallel and perfectly plane. This must be done by means of the Sun when it is near the Meridian, in the following Manner. Hold the Sextant vertically, and direct the Sight to some Object in the Horizon, or between you and the Sky, under the Sun; turn down the red Glass, and move the Index till the reflected Image of the Sun is in Contact with the Object seen directly: fix then the Index, and turn the red Glass round in its square Frame; view the Sun's Image and Object immediately, and if the Sun's Image is neither raised nor depressed, but continues in Contact with the Object below, as before, then the Surfaces of the darkening Glass are true.

*How to observe the horizontal Angle, or angular Distance, between two Objects.*

First adjust the Sextant, and if the Objects are not small, pitch on a sharp Top, or Corner, or some small distinct Part in each to observe; then having set the Index to 0 Deg. hold the Sextant horizontally, as above directed, and as nearly in a Plane passing through the two Objects as you can; direct the Sight through the Tube to the Left-hand Object, till it is seen directly through the transparent Part of the Horizon-glass: keeping that Object still in Sight there, move the Index till the other Object is seen by Reflection in the silvered Part of the Horizon-glass; then bring both Objects together by the Index, and by the Inclination of the Plane of the Sextant when necessary, till they unite as one, or appear to join in one vertical Line, in the Middle of the Line which divides the transparent and reflecting Parts of the Horizon-glass: the two Objects thus coinciding, or one appearing directly below the other, the Index then shews on the Limb the Angle which the two

Objects subtend at the naked Eye. This Angle is always double the Inclination of the Planes of the two reflecting Glasses to one another; and therefore every Degree and Minute the Index is *actually* moved from o, to bring the two Objects together, the Angle subtended by them at the Eye will be twice that Number of Degrees and Minutes; and is accordingly numbered so on the Arch of the Sextant; which is really an Arc of 60 Degrees only, but graduated into  $120^{\circ}$ . As this has been demonstrated in several Books and Pamphlets, it is needless to insist on it here.

The Angle found in this Manner between two Objects that are *near* the Observer, is not precise; and may be reckoned exact only when the Objects are above half a Mile off. For, to get the Angle truly exact, the Objects should be viewed from the Center of the Index-glass, and not where the Sight-vane is placed; therefore, except the Objects are so remote that the Distance between the Index-glass and Sight-vane vanishes, or is as nothing compared to it, the Angle will not be quite exact. This Inaccuracy in the Angle between near Objects is called the *Parallax* of the Instrument; and is the Angle which the Distance between the Index-glass and Sight-vane subtends at any near Object. It is so small, that a Surveyor will seldom have Occasion to regard it: but if it shall happen that great Accuracy is required, let him chuse a distant Object exactly in a Line with each of the near ones, and take the Angles between *them* and that will be the true Angle between the near Objects. Or, observe the Angle between near Objects, when the Sextant has been first properly adjusted by a distant Object; then adjust it by the Left-hand Object, which will bring the Index on the Arch of Excess beyond o Degrees: add that Excess to the Angle found between the Objects, and the Sum will be the true Angle between them. If one of the Objects is near, and the other distant, and no remote Object to be found in a Line with the near one; adjust the Sextant to the near Object, and then take the Angle between them, and the Error of Parallax will be removed.

*How to observe the Altitude of the Sun on the Sea, or on Shore at the Edge of the Sea.*

Adjust the Sextant by the Horizon (as is afterwards directed:) raise the Telescope, or Tube, as high as the Middle of the transparent Part of the

Horizon-glass; set the Index to 0 Degrees; turn in the darkening-glasses, or only one of them if the Sun is not very bright; turn your Face directly to the Sun; hold the Sextant in a vertical Position, with its Plane passing as near through the Sun's Center and the nearest Part of the Horizon as you can, and direct the Sight to that Part of the Horizon; then move the Index from 0 on the Limb, till you see the red Image of the Sun come down toward the Horizon. If there is any Difficulty in finding that red Image, it is either because the Sun is obscured by a Cloud, or because the Sextant is not in a vertical Plane that passes through the Sun, which must be discovered by glancing at the Sun with your naked Eye, and in the latter Case, by turning your Body more directly to the Sun. Having brought the Sun's Image almost down to the Horizon, then, with your Left-hand, swing the Sextant backward and forward, from one Hand to the other, and the Image of the Sun will seem to describe the Arch of a Circle convex to the Horizon: move the Index till the lower Edge of the Image just grazes on the Horizon; when it is at the lowest Point of the Arch apparently described, then stop the Index, and mark down the Degrees and Minutes shewn on the Limb; which call *The observed Altitude*. When this is corrected by the Sun's Semidiameter, the Dip of the Horizon, the Refraction, and the Index-error, if there is any, it will give the true Altitude of the Sun's Center.

*How to take an horizontal Angle between two Objects with the Sextant inverted.*

To observe in this Manner: set the Index at 0, and hold the Instrument as above directed; direct the Sight to the Left-hand Object till it is seen through the transparent Part of the Horizon-glass, and you will likewise perceive its reflected Image in the silvered Part: then advance the Index, and at the same time keep the reflected Image in Sight, by turning your Body and the Sextant toward the Right-hand, till you see the other Object through the transparent Part; make them join in the Middle of the Line of Separation of the Horizon-glass, and the Index will shew the Angle on the Limb.

This Method of observing will be often convenient, because, after a little Practice, an Angle may be taken more readily this way than with the Face of

of the Sextant uppermost ; and also because the faintest Object may often be found by direct vision through the transparent Part of the Glass, when it could not be so easily found by Reflection.

*How to adjust the Horizon-glass for Observation, by a horizontal Line.*

To adjust the Horizon-glass is, to make it parallel to the Index-glass when the Pointer of the Vernier is placed at 0, or the Beginning of the Degrees. This is the most important Adjustment of all for an Observer ; and should be done every Time one begins to make any Observations. The horizontal Lines fitteſt for this Operation are, the Horizon of the Sea, which is the leaſt ; and the Roof, or upper Part of the Wall of a large well-built House above half a Mile diſtant.

To adjust by the Horizon of the Sea, or any other horizontal Line : turn the Index, till the Pointer of the Vernier stands exactly at 0 on the Limb, and fix the Index there : hold the Sextant in a vertical Position, and obſerve whether the Edge of the Sea-horizon (or other horizontal Line) ſeen in the ſilvered Part of the Glass, makes one continued Line with the Edge of the Sea ſeen in the transparent Part : if it does not, and the Line appears broken, or one Part of it above the other, bring them into one ſtraight Line by turning the adjusting Lever at the Back of the Horizon-glass ; then fix the Lever by the Button-screw in the Middle of it, and the Horizon-glass is adjuſted, or parallel to the Index-glaſs.

To find the Error of this Adjustment : move the Index a little from its Place ; hold the Sextant in a vertical Position as before, and turn the Index till the Line of the Sea-horizon (or other horizontal Line) ſeen in the ſilvered Part of the Horizon-glaſs joins exactly in one ſtraight Line with the Edge of the Sea ſeen through the transparent Part : then the Number of Minutes by which 0 on the Vernier differs from 0 on the Limb is the Error of the Adjustment ; and muſt be allowed for in all the Angles taken with this Adjustment, in the following Manner. If the Pointer, or 0 on the Vernier, stands on the quadrantal Arch, or to the Left-hand of 0 on the Limb, then this Error is to be ſubtracted from the Angle ſhewn on the Limb by the

Index; but if  $\circ$  on the Vernier stands on the Arch of Excess, or to the Right of  $\circ$  on the Limb, then the Error of the Adjustment is to be added to the Angle shewn on the Limb.

To verify the Error of Adjustment found by this, or any other Method: set the Index so as to shew the exact Error found, and to shew it on the quadrant Arch, if that Error is to be subtracted, but on the Arch of Excess, if the Error is to be added. Then fix the Index, and directing the Sight to some distant Object, take notice whether the direct and reflected Images perfectly coincide or not: if they do, the Index Error is truly determined; if not, it must be corrected by the Method proper for it.

When the Index is moved from you, to bring an Image and Object together, it will give a different Angle from that same Angle taken by moving the Index towards you. This, probably, is occasioned by the Index bending or starting a little by the Pressure of the Hand. To correct this, when great Accuracy is required, take the Angles two or three times over each Way, and the Medium of all will be the true Angle.

*How to adjust the Horizon-glass by a straight vertical Line.*

The Lines fittest for this are, the upright Corner of a House, or a Pole set upright in the Ground, at least half a Mile off.

Bring the Pointer of the Vernier to  $\circ$  on the Limb; hold the Plane of the Sextant horizontal; view the vertical Line in the Middle of the transparent Part of the Horizon-glass, and its Image in the reflecting Part, and turn the Lever till both make one continued Line, neither appearing to the Left, or Right of the other; then screw the Horizon-lever fast, and the Horizon-glass is adjusted.

To find the Error of this Adjustment: move the Index out of its Place; then set the Vernier again to  $\circ$  on the Limb; hold the Plane of the Sextant horizontal; view the vertical Line directly through the transparent Part of the Horizon-glass, and its Image in the reflecting Part: if they make one continued Line, there is no Error; if they do not make one continued Line, move

move the Index a little to one Side or the other till they do so, and the Number of Minutes by which the Pointer of the Vernier differs from 0 on the Limb, is the Error of this Adjustment; which is to be subtracted from each Angle taken, if the Pointer stands on the quadrantal Arch, and added when it is on the Arch of Excess. In reading off the Minutes on the Arch of Excess, remember to take their Compliment to 30, in Place of what the Vernier shews.

*To adjust the Horizon-glass by the Sun.*

Set the Pointer of the Vernier to 0 on the Limb; hold the Sextant horizontal; raise the Telescope, or Tube, as high as the Middle of the transparent Part of the Horizon-glass; use a smoked, or red Glass to defend the Eye, and slip a Screen of Pastboard nine Inches square, with a Hole in the Middle, over the End of the Telescope, to keep the Face from being scorched; then view the Sun and its Image both through the transparent Part of the Horizon-glass, and by the Lever make them cover each other exactly, so as to appear as one Object; screw the Lever fast, and the Horizon-glass is adjusted.

The Adjustment may be made in the same Manner by the Moon, or a Star, without a Darkening-glass or Screen.

To find the Error of this Adjustment; move the Index a little from its former Place at 0; then, by the Index, make the two Images cover one another exactly as before; and the Number of Minutes by which the Pointer of the Vernier differs from 0 on the Limb, is the Error of Adjustment, and must be allowed for as above.

*Another, and rather a nicer, Manner of adjusting by the Sun.*

Set the Pointer of the Vernier to 0 on the Limb; hold the Quadrant horizontally, with the darkening Glass and Screen to defend your Eye and Face; measure the Diameter of the Sun, first on the quadrantal Arch, and next on the Arch of Excess; or with the Index, first on the Left-hand of 0

on the Limb, and then on the Right of it. If these two Diameters are equal, then is the Horizon-glass adjusted: if the Diameters are unequal, turn the Horizon-lever a little one way or the other, till upon measuring them again they are found equal: or, which is equivalent, and often more convenient, when the Error is small; mark the Difference between the two Diameters, and make Allowance of half that Difference in each Angle observed on the Limb, adding or subtracting it as before directed. In reading off the odd Minutes on the Arch of Excess, remember to take the Compliment to 30 of what the Vernier shews.

The Diameter of the Sun is measured thus. Set the Pointer of the Vernier at 0; hold the Sextant horizontally; by the Index bring the reflected Image in ~~external~~ Contact with the direct Image; note the Degrees and Minutes of the Pointer from 0; then bring them in Contact on the opposite Side of the direct Image, and note the Degree and Minutes; see whether the Diameter on the quadrant Arch, or the Diameter on the Arch of Excess, is greatest; mark the Difference, and reserve the half of it for Use, if you do not correct the Error by the Lever.

The Adjustment may be made pretty exact by taking two Diameters of a straight-sided Land-object in the same Manner.

*How to adjust the Sextant by any remarkable Land-object at a Distance.*

It will often happen on a Survey, that no straight Line, either vertical or horizontal, nor no celestial Body, can be seen to adjust by. In this Case, pitch on some small remarkable Object at a Distance, as the tapering Top of a Hill, or Rock, or a small Hummock, or Chimney-head, or detached Branch of a Tree, for this Purpose. To adjust by such an Object, set the Index to 0; hold the Sextant horizontal; view the Object through the transparent Part, and its Image in the reflecting Part, of the Horizon-glass; and by the Lever, bring the Image directly below the Object in the Middle of the Line of Separation, so as neither appears to one Hand of the other, but that if a vertical Line is imagined to be drawn through any Point of the Image, it would pass through that same Point of the Object; then screw the

the Lever fast, and the Horizon-glass is adjusted. This Method of adjusting may be sufficient for common Purposes, but is not so very exact as the other Methods.

*How to examine the Angle of the Sextant..*

The Angle of the Sextant, measured by the Arc of the Limb comprehended between 0 and 120 may be verified, first, by adjusting the Glasses as before directed, and then taking, at one Station, five or six horizontal Angles, comprehending a whole Circle round ; add these Angles together, and if their Sum amounts to 360 Degrees exactly, the Angle of the Sextant and its Graduations may be concluded to be exact. The Objects that form these Angles must be very remarkable and sharp, and all of the same apparent Height, to prevent Mistakes in taking them ; otherwise Inaccuracy in the Observer may be imputed to the Instrument.

Another way of examining the Angle of the Sextant is, by setting perpendicular in a level Ground three Poles, or Staves, between four and five Feet high, with Flags flying at each, so as to form a Triangle, whose greatest Angle is less than 120 Degrees ; let the Distance of the Poles from one another be half a Mile, or as much farther as they can be seen distinctly. After adjusting the Sextant, take the three Angles of the Triangle very exactly, keeping the Center of the Sextant exactly on the Center of the Staff you observe at ; and if the Sum of the Angles amounts precisely to 180 Degrees, the Sextant may be concluded to be good. The several Graduations, and their Subdivisions, may be examined by the Vernier, as directed for the Theodolite.

*How to examine the Accuracy of the two quadrantal Radij of an astronomical Quadrant.*

PLATE I. FIG. XIII.

1. On a level Line set two Poles, A and B, perpendicular in the Ground, about a Mile, or more, from one another ; let a Flag fly at the Top of each, to render them more conspicuous ; and let the Poles appear on some remarkable sharp Part (C) of a remote Object ; the farther the Object is, the better.

3.

2. Put

2. Put the Quadrant on its Stand at B, with its Plane set horizontal, either by a spirit Level laid on it, or by a Carpenter's Level, and its Center right above the Hole where the Pole stood, and in the Position as at b. If the Stand cannot support it horizontally, lay the Quadrant steady on a firm Table. Then, with the Index at 0 Deg. look through the Telescope, and turn the Quadrant till you see the Object C exactly at the Intersection of the Wires. Let the Quadrant be kept firm in that Position, and turn the Index to 90, and observe through the Telescope what Object the Intersection of the Wires bears on : if it bears on nothing remarkable, set up a Pole and Flag D, in that Direction, as far off as can be seen.

3. With the Index at 90, turn the Quadrant on its Center a Quarter round, till it is in the Position a, so that you see the Pole A through the Telescope at the Intersection of the Wires. Then keeping the Quadrant firm in that Position, turn the Index to 0 Deg. and if through the Telescope you see the Object D, exactly at the Intersection of the Wires, then the two Radij of the Quadrant, drawn through 0 and 90, are precisely at right Angles, and the Axis of the Telescope parallel to the Radius it is fastened to. If the Object D, does not appear exactly at the Intersection of the Wires, but on one Side of it ; then half its Distance from the Intersection is the Error of the Quadrant ; and is either in the Angle of the two Radij, or in the Axis of the Telescope, which is a Line supposed drawn through the Center of the Eye-glass and the Intersection of the Cross-wires. It is much more likely to be in the Telescope than in the quadrantal Radij ; and therefore to correct it, the Wires in the Telescope must be shifted by their proper Screws, half the Distance between the Object and Intersection of the Wires ; and the Quadrant examined over again as before. If it is not right on this second Examination, correct the Wire over again by the Screws, and examin the Quadrant a third Time ; and if the Object D, and the Intersection of the Wires still disagree, conclude the Error to be in the Angle of the Radij between 0 and 90. Mark how much that Error is, and by it correct all future Observations. This Method is liable to a little Inaccuracy, if the Plane of the Quadrant is not made exactly horizontal.

*How to examine and adjust an astronomical Quadrant.*

1. Set up the Quadrant on the Margin of a Lake, close to the Edge of the Water, so that the Center of the Quadrant is just over the Water: (or, in a well-sheltered Bay when the Sea is smooth, right over the High-water Mark) set one Side of the Quadrant perpendicular by the Plumb-line; and measure the Height of the Center of the Telescope above the Surface of the Water (or above the High-water Mark) in Inches and Parts of an Inch.
2. About half a Mile off (but not above a Mile) set a Pole straight up in the Edge of the Lake (or on the High-water Mark) with a conspicuous sliding Cross-piece on it: let one Edge of the Cross-piece be set as high above the Surface of the Water (or High-water Mark) as the Center of the Telescope was, and about 2 Inches higher, if the Distance of the Stations is half a Mile;  $4\frac{1}{2}$  Inches higher if  $\frac{3}{4}$  of a Mile, and 8 Inches higher if their Distance is one Mile; for the Height of the Tangent above the Surface of the Earth in these Arcs \*.
3. Set the Index of the Quadrant to 0 Degrees, or exactly to the horizontal Radius, and direct the Telescope to the Pole: if, while the vertical Wire runs along the Pole, the horizontal Wire runs exactly along that Edge of the Cross-piece which was set as high as the Center of the Telescope, the Quadrant and Telescope are right. If the horizontal Wire does not coincide

\* The Height of a Tangent to the Extremity of any small Arc above the other Extremity of that Arc, is a third Proportional to the Diameter of the Earth and the Length of the Arc. For, imagine a Diameter drawn to one End of the Arc, and produced till it meets a Tangent drawn to the other End; then, by Eucl. B. III. Prop. 36. the Rectangle of the Diameter (in this Case the same with the Secant) and produced Part is equal to the Square of the Tangent. If the Diameter is called  $d$ , the produced Part  $p$ , the Tangent  $t$ , and the Arc  $a$ ; the two equal Rectangles will be expressed thus  $d \times p = t \times t$ : and because a small Arc and its Tangent are so nearly equal, that one may be assumed for the other; in place of  $t \times t$ , substitute  $a \times a$ , and it will be  $d \times p = a \times a$ . When two Rectangles are equal, their Sides are reciprocally proportional (Eucl. B. VI. Prop. 14.) therefore  $d : a :: a : p$ ; that is, as the Diameter is to any small Arc, suppose one Minute, so is that Arc to the Height of the Tangent above it. By this Proportion the Numbers in the Text were found, supposing the Diameter of the Earth to be 41798147 Feet.

in that Manner, shift it by the proper Screws till it does so, and the quadrantal Angle and Axis of the Telescope will then be right.

The Screws for shifting the Wires are on the Telescope, at the common Focus of the Object and Eye-glass. If you see the Edge of the Cross-piece below the horizontal Wire, ease the upper Screw, and screw in the lower one \* ; if the Edge of the Cross-piece appears above the horizontal Wire, ease the lower Screw, and screw in the upper one, till the horizontal Wire coincides with the Edge of the Cross-piece ; and the Axis of the Telescope is then right.

This Adjustment may be verified, by placing the Quadrant where the Pole stood, and the Pole where the Center of the Quadrant was, and the Axis of the Telescope, and the Edge of the Cross-piece, at the same Height above the Surface of the Water as at first ; then make the Radius of the Quadrant perpendicular by the Plumb-line, set the Index to 0 Degrees, look through the Telescope, and if the horizontal Wire and Edge of the Cross-piece coincide again, the Angle of the Quadrant and Axis of the Telescope are certainly right. If any Difference is perceived, move the horizontal Wire, by its Screws, half that Difference, and it will be rectified.

The Graduations of a Quadrant may be examined in the same Manner as those of the Theodolite above mentioned, by applying the Vernier to the several Degrees, and Parts of a Degree.

*How to adjust Bird's twelve-inch Quadrant, as made before the Year 1751.*

For performing this, there are three principal Parts to be adjusted, 1. The Pillar is to be set perpendicular to the Horizon. 2. One Side of the Quadrant is to be set perpendicular. 3. The Axis of the Telescope, or Line of Sight, is to be made parallel to the horizontal Radius of the Quadrant.

\* As the Telescope inverts Objects, what appears below the horizontal Wire is really above it, and therefore, screwing in the lower Screw, which pushes up the Wire, will bring it nearer the Object. The contrary is true of the upper Screw.

*1. To set the Pillar perpendicular.*

1. Set the Brass Pedestal on a firm Support. An iron-bound Cask, with a Free-stone, or thick Slate, kept on purpose to cover the Top of it, is as good and convenient for a Surveyor as any thing. Then hang the Quadrant on the Pillar, and by the Spirit-level and Screws in the Feet, the Pillar may be set perpendicular in the following Manner. Turn the Quadrant till the upper Side is parallel to two opposite Feet, and with one Hand unscrew that Screw of the Foot toward which the Buble inclines, and, at the same time, with the other Hand screw in the opposite Screw, till the Buble rests equally between the two Marks, or Scratches, on the glass Tube: then turn the Quadrant, horizontally, half round; and if the Buble does not rest exactly between the Scratches as before, by unscrewing that Screw towards which the Buble inclines, and at the same screwing in the other, make the Buble move over half the Space in the Tube by which they differ, and remember toward which End it inclines. Then turn the Quadrant one Quarter round, till its upper Side is parallel to the other two Feet; and if the Buble does not rest in the same Place, with respect to the Scratches, as it did before, make it do so by the Screws in the Feet; and then bring it equally between the Scratches, by lowering that End of the Tube toward which the Buble stood, or raising the other End. For doing this, there are three Screws at each End of the Tube. To lower one End, slacken the right and left, or pushing, Screws, and screw in the Middle Screw, till the Buble rests equally between the two Scratches: then the Pillar and Level are right, or nearly so. To prove if they are exact, or to make them so, if they are not; begin the same Process over again, by first turning the Quadrant half round, and if the Buble rests not equally between the Scratches, move it half the Difference by the Screws in the Feet; and then quarter round, and adjusting the Level first so far by the Screws in the Feet, and next by raising or lowering one End of the Tube as before directed; till you get the Buble to stand in the same Part of the Glass when the Quadrant is turned half round and quarter round; and the Pillar and Level will then be right.

Let it be adverted here, that in turning the Screws of the Feet, the Feet themselves will be apt to shift insensibly on the Support; which, if not pre-

vented, will render this Part of the Adjustment troublesome and tedious. To guard against this, small Grooves, or Furrows, the Width of the Screw-ends should be sunk a little in the Stone, or Support, and the Screws placed in these Grooves; or some other equivalent Contrivance to keep the Feet from shifting.

2. *To make one Side of the Quadrant perpendicular.*

This is done by making the Plumb-line hang first over the fine Point at the Arrow-head, on the center Plate; and next to cover the other fine Point on the Arch below.

It is made to hang over the upper Point, by screwing the elastic Bit of Brass (over which the Wire of the Plummet goes) out, or in, by its upper Screw-nail. It is made to hang over the lower Point, by the adjusting Screw within the Quadrant, which is turned by a round brotched Brass-button on its Head.

The Plummet should be made to hang in a small Cup, or Vessel, filled with Water, to make it more steady, and rest sooner: and if the Vessel is fastened by an Arm to the Pillar of the Pedestal, so as to move round along with the Quadrant, it will be the more convenient.

The Plumb-line should be made to hang very near the Plane of the Quadrant, at the lower Arrow, but not to touch it: which is done by pressing with the Hand the elastic Bit of Brass before mentioned, a little backward or forward.

3. *To adjust the Telescope, or to make its Axis horizontal, when the Index is at 90 Degrees.*

Set the Index to 90, keeping the Plummet and Level right: then direct the Telescope to some distant Object above a Mile off, (but the farther the better) till you see a remarkable Part of it exactly at the horizontal Wire: then take off the Quadrant from the Pillar, by unscrewing the two large Iron-screws; and invert the Quadrant, by screwing it on again with the Arch

Arch uppermost: then screw to the Arch of the Quadrant a crooked Bit of Brass, with a Notch in it, as near the Arrow as you can, and hang the Plummёт over that Bit of Brass, moving it to one Side or the other, till the Line covers the Point on the Arch; and if it does not then cover the Point on the Center-plate also, make it do so by the Screws in the Feet: then look through the Telescope; and if the horizontal Wire cuts the same Part of the Object it did before Inversion, the Telescope is right. If the Object appears higher or lower, the Wire in the Telescope must be moved half the Difference, by the two Steel-screws near the Eye-end of the Telescope: if the Object appears higher than the horizontal Wire, ease the lower Screw, and screw in the upper one; if the Object appears below, ease the upper Screw, and screw in the lower one; till the horizontal Wire cuts the distant Object at half the Difference it did before inverting it: then it is adjusted for Observation.

The Wire in the Telescope that is applied to the Limb of the Sun, in taking its Altitude, ought to be exactly horizontal when the Plane of the Quadrant is perpendicular. To find if it is so: adjust the Quadrant by the Level and Plummёт, as before directed; then point the Telescope to some small distinct Object at a Distance, as the Top of a Rock, a Chimney-head, &c. and remark a particular Part of it which the Wire crosses: turn the Quadrant a little about gently, towards one Side; and if all Parts of the Wire pass over the same Part of the Object, the Wire is truly horizontal. But if one End of the Wire crosses the Object higher, or lower, than the other, press the two small Screws that govern the Wire between your Finger and Thumnb, and at the same Time turn them side-wise, but in opposite Directions, till, upon looking again, and turning the Quadrant, you find all Parts of the Wire pass over the same Part of the Object; then it is truly horizontal.

These Adjustments may seem difficult and tedious at first; but they are necessary: and by practising two or three Times within Doors, till a right way of handling the several Parts of the Quadrant becomes familiar, they will be found much easier than may at first be imagined.

When

When a Quadrant of this Sort, made by a good Artist, has been first verified and adjusted, the simplest and easiest way of examining a Theodolite or *Hadley's Quadrant* is, first to take an Angle between any two smart Objects with the verified Quadrant, and then with the Instrument under Examination; and if they differ, to ascribe it to the latter.

*How to examine a magnetic Needle.*

First set the Box level; by raising, or lowering, one Side of it, till each End of the Needle, when it plays from Side to Side, runs along the Circumference of the graduated Circle in the Box; and both Ends of the Needle, when at rest, point equally high on it. Then turn the Box gently, till one End of the Needle rests at the Beginning of the Graduations exactly, and if the other End points exactly to  $180^{\circ}$ , the opposite Graduation, the Needle is so far right centered. Next turn the Box 90 Degrees round, and if the Ends of the Needle then point precisely to the opposite Degrees, then the Needle is truly centered. Each Degree may be tried in the same Manner.

Observe likewise if the Cap in the Center of the Needle is small and tapering in the Inside, or large and spherical: if it is a Portion of a large Sphere, the Point of the Steeple will be apt to rest on different Parts of it, and by that Means vary the Center of the Needle, and cause it to point wrong.

## C H A P. V.

Meridional Problems by the Stars; and Variation of the magnetic Needle by the Sun.

*How to fix a meridian Line by a Star, when it can be seen at its greatest Elongation on each Side of the Pole.*

PROVIDE two Plummets. Let one hang from a fixt Point: let the other hang over a Rod, supported horizontally about 6 or 8 Feet above the Ground, or Floor in a House, and so as to slide occasionally along the Rod: let the moveable Plummets hang four or five Feet northward of the fixt one. Some time before the Star is at its greatest Elongation, follow it, in its Motion, with the moveable Plummets, so as a Person a little behind the fixt Plummets may always see both in one, and just touching the Star. When the Star becomes stationary, or moves not beyond the moveable Plummets, set up a Light on a Staff, by Signals, about half a Mile off, precisely in the Direction of both the Plummets. Near twelve Hours afterwards, when the Star comes towards its greatest Elongation on the other Side of the Pole, with your Eye a Foot or two behind the fixt Plummets, follow it with the moveable Plummets, till you perceive it stationary as before: mark the Direction of the Plummets then by a Light put up precisely in that Line: take the Angle between the Places of the two Lights with Hadley's Quadrant, or a good Theodolite, the Center of the Instrument at the fixt Plummets; and a Pole set up in Day-light, bisecting that Angle, will be exactly in the Meridian seen from the fixt Plummets.

This Observation will be more accurate if the Eye is steadied by viewing the Plumb-lines and Star through a small Slit in a Plate of Brass stuck upright on a Stool, or on the Top of a Chair-back.

This Problem depends on no former Observations whatever: and as there is nothing in the Operation, or Instruments, to affect its Accuracy, but what any one may easily guard against, it may be reckoned the surest Foundation for all subsequent celestial Observations that require an exact Meridian Line. The only Disadvantage is, that it cannot be performed but in Winter, and when the Stars may be seen for 12 Hours together; which requires the Night to be about 15 Hours long.

*How*

*How to fix a Meridian-line by two circumpolar Stars that have the same Right-ascension; or differ precisely 180 Degrees.*

Pitch on two Stars that do not set, and whose Right-ascensions are the same, or exactly 180 Degrees different; take them in the same vertical Circle by a Plumb-line, and at the same Time let a Light be set up in that Direction, half a Mile, or a Mile off, and the Light and Plummets will be exactly in the Meridian.

In order to place a distant Light exactly in the Direction of the Plumb-line and Stars, proceed in the following Manner. Any Night before the Observation is to be made, when the two Stars appear to the Eye to be in a vertical Position, set up a Staff in the Place near which you would have the Plummets to hang; and placing your Eye at that Staff, direct an Assistant to set another Staff upright in the Ground, 30 or 40 Yards off, as near in a Line with the Stars as you can. Next Day, set the two Staffs in the same Places; and in their Direction, at the Distance of a Mile, or half a Mile northward, cut a small Hole in the Ground, for the Place where the Light is to stand at Night; and mark the Hole so, that a Person sent there at Night may find it.

Then chuse a calm Night, if the Observation is to be made without Doors; (if it is Moon-light, so much the better) and half an Hour before the Stars appear near the same Vertical, have a lighted Lanthorn ready tied to the Top of a Pole, and set upright in the distant Hole marked for it the Night before; and the Light will then be very near the Meridian, seen from the Place marked for the Plummets. At the same Time, let another Pole, or Rod, 6 or 8 Feet long, be supported horizontally where the Plummets is to hang, six or seven Feet from the Ground, and hang the Plumb line over it, so as to slip easily along it either to one Hand or the other, as there may be Occasion (or tie a Staff firmly across the Top of a Pole six or seven Feet long; fix the Pole in the Ground, and make the Plumb-line hang over the Cross-staff.) Let the Weight at the End of the Line be pretty heavy, and swing in a Tub of Water, so as it may not shake by a small Motion of the Air. Then shift the Plumb-line to one Hand, or the other, till one Side of

the Line, when at Rest, cuts the Star which is nearest the Pole of the World, and the Middle of the Light together: as that Star moves, continue moving the Plumb-line along the Rod, so as to keep it always on the Light and Star, till the other Star comes to the same Side of the Plumb-line also; and then the Plummets and Light will be exactly in the Meridian.

There are not two remarkable Stars near the North Pole, with the same Right-ascension precisely, or just a Semicircle's difference: but there are three Stars that are very nearly so, viz. the Pole-star,  $\epsilon$  in Ursa Major, and  $\gamma$  in Casiopeia. If either of the two last, particularly  $\epsilon$ , are taken in the same Vertical with the Pole-star, they will then be so very near the Meridian, that no greater Exactness need be desired for any Purpose in Surveying. At London  $\gamma$  is about  $1'$  West of the Meridian then; and  $\epsilon$  much nearer it, on the West Side likewise. Stars towards the S. Pole proper for this Observation are,  $\beta$ , in the Head of the Cross;  $\alpha$ , in the Foot of the Cross; and  $\alpha$ , in the Head of the Phenix.

*How to find a Meridian-line by a circumpolar Star, when it is at its greatest Azimuth from the Pole.*

By a circumpolar Star is here meant, a Star whose Distance from the Pole is less than either the Latitude, or Co-latitude, of the Place of Observation.

1. Find the Latitude of the Place in which you are to observe the Star.
2. Pitch on a Star whose Declination is known, and calculate its greatest Azimuth from the North, or elevated Pole.
3. Find at what Time it will be on the Meridian the Afternoon you are to observe; so as to judge about what Time to begin the Observation.
4. Prepare two Plummets, and follow the Star with one of them, (as directed in Page 47.) until the Star is stationary, then set up a Light half a Mile, or a Mile off, in the Direction of the Plumb-line and Star, and mark the Place of the Light in the Ground, and also of the Plumb-line.
5. Next day set up a Pole where the Light stood, with a Flag flying at it. Then with a Theodolite, or Hadley's Quadrant, set the Index to the Degree

H and

and Minute of the Star's Azimuth from the North, found before; direct, by waving to one Hand or the other, an Assistant to set up a Staff on the same Side of the Pole with a Flag as the Pole of the World was from it, so as at the Plumb-line these two Lines may make an Angle equal to the Azimuth of the Star, and the Plumb-line and Staff will then be in the Meridian.

When any Star is descending, it is on the West-side of the Pole of the World; while it ascends, it is on the East side of it.

The Pole of the World is always between the Pole-star and Ursa-major: so that when Ursa-major is W. or E. of the Pole-star, the Pole of the World is W. or E. of it likewise.

To find the greatest Azimuth of a circumpolar star from the Meridian, use the following Proportion.

*Cofine of the Lat. : R :: S. of the Star's Distance from the Pole : S. of its greatest Azimuth.*

The greatest Azimuth is when the Star is ~~below~~<sup>above</sup> the horizontal Diameter of its diurnal Circle ~~as many Degrees, (measured on that Circle)~~ as is equal to the Star's polar Distance.

On the N. Side of the Equator, the Pole-star is the most convenient for this Observation; for the Time when it is at its greatest Elongation from the Pole, may be known sufficiently near by the Eye; by observing when  $\epsilon$  in Ursa-major and  $\gamma$  in Cassiopeia appear to be in a horizontal Line, or parallel to the Horizon; for that is nearly the Time. Or, the Time may be found more precisely by making fast a small Piece of Wood along the Plumb-line when extended, with a Cross-piece at right Angles to the Top of the upright Piece, like the Letter T; when the Plumb-line is at Rest, and both Stars are seen touching the upper Edge of the Cross-piece, then they are both horizontal. It is another Convenience in making use of this Star, that it changes its Azimuth much slower than other Stars, and therefore affords more Time to take its Direction exact.

On the South-side of the Equator, the Head of the Cross is the most convenient for this Observation, being nearest to the S. Pole; and the Time of

its coming to the horizontal Diameter of the diurnal Circle, when it appears in a horizontal Line with the Foot of the Cross, or the Head of the Phenix; and <sup>at the Equator</sup> its greatest Azimuth one Hour sixteen Minutes before or after that, as it is E. or W. of the Meridian. The Pole of the World is between the Head of the Cross and the last of these Stars.

To find when any Star will come to the Meridian, either in the South or North. Find the Star's right Ascension, in Time, from the most correct Tables, also the Sun's right Ascension for the Day and Place proposed; their Difference will shew the Difference between their Times of coming on the Meridian in the South, or between the Pole and Zenith; which will be after Noon if the Sun's right Ascension is least, but before Noon if greatest. Eleven Hours 58 Minutes after the Star has been on the Meridian above the Pole, it will come to the Meridian North of it, or below the Pole.

*Given the Latitude of the Place, and the Declinations and right Ascensions of two Stars in the same vertical Line; to find the horizontal Distance of that Vertical from the Meridian; the Time one of the Stars will take to come from the Vertical to the Meridian; and the precise Time of the Observation.*

### C A S E I.

*When the two Stars are northward of the Zenith.*

### P L A T E I. FIG. XV.

Let  $ZP\odot$  be a Meridian,  $Z$  the Place of Observation,  $HO$  its Horizon, and  $ZV$  a vertical Circle passing through two known Stars,  $s$  and  $S$ ;  $P$  the Pole,  $EQ$  the Equator,  $PsD$  and  $PSd$  Circles of Declination (or right Ascension) passing through these Stars: then is  $ZP$  the co. latitude of the Place,  $Ps$  and  $PS$  the co. declination of the two Stars respectively, and the Angle  $sPs$  the nearest Distance of their Circles of right Ascension;  $VO$  the Arc of the Horizon between the vertical Circle and the Meridian; and  $dQ$  the Arc of the Equator between the Star  $S$  and the Meridian. When, in the Triangle  $PZs$ , the Angles  $PZs$  and  $sPZ$  (measured by the Arcs  $VO$  and  $DE$ ) are found, the Problem is solved.

## S O L U T I O N.

1. Begin with the oblique-angled Triangle  $P_s S$ ; in which are given two Sides  $P_s$  and  $P_S$ , the co. declinations of the two Stars respectively, and the included Angle  $S P_s$ ; which Angle is the Difference of their right Ascensions when it is less than 180 Degrees; but if their Difference is more than  $180^\circ$ , then the Angle  $S P_s$  is equal to the lesser Right-ascension added to what the greater wants of  $360$  Degrees. From hence find the Angle  $S_s P$  by the following Proportions.

*As Radius*

*Is to the co. sine of the given Angle ( $S P_s$ )*

*So is the Tan. of the Side opposite the required Ang. ( $P_s$ )*

*To the Tan. of an Angle, which call M.*

$M$ . is like the Side opposite the Angle sought, if the given Angle is acute; but unlike that Side, if the given Angle is obtuse\*.

Take the Difference between the Side ( $P_S$ ) adjacent the required Angle and  $M$ ; call it  $N$ . Then,

Sine  $N$  : Sine  $M$  :: Tan. of the given Angle ( $S P_s$ ) : Tan. of the required Angle ( $S_s P$ ) which is like the given Angle if  $M$  is less than the Side ( $P_S$ ) adjacent to the required Angle; but unlike the given Angle if  $M$  is greater than  $P_S$ .

2. Next in the oblique-angled Triangle  $P_s Z$  there are given two Sides,  $P_Z$  the co-latitude of the Place of Observation;  $P_s$ , the co. declination of the Star  $s$ ; and the Angle  $P_s Z$  opposite to one of them, which is the Supplement of  $P_s S$  (last found) to  $180^\circ$ : from thence find the Angle  $P_Z s$  opposite the other Side, by the following Proportion.

. Sine of  $P_Z$  : Sine  $P_s Z$  :: Sine  $P_s$  : Sine  $P_Z s$ . either acute or obtuse.

This Angle  $P_Z s$ , measured by the Arc of the Horizon  $V O$ , is therefore equal to the horizontal Distance of the Vertical of the two Stars from the

\* Two Arcs, or Angles, are said to be like, or of the same Kind, when both are less than  $90^\circ$ , or both more than  $90^\circ$ : but are said to be unlike, when one is greater and the other less than  $90^\circ$ : and are made like, or unlike, to another, by taking the Supplement to  $180^\circ$  of the Arc, or Angle, produced in the Proportion, in place of what the Proportion brings out.

Meridian.

Meridian. Let the Direction of the Vertical be taken by a Plumb-line and distant Light, as before directed, or by two Plumb-lines, and marked on the Ground. Next Day, the Degrees and Minutes in the Arc V O may be added to it by Hadley's Quadrant, and a Pole set up there, which will be in the Direction of the Meridian from the Pumb-line.

3. Last, in the same Triangle P's Z, find the Angle s P Z between the given Sides, by the following Proportions.

*As Radius*

*Is to the Tan. of the given Angle (P s Z)*

*So is the co. s. of the adjacent Side (P s)*

*To the co. t. of M.*

M is acute, if the given Angle and its adjacent Side are like; but obtuse, if the given Angle and the adjacent Side are unlike.

*As the co. t. of the Side adjacent to the given Angle (P s)*

*Is to the co. t. of the other Side (P Z)*

*So is the co. s. M*

*To the co. s. of an Angle, which call N.*

N is like the Side opposite the given Angle, if that Angle is acute; but unlike the Side opposite the given Angle, if that Angle is obtuse.

Then the required Angle s P Z is either equal to the Sum or Difference of M and N, as the given Sides are like or unlike.

The Angle s P Z, thus found, added to s P S, and their Sum subtracted from  $180^{\circ}$  will leave the Angle d' P Q, or the Arc d Q that measures it, which is the Arc of the Equator the Star S must pass over in coming from the Vertical, Z V, to the Meridian. Which converted into Time, and measured by a Clock, or Watch, beginning to reckon the precise Moment that a Plumb-line cuts both Stars, will shew the Hour, Minute, and Second that S is on the Meridian.

Find, by the right Ascension of the Star and Sun, at what Time that Star should come to the Meridian in the North the Night of the Observation; subtract from it the Time the Star takes from the Vertical to the Meridian, and the Remainder, corrected by the Sun's Equation, will be the Time when the two Stars were in the same Vertical.

The

The lowest of the two Stars comes soonest to the Meridian below the Pole ; the highest of them comes soonest to that Part of the Meridian which is above the Pole.

The nearer in Time one of the Stars is to the Meridian when the Observation is made, it is the better : for then an ordinary Watch will serve to measure the Time sufficiently exact. It is still more advantageous if one of them is above the Pole when the other is below it.

The nearer one of the Stars is to the Pole, and the farther the other is from it, the more exact will this Observation be ; because the Change of the Vertical will be the sooner perceived. For this Reason, in North Latitudes, Stars northward of the Zenith are preferable to those that are southward of it,

If the two Stars are past the Meridian when they are observed in the same Vertical, then the Arc  $d\ Q$  gives the Time  $S$  took to come from the Meridian to the Vertical ; and must be added to the Time when that Star was on the Meridian to give the Time of the Observation : and the Arc of the Horizon  $V\ O$  must be marked on the Ground on the Side of the Vertical contrary to what it would have been in the foregoing Supposition ; that is, eastward *below* the Pole, and westward *above* it.

When two Stars come to the vertical Line near the Meridian, it may be difficult to judge on which Side of it they are at that Time ; for determining this, the following Rules may serve.

The right Ascension of two Stars may be either each *less*, or each *more* than 180 Degrees, or *one more* and the *other less*.

When the right Ascension of each of the Stars is *either less, or more, than 180 Degrees*, they will come to the same Vertical on the *East-side* of the Meridian when the Star with the *greatest* right Ascension is the *lowest* ; but on the *West* Side of the Meridian when it is *highest*.

When

*When the right Ascension of one of the Stars is more than 180 Degrees, and that of the other less.*

If the right ascension of the highest is less than  $180^{\circ}$ , but greater than the Excess of the other's right Ascension above  $180^{\circ}$ , then they come to the Vertical on the *East Side* of the Meridian. But if the right Ascension of the higher Star is less than that Excess, they come to the Vertical on the *West Side* of the Meridian.

If the right Ascension of the higher Star is *more* than  $180$  Degrees, and that Excess is *less* than the right Ascension of the lower Star, then they come to the Vertical on the *West Side* of the Meridian ; but if the Excess of the higher Star's right Ascension *above*  $180^{\circ}$  is *more* than the right Ascension of the lower Star, then they come to the same Vertical on the *East Side* of the Meridian.

## C A S E II.

*When the two Stars are in the same Vertical southward of the Zenith.*

### PLATE I. FIG. XVI.

When two Stars are observed in the same vertical Line *southward* of the Zenith (or toward the depressed Pole) the Operations and Solutions are nearly the same as in Case 1. For let  $PZO$  be a Meridian,  $Z$  the Place of Observation,  $OH$  its Horizon,  $ZV$  a vertical Circle passing through the two Stars  $S$  and  $s$ ;  $P$  the Pole,  $QE$  the Equator,  $PD$  and  $Pd$  two Circles of Declination (or right Ascension) passing through the two Stars respectively : then is  $ZP$  the co. latitude of the Place,  $PS$  and  $P_s$  the co. declination of the two Stars respectively, the Angle  $SPs$  the Difference of their right Ascensions, the Angle  $SZQ$  (or  $OV$  the Arc of the Horizon which measures it) the Distance of the Vertical from the Meridian ; the Angle  $s.PZ$  (or  $d.Q$  which measures it) the Arc of the Equator, which  $s$  must pass over from the Vertical to the Meridian, which converted into Time, and measured by a Watch, will shew when  $s$  is on the Meridian ;

and substracted from the calculated Time that the Star should come to the Meridian, will (when corrected by the Sun's Equation) give the true Time of the Observation:  $s P s$  is the Triangle to begin the Solution with; and in the Triangle  $s P Z$ , the Angles  $s P Z$  and  $s Z P$ , when found, will give the Solution of the Problem, as in Case 1. For the Supplement of  $s Z P$  to  $180^\circ$  is the Angle  $V Z O$  (or its Measure  $O V$ ) the horizontal Distance of the Vertical from the Meridian; the Angle  $s P Z$  (or its Measure  $d Q$ ) is the equatorial Distance of the Vertical from the Meridian: for all which the Solution in Case 1. properly applied, will serve.

On the South Side of the Zenith, when the *highest* of the two Stars has the *least* right Ascension, they come to the same Vertical on the *East Side* of the Meridian; but when the highest Star has the *greatest* right Ascension, they come to the same Vertical on the *West Side* of the Meridian.

*How to fix a Meridian-line by the Equality of the Time of Ascent and Descent of a circumpolar Star toward the Meridian.*

For performing this Problem it is necessary to have a good Clock, or a good Watch, in the Room where the Observation is made, or in a Room adjoining to it. The Clock-case should be set perpendicular by a Plummeter, and the Back of it screwed firm to wooden Wedges, or to Holdfasts of Iron, driven tight in a Brick, or Stone Wall; or to an upright massy Post, or Beam, set two or three Feet into the Ground below, and supported higher up by inclining Stays on two or three Sides: for if a Clock is not fixed firm, but set up against an ordinary Wainscoting, or is liable to any Shake by Persons walking on the Floor, it will never be made to go exactly. There should likewise be some Contrivance to prevent the lengthening or shortening of the Pendulum by Heat and Cold. When the Clock has been placed perpendicular and immoveable, then it must be exactly regulated to keep equal Time, which may be done in the following Manner.

Take notice of any Star that will be hid from your Sight, any Time of the Night, by a Chimney-head, the Corner of a House, or a Post set firm in the Ground for that Purpose. Make a small Hole in a Window-shutter, or a Door, or in a thin Plate of Metal fixed in either of them, through which

which the Star may be seen. A little before the Star will be hid, stop your Clock, or Watch, with the Second-hand at 60; write down the Hour and Minute at which it was stopt; continue looking to the Star through the Hole; and the Instant the Star disappears behind the Chimney-head, or other Object, set the Clock or Watch agoing. If a Watch is made use of, it may be held in your own Hand, with your Finger at the Stopper, while you look to the Star through the Hole. If it is a Clock, you must have an Assistant ready to let go the Pendulum the Moment you call out, *Now*. Next Night, about the same Time, look through the same Hole, and watch the coming of the same Star toward the same Object, and the Moment it disappears, stop the Clock, or Watch, till you write down the Hour, Minute, and Second at which it was stopt; and if the Time between the two Observations shewn by it, amounts precisely to 23 Hours, 56 Minutes, 4 Seconds, the Clock, or Watch has kept Time exactly. If the Time shewn by it is more than that, then it goes so much faster in a Day than it ought to do: if the Time shewn is less than that, then it goes so much slower than it should do to keep equal Time: and the Error in a Day should be noted down, so as to be allowed for afterwards, in case the Clock or Watch cannot be brought precisely to equal Time, but may fall short, or exceed it, by a certain Number of Seconds each Day. If there is Time enough, it is better to let several Days, or a Week, intervene between the two Observations, and for each Day elapsed, to allow 3 Minutes 56 Seconds for the Acceleration of the Star's Revolution. That is, if the Clock or Watch goes equally, the Star will disappear behind the Object 3 Minutes 56 Seconds sooner the second Night, or in one Day, than it did the first; in two Days, it will disappear 7 Minutes 52 Seconds sooner; the fourth Night (or in three Days) 11 Minutes 48 Seconds sooner by the Clock or Watch than the first Night; and so in Proportion for any Number of Days; or, a little more precisely, according to the following Table.

When the Clock has been exactly regulated to go equal Time, then pitch on a Star which does not set, and is neither very near the Pole, nor too far from it. If it is very near, it will change its Vertical slowly, and the precise Time of its coming to it will not be so easily distinguished: if it is too far from the Pole, it will come to the highest Part of its diurnal Circle, too near the Zenith to be observed accurately at the Plumb-line. Some Star in the great Bear may be the most convenient. Hang two Plummets with heavy Bobs, swinging in Water, near a large open Window, or Door, facing the North, and as nearly in the Direction of the Meridian as you can; let the Plumb-line next your Eye be fixt, and the other moveable at pleasure from one Hand to the other. The Eye should not be too near the fixt Line, nor the farthest Line of too small a Size; for then they will appear misted, transparent and ill-defined, and the precise Time of the Star's Appulse will not be known; but the Distance of the Eye and the Thicknes of the Lines should, by previous Trial, be suited to distinct Vision. Stop the Clock at any Hour, with the Minute and Second Hands at 60, and write down the Hour they point out. Near the Time that the Star comes in the Direction of both Plumb-lines, let an Assistant hold back the Pendulum of the Clock, so as to be ready to let it go when you call to him: keep both Plumb-lines in one from your Eye, and the Moment the Star touches one Side of them, call out, *Now*; that your Assistant may immediately let go the Pendulum of the Clock. Take care the Plummets be not shifted in the least; and, near twelve Hours afterward, when the Star next approaches near to the Direction of the Plumb-lines, take them in one from your Eye, and let your Assistant watch the Second-hand of the Clock, counting each Second audibly by the Beats; and the Moment you see the Star at the same Side of the Plumb-lines as in the first Observation, call out, *Now*: your Assistant must immediately call out the last second he counted, so as you may know it and mark it down, while he looks to the Minute and notes it down; then write down the Hour, Minutes and Seconds all together. In the same Manner, near twelve Hours thereafter, when the Star approaches near to the Plumb-lines on the other Side of its diurnal Circle, take the Plumb-lines in one with your Eye; let your Assistant count the Seconds by the Beats of the Clock.

Days.	M.	S.	3ds.
1.	3..	55..	54
2.	7..	51..	48
3.	11..	47..	42
4.	15..	43..	36
5.	19..	39..	30
6.	23..	35..	24
7.	27..	31..	18
8.	31..	27..	12
9.	35..	23..	6
10.	39..	19..	—

Clock, and the Moment the Star touches the same Side of the Plumb-line as before, call out, *Now*; that he may repeat, audibly, the Second he counted last, which you must note down, while he looks the Minute, and notes it down. If the Time of the Star's Ascent on one Side of the Plumb-lines is exactly equal to the Time of its Descent on the other, then the Plumb-lines hang precisely in the Meridian. If these Times are unequal, shift the moveable Plummet a little toward that Side where the Star took the longest Time; and begin the same Operations again, and continue them till you find the Star takes equal Time on each Side of the Plumb-lines.

In Observations of this kind, where the Time is required to be known, it may be managed without an Assistant, after some Practice. For one may catch the Second from the Clock by his Eye a little before the Time of Observation, and continue counting the Seconds or Beats by the Ear, while his Eye watches the Star; and the Instant it is hid, he should repeat the last Second till he sees the Minute on the Clock, and then write down both together. Or, begin to count the Seconds by the Ear from the Moment of the Observation, till you can see the Time shewn by the Clock, from which subtract the Number of Seconds you counted, and you will have the Time of the Observation according to the Clock.

If the Pendulum of the Clock is contrived to keep of an equal Length, notwithstanding the Variations of Heat and Cold; if it is not wound up between the Observations, or does not stop in the Time of winding; and if it is well regulated immediately before the Observations begin, and again verified by the diurnal Revolution of a Star when they are over; a very exact Meridian Line may be found this way. But such Clocks are rarely to be met with; the Operation is often tedious; and is practicable only in Winter, and when the Sun does not set later than about half an Hour after four; for the Stars are seldom seen in Winter distinctly, till the Sun is near an Hour and a half set.

In this Operation, a Watch is managed much in the same Manner as a Clock; only, when it is to be stopt during the Observation, the Observer may do it himself, by holding the Watch in both Hands, with his Finger at

the Stopper. When the Time is to be noted without stopping the Watch, then, near the Appulse of the Star to the Plumb-lines, an Assistant must count the Seconds by the Motion of the Second-hand; and, when the Observer calls out, *Now*, must repeat audibly the last Second he counted, that the Observer may note it down, while the Assistant looks to and notes the Minute.

Though other Methods of finding a Meridian-line than those above, except the last, have been generally practised and recommended by Astronomers, yet it will be evident to Theorists, especially such as are accustomed to celestial Observations, that when the Places of the Stars are exactly known, this important Problem may be performed more easily, and with more Certainty, by means of a Plummet, *which adjusts itself*, than by any other Instruments, or Observations whatever on the Sun. For it is well known, that in solar Observations, the Air, the Object, the Instruments, the Observer, may each occasion some Inaccuracy, which the most Expert cannot altogether avoid; but they are either quite removed, or easily prevented, in our way, by the most inexperienced Performer.

The polar Distances and right Ascensions of these six Stars in 1770, were as follows: + signifies increasing, — decreasing.

	Char. and Mag.	Distance from the Pole.	Variation in 10 Years.	R. Ascension		Increases in 10 Years.	R. Ascension in Time.	Increases in 10 Years.					
				D.	M.	S.	'	"	H.	M.	S.	'	"
Northern Stars.													
Pole Star . . . . .	z. 2	0. 1. 56	— 3. 17	10.	44.	0	25.	0	0.	42.	56	3.	17
First in Tail of the G. Bear	z. 2	32. 47. 0	+ 3. 17	190	45.	5	7.	0	12.	43.	0	0.	27
Middle of Casiopeia's Chair	γ. 3	30. 32. 20	— 3. 17	10.	27.	16	9.	0	0.	41.	48	0.	35
Southern Stars.													
Head of the Cross . . .	γ. 2	19. 9. 0	— 3. 18	184.	29.	42	8.	3	12.	19.	0	0.	40
Foot of the Cross . . .	z. 1	28. 10. 0	+ 3. 20	183.	29.	0	8.	3	12.	14.	0	0.	32
Head of the Phenix . . .	z. 2	46. 30. 0	+ 3. 20	3.	43.	15	7.	30	0.	15.	0	0.	30

## P R O B.

*How to find the Sun's Amplitude at Rising, or Setting, and from thence the Variation of the magnetic Needle.*

Make the Needle level with the graduated Circle in the Box. Then, when the Sun's lower Edge is a Semidiameter above the Horizon, take the Bearing of its Center (from the N. or S. which ever is nearest) through the Sights, making the Thread bisect the Sun's Disk, and that subtracted from  $90^{\circ}$ , will be the Sun's magnetic Amplitude, or Distance from the E. or W. Points by the Needle.

Next, calculate the Sun's true Amplitude for that Day, by the following Proportion.

*As the co. sine of the Latitude,  
Is to Radius,  
So is the Sine of the Sun's Declination at Setting or Rising,  
To the Sine of his Amplitude from the W. or E.*

Which will be N. or S. as the Sun's Declination is N. or S: and the Distance in Degrees and Minutes between the true E. or W. and the magnetic, is the Variation of the Needle.

An easy and sure way to prevent Mistakes, which the Unexperienced are liable to in this Calculation is, to draw a Circle by Hand, representing the visible Boundary of the Horizon, and on it to mark the several Data by Guess; then by inspecting the Figure it will easily appear how the Variation is to be found, whether by Addition or Subtraction, and on which Side of the North it lies. For Example:

## PLATE I. FIG. XIV.

Suppose the Variation was sought at Sun-setting. Draw by Hand a Circle N W S E, to represent your visible Horizon: in the Middle of it mark the Point C, for your Station: from C, draw the Line C W, to represent the true West; then on the North or South Side of that Line, according as the Sun set northward or southward of the true West, draw the

Line

Line  $C\odot$ , representing the Direction of the Sun's Center at setting, and another Line  $Cw$ , for the magnetic West, either on the North or South Side of  $\odot$ , as it was observed to be, and at its judged Distance. Then by observing the Situation of these Lines, it will easily occur whether the magnetic Amplitude and true Amplitude are to be added, or subtracted, to give the Variation; and on which Side of the true North the Variation lies. In the present Supposition,  $w\odot$  is the magnetic Amplitude, and  $\odot W$  the true Amplitude;  $\odot W$  therefore must be subtracted from  $w\odot$  to give  $wW$  the Distance of the one from the other: and  $n$ , the magnetic N. 90 Deg. from  $w$ , must be westward of  $N$ , the true North.

This Method is sufficiently exact for finding the Variation; but it is not exact enough for fixing a precise meridian Line: because of the Uncertainty of the Refraction, and of the Sun's Center: but if the Sun ascends, or descends, with little Obliquity, the Error then will be very little.

*How to find the Sun's Azimuth, and from thence to find the Variation of the Needle.*

First let the Latitude of the Place be exactly found. Next let the Quadrant be carefully adjusted for Observation. Then, two or three Hours before or after Mid-day, take the Altitude of the Sun's Center as exactly as possible, making the vertical Wire of the Telescope bisect the Sun's Disk; and, without altering the Plane of the Quadrant in the least, move the Telescope vertically till you see some distant sharp Object on the Land, exactly at the vertical Wire; and that Object will be in the Direction of the Sun's Azimuth when the Altitude of its Center was taken. If no such Object is to be seen, let a Pole be set up in that Direction, about half a Mile off, or as far as can be seen easily.

Next calculate the Sun's Azimuth by the following Rule.

Add the Complement of the Latitude, the Complement of the Altitude, and the Complement of the Sun's Declination to  $90^\circ$  together, and take the half of that Sum, and note it down: subtract the Complement of the Declination from the half Sum, and take the Remainder; then take the Complement

plement arithmetical \* of the Sines of the Complement of the Altitude, and of the Complement of the Latitude, and add them together, and to them add the Sines of the forementioned half Sum and Remainder: half the Sum of these four Logarithms is the co. sine of half the Azimuth required. Therefore find by the Tables what Angle that co. sine answers to; double that Angle, and that will be the Sun's Azimuth from the North.

If the Sun's Declination is S. in *North Latitude*, or N. in *South Latitude*; in place of taking the Complement of the Declination to  $90^\circ$ , add  $90^\circ$  to it, and proceed as before.

In South Latitudes the Azimuth is found in the same Manner; only, the Sun's Azimuth is found from the S.

Then, to find the Variation, place your Needle below the Center of the Quadrant, set it level, and find how many Degrees the Pole, or Object, in the Sun's Azimuth, bears from the North by the Needle; and the Difference between that and the Azimuth found by Calculation, is the Variation of the Needle sought.

If the Sun ascends, or descends with little Obliquity, a Meridian Line may be fixed pretty exactly this way, because a small Inaccuracy in the Altitude of the Sun's Center will not be sensible in the Azimuth. But when the Sun does not rise high on the Meridian, this Method is not to be relied on when great Exactness is necessary; for then every Inaccuracy in Latitude, Altitude, and Refraction, occasions severally a greater Error in the Azimuth. To mark the Meridian-line on the Ground, place the Center of a Theodolite, or *Hedley's Quadrant*, where the Center of the Quadrant was when the Sun's Altitude was taken, and putting the Index to the Degree and Minute of the Azimuth, direct (by waving your Hat towards one Side, or the other) a Pole to be set up, making an Angle with the former Pole placed in the Azimuth, equal to the Sun's Azimuth found; and that

\* The Complement arithmetical of a Logarithm is found thus: begin at the Left-hand of the Logarithm, and subtract each Figure from 9, and the last Figure from 10, setting down the several Remainders in a Line; and that Number will be the arithmetical Complement required.

last-placed Pole will be in the Meridian, seen from the Center of the Quadrant.

*N. B.* The Sun's Declination in the Tables must be corrected by the Variation arising from the Difference of the Time between your Meridian and that of the Tables ; and also for the Variation of Declination for the Hours before, or after Noon, at which the Sun's Altitude was taken.

The two last Problems are constructed and explained in every Treatise of Navigation ; it is therefore needless to be more particular here.

P A R T

# MARITIM SURVEYING.

## P A R T II.

*Of the Procedure and Operations in surveying Sea-coasts,  
according to their various Circumstances.*

CHAP. I. A Stasimetric Scheme. 2. Procedure in surveying Coasts, under ordinary Circumstances. 3. In surveying Coasts unfavourably circumstanced. 4. Rocks and Shoals; Tides and Soundings; Descriptions; copying and reducing Draughts; Instruments. 5. Longitude; continuing a Meridian Line.

## C H A P. I.

*How to form a Stasimetric Scheme of Points, by which the Distances along the Coast may be determined.*

IN order to survey a Bay, Harbour, River, or any Part of the Sea-coast, with sufficient Accuracy and Expedition, three principal things are requisite. The first is, to measure a *fundamental Base-line*: the second, to form a *stasimetric Scheme* of Points, by which other Distances along the Coast may be determined: the third is, to delineate on Paper a *Figure similar to the proposed Coast*. The several Methods of finding the Length of a Base-line having been explained in Part I. Chap. 3. the next is, how to form the *stasimetric Scheme*; which may be done in the following Manner.

## P L A T E II. F I G. I.

1. Having first taken a general View of the Place to be surveyed from an Emiñence, in order to find out proper Objects for the Scheme, and to plan

K in

in your Mind the Procedure of the Survey; pitch on the smooth Sand X Y for a fundamental Base-line, as near the Middle of the Harbour, or Part to be surveyed, as can be had conveniently: on it measure the straight Line X Y with a Pole, or Chain, as directed p. 4; mark its Extremities on the Ground, and write the Number of Poles, or Chains, in a Field-book prepared for the Purpose.

2. At X, with a Theodolite, take the Angles, that some of the most conspicuous and sharpest Objects (such as Steeples, Towers, remarkable Tops of Hills, &c.) on each Side of the Bay, make with the Base-line; as Y X A, Y X B, Y X C, Y X D: and as many other Parts along the Coast as are distinctly seen, and may be judged convenient. If there are no remarkable Objects to be seen that are small enough, cause Turrets, or Signals, of Stone or Turf, to be built for that Purpose, on the most convenient Hills or Eminences. Also, at Y, take the Angles X Y A, X Y B, &c. then take the Bearing of X from Y, with a good magnetic Needle and Sights: write these Angles distinctly in the Field-book, together with the Name, or Description, of each Object, and the Part of it which you observe, so that they may be known again when needed, or when you go to them: write down likewise, on which Side of the Base-line each Object lies, whether on the right or left, when you look towards the farthest End of it.

3. Near the Middle of a Sheet of large Paper (so as there may be room for laying the Protractor, and drawing out the Angles) draw the magnetic Meridian right up and down. In a convenient Part of that Line, mark a Point X, to represent one Extremity of the Base-line; and at X, make an Angle with the Meridian equal to the Bearing of Y from it. On that Bearing set off the Length of X Y, in Miles and Parts of a Mile, taken from a Scale of equal Parts, not less than an Inch to a Mile.

4. From X and Y on the Paper, protract carefully, with a large Protractor, the several Angles written in the Field-book; and also calculate trigonometrically the most material Distances: judge of the Accuracy of the Protraction by its Agreement with the Calculation; and if any of them differ, protract, or calculate these again till the Error is corrected, and the Points representing the several Objects are marked in their proper Positions and Distances on the Paper.

5. To

5. To verify these protracted Distances, go to any of the Objects, as D, take the Bearing of X and Y, to find if they agree with the Protraction; if so, then take the Angle X D Y, add it to the two other Angles of that Triangle, and if their Sum is 180 precisely, then is the Position of D exact: there also take the Bearings, by the Needle, of the other Objects; and if their Bearings from you agree with their Bearings from D on the Paper, then your several Operations have been rightly performed, and the stasimetric Scheme is exact. If the Bearing of any of the Objects by the Needle differs from that of its representative Point on the Paper, take the Angles of that Object again from the two Extremities of the Base-line, and protract them anew, till you discover the Error, and find all agree: then, and no sooner, should you proceed to delineate the Coast-line on the Paper.

Note 1. Very high Mountains, whose Tops are often hid in Clouds, are on that account not so convenient Objects for a stasimetric Scheme, as Hills of a more moderate Height, which may be seen more frequently.

Note 2. As the Use of this stasimetric Scheme is for finding the Distances of other Parts along the Coast, by one or other of the longimetical Problems in Part I. therefore great Care ought to be taken to get the Distances and Positions of the Points in it exact: wherefore if any of the Objects lies too oblique to the Base-line, or is too far from it, to have its Distance determined precisely; let *one* of the Angles only be taken at the Base-line, and the other from one of the Objects whose Distance is ascertained, and makes a more direct Intersection with it.

Note 3. If no level Sand can be had within the Bay, or Harbour, for a Base-line; then a smooth Plane near it, on either Side, may be measured, and the Distances of the Objects on each Side of the Bay deduced trigonometrically from thence.

Note 4. If no Plane fit for a Base-line is in, or near, the Harbour to be surveyed, and if it is but a few Miles in Extent; then the Distance of Objects properly situated may be measured by the Velocity of Sound (as above directed) and that Distance made a Base-line, from which the stasimetric Scheme may be deduced.

Note 5. If the Bay, or River is above eight Miles in Extent, and no sufficient Plane in, or near it; then take the Latitudes of two remarkable Hills, or Objects, at least 15 Miles asunder, and not above 45 Degrees from the Meridian; from thence find their Distance in Miles and Parts of a Mile (by Prob. 2. page 12.) make that the Base-line, and from thence calculate and protract the Distances in the stasimetric Scheme.

*The Form of a Field-book.*

Measured in a straight Line on the Sand on the S. Side of the Entry of Great-Harbour, from X (at the High-water Mark of White-point) to Y (at the Edge of the Gras on Green-point) 416 Chains, 25 Links; each Chain 60 Feet long, and each Link one Foot; amounting to 4½ Miles and 408 Yards more.

At X, (by the Needle) Y bears — W.  $18\frac{1}{2}^{\circ}$  N.

*Angles taken at X, on the Side XY. Y bearing W.  $18\frac{1}{2}^{\circ}$  N.*

		Deg.	Min.
	a	Signal on Flat-hill	23 24
Right Side.	b	Steeple of New-church, and Point of a Clay-cliff about a Mile beyond it	36 15
	c	Sharp Rock on Black-head, and north End of a	
	d	Spring-tide Rock near the Middle of the Bay	41 48
	e	Castle-strong, W. Corner	9 15
Left Side.	f	Low rocky Point N. of that Castle	11 54
	g	E. Gavel of Wood-house	134 50

*Angles taken at Y, on the Side XY, X bearing E  $18\frac{1}{2}^{\circ}$  S.*

		Deg.	Min.
Left Side.	a	Signal on Flat-hill	64 40
	b	Steeple of New-church	57 22
	c	Sharp Rock on Black-head	95 16
		Hillock near Sandy-point, and Point of a high	
		rocky Cliff, supposed three Miles off	38 27
		W. Gavel of Cot-house	134 15
Right Side.	e	Castle-strong, W. Corner	163 10
	g	E. Gavel of Wood-house	18 49
		Ben-more Hill, the westmost and most tapering	
		Top; also a breaking Shoal about a League off	165 52

In taking the Angles at Y, take the same Part of each Object which was done at X; and before protracting, prefix the same small Letters to the Names of the same Places, or Objects. These Letters serve to mark the several Directions of the Angles in protracting, so that the Intersection of the corresponding Sides may be more readily found on the Paper.

## C H A P. II.

Examples of the Procedure in surveying Sea-coasts, under the most common Circumstances.

## E X A M P L E I.

*How to survey a Bay, Harbour, or River, and to delineate the Coast-line on Paper.*

## P L A T E II. F I G. I.

**I**N this Operation it will save much Time and Trouble, if the Surveyor has been accustomed to estimate, by the Eye, small Distances, such as a Mile, half a Mile, a Quarter, and 200 Yards, something near the Truth. This Dexterity may be acquired, sufficiently exact for the Purpose it is to serve, by frequently viewing with Attention such Distances when actually measured, and comparing other Distances with the Ideas of these in the Memory; especially when the Nature of the Ground compared, and the Darkness, or Clearness of the Weather, are nearly the same.

First, form a stasimetric Scheme of Points, as above delineated; and that the Points may be more distinct, and the Paper not incumbered with a Number of unnecessary Lines, transfer the magnetic Meridian, and the several ascertained Points in the Scheme, to a convenient Part of a large Sheet of clean Paper (or of several Sheets pasted together) so as there may be Room on it to contain and protract the rest of the Coast: then draw a small Black-lead Circle, or Dots, round each of these Points to render them more apparent, and write near each the Name of the Object it represents. Next, divide the Paper all over into small Squares, by Black-lead Lines about an Inch asunder, drawn parallel, and perpendicular, to the magnetic Meridian.

These

These Parallels serve for laying the Protractor North or South, East or West by the Eye readily, in any Point from whence a magnetic Bearing is to be drawn. If one Direction only of any other sharp Object has been taken, draw it out likewise from the Station at which it was observed; and write its Name along that Line, with a (D:) before it, to signify that it is the Direction of that Object; or an (L:) if it is the Limit of any Head, or Point. If Fig. 1. is the Harbour, the new stasimetric Scheme will then look as in Plate II. Fig. 2.

In the next Place; With this Scheme of Points and angular Directions, go to any Station on the Shore whose Place is marked on the Paper, as X, and there, with a good magnetic Needle, take the Bearing of a from it; draw it on the Scheme from X, the corresponding Point there, and judge its Distance as near as you can; take that Distance in your Compasses from the same Scale of equal Parts by which X Y was laid down; mark its Extent from X on the Bearing-line, and trace with the Point of your Compasses on the Paper the Curvature of the Coast between X and a, as exactly as you can: take also the Bearing of b from X, and draw it on the Paper. Next go to Y; at it take by the Needle the Direction of b, draw it on the Paper, and it will intersect the other Direction taken at X in b; trace the apparent Curvature between Y and b: take also the Bearing of a, and that drawn out will intersect its former Direction from X in the true Place of a; sketch the Curvature between a and b, and black-lead the whole from Y to X, which will finish that Part of the Coast-line. Next take the Bearing of c from Y, draw it out on the Scheme, and on that Line mark its judged Distance, and with your Compasses trace on the Paper the Curvature of the Coast between them. At c intersect that Bearing by the Bearing of C, or D, and correct that Part of the Curvature which is next c, and pencil the Tracing between c and Y: then take the Bearing of d, draw it on the Scheme, mark its judged Distance on that Bearing, and trace the Curvature between c and d with your Compasses: take also the Bearing of g, draw it on the Paper, and mark it so as to be found again when you come to g. At d take the Bearing of D, to intersect the former Bearing from c in d; correct the Curvature between them, and pencil it. Take also the Bearing of c, mark its judged Distance, and trace the Curve between them: at e, by the Bearing of D, intersect its Bearing from d, correct the Curvature, and pencil as before. In the same Manner, fix by magnetic Intersections the

the Points f and g, and trace the Curvature and Windings of the Coast between them. But because the Point h is more in the Way of Ships sailing up the Bay than any of the former Parts between c and h, its Position should be determined in the exactest Manner. Therefore at h, with a Theodelite, or Hadley's Quadrant, take the Angles A h B, B h C, and from thence (by Longim. Prob. 3.) find the Point of Station h; which will give its true Place independent of the former Stations. Next draw the magnetic Bearings of the Points i, k, l, m, n; intersect these Lines by the Bearing of C taken at each of them; delineate the Curvatures between them, and the Coast-line of the Bay will be so far finished. In this Manner proceed along the whole Shore of the Harbour, *first* fixing the Place you stand on by Intersections of the Needle; but at proper Distances, and Parts that are of Consequence to Navigation, let these be determined by two Angles taken more accurately, and protracted by one or other of the longimetical Problems in Part I.

If all the Stations along a Coast are determined by magnetic Intersections, the natural Defects of the Needle will, in Time, produce sensible Errors. If all the Stations are determined by the Angles between three Objects (according to a Longim. Prob. in Part I.) though that in Strictness is the most exact Way, yet it would be much more tedious, confuse the Draught with a Multitude of intersecting Lines and Arches, and answer no Purpose in Navigation better, nor in any Respect render a Draught more serviceable than if only some of the most material Stations, and principal Projections of the Coast here and there, were so determined, and the intermediate Parts found by Intersections of the Needle.

If, at any Part of the Coast, two, or three of the necessary stasimetric Objects cannot be seen, because of interposing Hills; or are not in a convenient Position for determining the Station wanted; go to some Eminence near that Station, from whence the Objects may be seen more advantageously, find the Point of Station on that Eminence by one of the longimetical Problems in Part I. protract it on the Paper, and by it, with one or two of the other Objects, find the Station wanted on the Coast.

If at any Station, the Angle between two Hills, or Objects, is wanted, but interposing Eminences, or other Objects, hinder one of them to be seen; in that Case, go to the Top of the interposing Eminence till you

see the Hill and Station from it, and exactly in a Line with them set up two Poles; and the Angle between any of the Poles and the other Object, or Hill, taken at the Station, will be the same with the Angle between the two Hills. Thus may the Angle, or Bearing of any Objects, be found from a Place where they cannot be seen, which will be necessary or convenient on many Occasions.

Mountains, Rocks, and Stones, are frequently endowed with a magnetical Quality, which disturbs the Polarity of a Needle, and, if not adverted to, will create much Trouble. To prevent this, it is necessary to carry two Needles always along with you, and to set both down at each Station, beyond the Reach of each others Influence; if they both agree in the Bearing of the same remote Object, it may be concluded that they are not disturbed by Magnetism.

At every Station take Care to insert, and to distinguish in the Draught, the Appearances and Nature of the neighbouring Parts of the Coast, such as rocky Cliffs, grassy Cliffs, sandy Shores, rocky Shores, &c. and remarkable Hills, Houses, Trees, &c. but especially such as may serve for Landmarks to direct Ships through Channels, or to avoid Rocks and Shoals. The Hills and Houses ought to be represented in such Prospects as will enable Sailors to know them when they are seen from the Sea, without regarding much the Diversity of Perspective which this will occasion in the same Draught; for Uniformity ought always to give Place to Utility.

If the Survey is to be carried on without the Harbour, along the Coast on each Side, it may be continued on the Foundation laid before, and by the same Procedure, to as great a Distance as the stasimetric Objects can be easily seen, and while the Angles they subtend are not too acute to be relied on, or protracted with Certainty. Before this happens, it will be necessary to pitch on other stasimetric Objects as you go along the Coast, to determine their Distances carefully by one of the longimetric Problems, and make these supply the Places of such of the former as are too remote; and by these last, others still farther along may be determined, to serve the same Purpose to a greater Extent. But let it be adverted here, that if too long a Series of Triangles is continued from one Base-line, the unavoidable Imperfection of Instruments,

struments, Observations, and Protraction, will, in Time, produce very sensible Errors: therefore it is adviseable to break off the Series at proper Distances, and to lay a new Foundation by measuring another Base-line, and proceeding by it as at first. When this is done, remember that two Objects, or Points, in the first Draught, at a proper Distance and Position, be likewise included in the second; that by these Points the two Draughts may be easily joined together, and make one Map when it is needful.

When the Coast of a Bay, Harbour, or River is surveyed, or any considerable Part of it, then find the Distance and Extent of all the Rocks, Sand-banks and Shoals that lie near it; observe Leading-marks to them; and Land-marks for avoiding them; sound the Depth of the Water to, round, and from each; and also along the Shore, as near to it as Vessels can, or ought to sail. Also sound particularly the several Places of Anchorage, and take Land-marks, or Bearings for finding them, and insert an Anchor in the proper Part of the Draught, for pointing them out more readily. Insert also the Time of High-water on the Days of new and full Moon, the ordinary Rise of spring and neap Tides, the Direction of the Stream of Flood, and what Irregularities may be observed in these, both within and without the Harbour.

### E X A M P L E II.

*How to survey an Island.*

CASE 1. *If the Island has several conspicuous Hills in it:* first measure a Base-line on a Plane, or by two Latitudes. 2. On that Base-line form a stasimetric Scheme of Points, as was explained in Part II. Chap. 1. § 3. Survey and delineate the Coast of the Island as directed for a Harbour in the last Example.

### P L A T E III. F I G. I.

CASE 2. *If the Island is but a few Miles long, and has but one remarkable Hill, or Object in it;* as A, which may be seen from all, or most Parts of the Coast: by one of the longimetreal Problems in Part I. find the Dis-

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tance

tance of any of the Points, or Heads, in the Island (suppose B) from A. At B, take the magnetic Bearing of A, in Degrees and Parts of a Degree; lay down that Distance on Paper according to the Bearing observed; and draw black-lead Lines all over the Paper, parallel and perpendicular to the magnetic Meridian, about an Inch from one another. At B, with a Theodolite, or Hadley's Quadrant, take the Angles A B C and A B F; write them down first in your Field-book, and then draw them out from the Point B: also take the Bearing of a and b by the Needle; judge of their Distances from B; draw them, and sketch the Curve of the Coast between you and them. Then go to C; there take the Angles A C B, A C D; write them down; add A B C and A C B together, subtract their Sum from  $180^{\circ}$ , and the Remainder will be the Angle B A C; draw that Angle from the Point A, and it will intersect the Side B C in the Point of Station C. At C, take the Bearing of a, and that will intersect its former Bearing from B in the corrected Place of a: then sketch the Curvature between C and B, and pencil it. Next draw the Angle A C D; at D, take the Angles A D C, A D E; the former added to A C D and subtra&ted from  $180^{\circ}$  will leave the Angle C A D; draw that Angle, and it will intersect the Side C D in the Station D: sketch and pencil the Curve between C and D. Then draw the Angle A D E, and the Bearing D c; fix the Point E by the Angle at E, and finish the Bay between D and E as before. Go on in this Way from Point to Point round the Island, till the whole is finished, remembering to insert remarkable Cliffs, Houses, or other Objects at each Station.

If at any Point, as G, or H, A cannot be seen; take the Direction of A from that Point, by setting two Poles exactly in a Line between them, as mentioned in Page 72, and take the Angle between one of the Poles and the other visible Object. By such a Procedure as this, the Circumference of any Island will be found to meet on Paper, without arbitrary Alterations.

When an Island is small, and great Exactness unnecessary; if the first Distance, or Base-line, is true, and long enough, all the other Distances may be got sufficiently exact by Intersections of a good Needle, without any other Instrument to take the Angles.

When

When one Side, or Part, of the Island is delineated on the Paper, then find the Distance and Dimensions of the several Rocks, Sand-banks, Shoals and Ledges along it; found the Depth of the Water towards and round each of them, and take Leading-marks on them, and Land-marks for avoiding them: found also along each Part of the Coast after it is delineated, as near the Shore as Vessels can, or ought to sail; and when the whole Coast is finished, found round it a Mile, or more, from the Land.

## EXAMPLE III.

*How to proceed in surveying an extensive Coast.*

CASE I. If the Coast extends northward, or southward, take, carefully, the Latitudes of two remarkable Hills, or Promontories along the Coast, as near the true Meridian as they can be found, and as far from each other as can be seen distinctly: suppose 20, 30, or 40 Miles. From their Difference of Latitude, Bearings and Variation of the Needle, find their Distance in Miles and Parts of a Mile (by longim. Prob. 2. p. 12.) make the Chord of that Arc, or Distance, the Base-line, and by it form a stasimetric Scheme of Points; one, or two of them, representing remarkable and sharp Objects. If one, or more, of the Objects lie off the Coast at Sea, it will be a Convenience; for there will be fewer Objects to intercept the Sight of these. When the stasimetric Scheme is verified, and transferred to some Sheets of clean Paper, and a Number of magnetic Meridians, and east-and-west Lines, drawn over it, then begin to survey and delineate the Coast, as directed in Example II.

It will sometimes happen that no one proper Object is to be seen from both Ends of so long a Base-line, with which to form a stasimetric Scheme; but if any remarkable intermediate Object, properly situated, is seen at one End of the Base-line, and the other End seen at that Object, its Distance may be found with equal, or rather more Accuracy, by taking one Angle of the Triangle, formed by the Base-line and Object, at that End of the Base-line where it is seen, and the other Angle at the Object; thence the third Angle is found; and the Distance of the Object. These three determined Distances will be sufficient for finding all other Distances between

the two Extremities of the Base-line ; and also for determining other stasimetric Objects necessary for continuing the Survey far beyond these Extremities.

Let it be observed, that if the Quadrant with which the Latitudes of the two Places were found, and the Observations, are good, a Base-line of 30, or 40 Miles measured in that Manner, is more to be relied on than such a Distance determined by a Base-line of three or four Miles measured on a Plane : because it is rare to find Planes so long without some sensible Irregularities in them ; and more rare to meet with two so remote Objects sharp enough for taking the Angles at each End of such a Base-line with Precision. These two Sources of Inaccuracy may occasion a greater Error in a Distance of 20, or 30 Miles, than can be supposed when the Latitudes are carefully taken with a good Quadrant.

When the Survey has been continued by a Train of stasimetric Triangles a considerable Length beyond the Base-line, it will then be proper to discontinue the Procedure on that Foundation, and to measure a new Base-line ; either on a level Plane, or by two Latitudes, as before ; taking Care to have two determined Points common to both Draughts, for connecting them into one.

If the Instrument with which the Angles are taken does not give them minutely enough ; or if the Objects that form the stasimetric Triangles are not sharp enough, Errors may be expected, and will undoubtedly become sensible in a long-continued Series of Triangles. How far they have actually taken Place in the Draught may be discovered by comparing the observed Bearings of distant Hills or Head-lands, whose Positions have been determined by former Observations, with their Bearings in the Draught ; or, by comparing such moderate Distances as one can judge of by the Eye, with their protracted Distances on the Paper.

When a considerable Length of the Coast has been surveyed, the Soundings marked near it, and all the Rocks, Shoals, Banks, remarkable Hills, Buildings, Groves of Trees, and other Distinctions of the Coast inserted and expressed in the Draught ; then sail along it, six, eight or ten Leagues from

## Chap. 2. MARITIM SURVEYING.

the Land, according as it can be seen distinctly ; sound the Depth of the Water, observe the Setting of Tides and Currents, and sketch Views of the Coast as you sail, inserting in them the Names of the most material Hills, Heads, Entries of Rivers, Harbours, &c. so that Seamen may know, by the Eye, where the principal Places on the Coast lie, and how to steer for them.

*CASE 2.* If the Coast to be surveyed extends eastward, or westward ; chuse a remarkable Hill, or Head, near the Coast, and another Hill, or remarkable Object up the Country, northward or southward ; find their Distance by measuring a Plane, or by the Latitudes, and make that Distance the Base-line ; from thence form a stasimetric Scheme of Points ; and with it proceed to survey and delineate the Coast as before directed.

If any Part of a Coast that extends eastward or westward is so circumstanced, as neither to have in it a level Plane fit to be measured, nor any Hill or remarkable Object up the Country, or lying off the Coast, far enough distant to become a Base-line by taking their Latitudes : in that Case, build a Wall or Turret, of Earth or Stone, on the most conspicuous Part of the Shore ; and another Turret three or four Miles from it up the Country, and so large as to be seen five or six Miles off, or farther : measure the Distance of the two Turrets by the Velocity of Sound, and make that a Base-line from whence to determine trigonometrically the Distance of the other Stations and Signals set up along the Coast for that Purpose : from these last, find other Distances : then, if you meet with no Plane fit to be measured, nor any remarkable Hill or Object, at a sufficient Distance and Position for determining a new Base-line by the Latitudes ; measure a new Base-line by Sound, and proceed as before. Such a Case as this is very rare ; but when it happens, a good portable Telescope, or Spy-glaſs, will be found convenient, and should be provided accordingly.

That the maritim Survey of a Kingdom, or large Tract of Continent, may be carried on with Expedition and Accuracy together, it is necessary one Superintendant, or Head-surveyor, expert in Theory and Practice, should have two Assistants under him, who are capable of executing his Orders. Their Duty is, to conform to his Directions diligently and faithfully, as far as can be done ; to omit no Part of the Coast, nor neglect any Rocks, Shoals,

Shoals, Channels, Tides, or necessary Soundings; to be at Pains to get Information concerning them from the Inhabitants, or Pilots, wherever they come, but to insert nothing in their Draughts but what has been actually examined by themselves; to keep a daily Journal of their Operations, Observations, and likewise of what Informations they may receive from others that have not been examined by themselves.

The *Head-surveyor's* Duty is, to plan and direct the Procedure of the whole Survey; to order the Vessel, Boats and Men on the Service when and where he sees it necessary; to chuse proper Planes and Distances for measuring fundamental Base-lines; to see the Mensuration, or celestial Observations himself; to pitch on proper Objects for the stasimetric Scheme; and see the Angles taken that determine their Distances; to inspect the Calculations and Protraction; to verify the Scheme when protracted; to cause a clean Copy of it to be made out for himself, and one for each of the Assistants; to send one of them to survey on one Side, or towards one End of the Base-line; another on the other Side or End of it, and himself to remain with the Vessel that attends the Survey, and to survey in that Neighbourhood; to examine their several Performances when they return to the Vessel; to compare the most material Distances in their Draughts with the Observations by which they were determined; to point out Mistakes, or Defects, and cause them to be corrected; to insert the several Observations, Measurements, Descriptions, and Sailing-directions regularly in a Book; to join the several Parts of the Coast, as they are compleated, into one Draught; and when that is of a sufficient Extent, to cause a clean, distinct Copy to be made of it aboard: then to sail in the Vessel to the next stationary Harbour; to cause Soundings, and useful Views of the Coast to be taken by the Way, and such remarkable Objects on Land to be inserted in the Draught, as may have been omitted by the Assistants. There will be no great Advantage in having more than two Assistants under one Superintendant; for this would often occasion either Delays in waiting for one or other of them before the Scene of Operation could be shifted; or else a superficial Inspection of their Performances.

Toward the End of Harvest, when the Days are turning short, and bad Weather may be expected more frequently, a Survey will be sooner dispatched, if the Examination of Shoals and Sand-banks, that lie at a Distance

from

from the Land, is postponed till the End of the following Spring; and the Survey of the Coast *only*, and the Soundings near it, or of Rivers and narrow Arms of the Sea, are taken in the Winter and Spring Seasons.

Though in general it is better that a considerable Part of a Coast be surveyed before the Soundings are taken near it, yet often both may be dispatched together with sufficient Exactness; by making an Eye-sketch of the small Bays and Windings of the Coast as you go from Point to Point in a Boat to determine their Distances, taking the Soundings by the Way, and inserting them in the corresponding Parts of the Sketch as near as you can judge; and at the same Time marking down the Direction of the Boat, or on what Object her Head is kept, in sailing or rowing from Place to Place; and the Bearing of one or two Objects when the principal Soundings were taken: when that Part of the Coast is surveyed, these Soundings may then be transferred to the Draught by the Bearings, and by the Direction in which the Boat was steered. If any Shoals are met with, take Marks on them, or two contiguous Angles by *Hadley's Quadrant*, and examine them more particularly afterward, when the Survey of the adjacent Part of the Coast is finished.

#### E X A M P L E IV.

*How to proceed in surveying a large Cluster of Islands.*

#### P L A T E III. F I G. II.

Pitch on two remarkable tapering Hills, Rocks or other Objects, at a competent Distance, as A and B, as near North, or South, of each other as you can; and if a Line connecting them runs about the Middle of the Cluster it is the better. Find the Distance of A and B, either by measuring a straight Line on a Plane, or by their Latitudes, and make that the Base-line: select amongst the other Islands one, or more, remarkable Objects on each Side of the Base-line; as C and D; of these form and verify a stasimetric Scheme, as before directed. With this Scheme go to one of the Islands, and by longim. Prob. 4, or 7, find the Distance of two proper Parts, or Points in it; make that a Base-line for surveying it, as explained in Example 2. Do the same for each Island, sounding the Channels between them, and observing the Setting and Strength of the several Streams of Tide as

you pass and re-pass, taking Notes of all and inserting them in the Draught at the Time they are observed. When all the Islands are surveyed and Channels sounded, and every necessary Particular is inserted in the Draught, then sail round the whole, between four and ten Leagues from the Land, and found the Depth at that Distance, and draw Views of the Land as you sail.

If the Islands extend northward and southward in a narrow Chain, (as in Fig. 3.) and are so small, or so uneven, that no sufficient Base-line, for the whole, can be measured on either of them ; then find by their Latitudes the Distance of two of the most conspicuous Hills, or Rocks, that are situated along one Side of the Cluster, and at a sufficient Distance from each other, as A and B. At A and B with *Hadley's* Sextant, or a Theodolite, take the Angles B A C, B A D, A B C, A B D, and from thence and the Side A B, find the Places of C and D any two remarkable Objects on the other Side of the Cluster of Islands, which may make the Angle B A C and A B D as near to right Angles as may be conveniently had. From any three of these Points the Position of a Place on any of the Islands may be found ; or by Intersections of the Needle from *two* advantageously situated ; and each Island delineated separately, as before directed.

### C H A P. III.

Examples of the Procedure in surveying Sea-coasts under unfavourable Circumstances.

#### E X A M P L E I.

*How to survey small Islands that extend East or West in a long narrow Train.*

#### P L A T E III. F I G. IV.

**I**F the Islands are small, and extend eastward or westward in a long narrow Train, as Fig. 4. then no sufficient Base-line will be found either by measuring a Plane, or by the Latitudes of any two Islands : in that Case, if the whole Extent of the Islands is not above 12 or 15 Miles, a Base-line may

may be measured by the Velocity of Sound, a stasimetric Scheme protracted from thence for determining the Distances of the several Islands from one another, and each surveyed separately, as before directed. But if the Islands extend above 15 Miles in Length, such a Procedure will not be found sufficiently exact, because a Distance cannot be measured by Sound with Precision: then the following Method is more to be relied on.

Pitch on the most remarkable Head, Hill, or Object that is to be found near the Middle of the Range of Islands, as X: find its Latitude, and the Variation of the Needle there. Then five or six Leagues northward, or southward of X, anchor a Vessel, as at V: on a prefixed Day, let the Latitude of the Vessel be taken aboard with *Hadley's Quadrant*, as exactly as possible: and at the same Time, while the Vessel remains steady in the same Place, let the Angles X V C, X V D, be also taken with *Hadley's Sextant*, and as many more Angles between them as there are remarkable Objects to be seen. About Mid-day, while the Sun's Altitude is taking on board, let one Assistant at X find the magnetic Bearing of the Vessel from thence; and also set up a Pole half a Mile, or a Quarter of a Mile from it, exactly in the Direction of the Vessel; and let the Place of the Pole be marked on the Ground, for future Use. Let two other Assistants be stationed one at A, and the other at B, four or five Miles on each Side of X, and there mark the Direction of the Vessel from them at Mid-day, by Poles in the Ground as was done at X, so that the Angles V X A, V X B and X B V, X A V, may be taken afterwards at Leisure, though the Vessel is removed. Find the Distance of X V (by longim. Prob. 2.) make that the Base-line; on it protract the several Angles observed on Shore, and aboard, and form a stasimetric Scheme of Points; in which the Stations X, V, A, B, will be determined; and also one Direction of C and D, and of the other Objects observed at V. From the Angles which two, or three of these Objects subtend, the Distances of the other Islands, or of any Station in them, may be found, and each Island surveyed separately.

If the whole Extent of the Islands cannot be included in the Angles taken at V, the same Operations must be begun again where the former ended, and the rest of the Islands surveyed in the same Manner.

## EXAMPLE II.

*How to proceed in surveying a Coast covered with Wood, or Bushes, so that in-land Objects cannot be seen from the Shore.*

## PLATE III. FIG. V.

CASE 1. *If it is a River, Bay, or Harbour;* find stasimetric Objects on the Side opposite to that which you are to delineate, and by these determine your stationary Distances by two Angles, or by Intersections of the magnetic Needle.

CASE 2. *If it is an open Coast, that has no Land opposite to it near enough.*  
 1. Measure a Base-line on a Plane, or by two Latitudes, and from thence find the Distances of three remarkable in-land Hills, or Objects, as Y, R, W, that may be seen from a Boat at Sea, over the Trees; and protract a stasimetric Scheme on Paper including these three Objects. If three remarkable Objects cannot be had; two conveniently situated for magnetic Intersections along the Coast will serve.

2. Prepare eight Poles, each with a flag flying at its Top for Signals; two of the Flags white, two red, two yellow, or striped, and two black: the white to represent one of the stasimetric Objects, the red another, the yellow or striped a third, as shall be agreed on between the Surveyor and his Assistant; the black any new stasimetric Object, whose Place may be found necessary to be determined as the Survey goes on. Let these Colours be written on the stasimetric Scheme, each at the Point, or Object, signified by it; and a Copy of the Scheme given to the Assistant to direct him, and prevent Mistakes; together with four of the Signals, each of a different Colour: the other four are to be carried along with the Surveyor.

3. At the first Point, or Station on Shore, whose Distance is to be determined, as A, let the Surveyor set up a Signal of a Colour indicating what stasimetric Object is to be first observed, as Y, or yellow. Then the Assistant must go off in a Boat to X, as far from the Shore as to see the stasimetric Objects over the Trees, and take that Signal in a Line with the Object it denotes, keeping the Boat steady at Anchor in that Place, and putting up

up a Signal of the same Colour with the Surveyor's, to signify that the Boat is then in the right Place. Let the Surveyor then take exactly the magnetic Direction of the Signal X in the Boat; which will also be the Direction of his Station B from the indicated stasimetric Object Y. Let it be drawn out on the Paper from its proper Point Y, and a Signal of another Colour R, set up in the Place of the former, to signify to the Assistant to move the Boat till he takes that Signal in a Line with the Object R it represents; and there let the Assistant, when ready, put up the like Signal at Z, and keep the Boat at rest till the Surveyor has taken its magnetic Direction, and drawn it on the Paper; which will intersect the former Direction in the Point of Station A, on the Shore.

4. Let the Surveyor then take the magnetic Direction of the next Point of Land B, along the Shore; draw it on the Paper; estimate its Distance by the Eye; mark that Distance on the Line of Direction; sketch with the Point of his Compasses the Figure of the Curvature of the Coast between the two Points A and B; take down his Signal and wave it, for a Sign to the Boat to move towards the next Point B, on the Coast.

5. When the Surveyor comes to that Point, let him put up a Signal W, or White, intimating what stasimetric Object the Assistant is to take in a Line with it; which when he has done, let the Assistant set up the like Signal in the Boat, and keep it steady at Anchor till its Bearing is taken on the Shore; which will intersect the Line of Direction taken at A the first Station, in the corrected Distance of B the second Station: let the Curve of the Coast between them be also corrected, and drawn with Black-lead, and the Coast-line between these two Points of Land will be finished.

Let the Distance of the next Point C be determined, and the Coast-line sketched, in the same Manner; and go on thus till the stasimetric Objects can be seen by the Surveyor on Shore, when there will be no Occasion for the Boat.

If any of the stasimetric Objects shall happen to be hid from the Assistant's Sight by interposing Hills; he must then first set up the Surveyor's Signal, and above it set up that Colour which points out some other stasimetric Object that is seen: which when the Surveyor observes, he must set up

another Signal of the same Colour above his former one, and then take the Direction of the Signals in the Boat; which will be the Direction of his Station from the Object represented by the uppermost Colour.

When any of the stasimetric Objects begins to disappear, or to be too remote, or oblique, for Accuracy; then, when the Surveyor is at a proper Station, the Assistant must set up the black Signal; to signify that the Place of a new stasimetric Object is to be determined. The Bearing of the black Signal taken when the Boat is at Rest, and drawn out from the Surveyor's Point of Station, in the opposite Direction, will be one Direction of the new stasimetric Object. When the Surveyor comes to another Station proper for another Intersection, the black Signal must be set up again by the Assistant: and when the Boat is at Rest, the Surveyor is to take its magnetic Direction, which will intersect the former in the Place of the new stasimetric Object; which must continue to be pointed out thereafter by the black Signal; till, by Agreement, it shall be changed for another Colour, that the black Colour may be again reserved for another new Object, in case it shall be necessary.

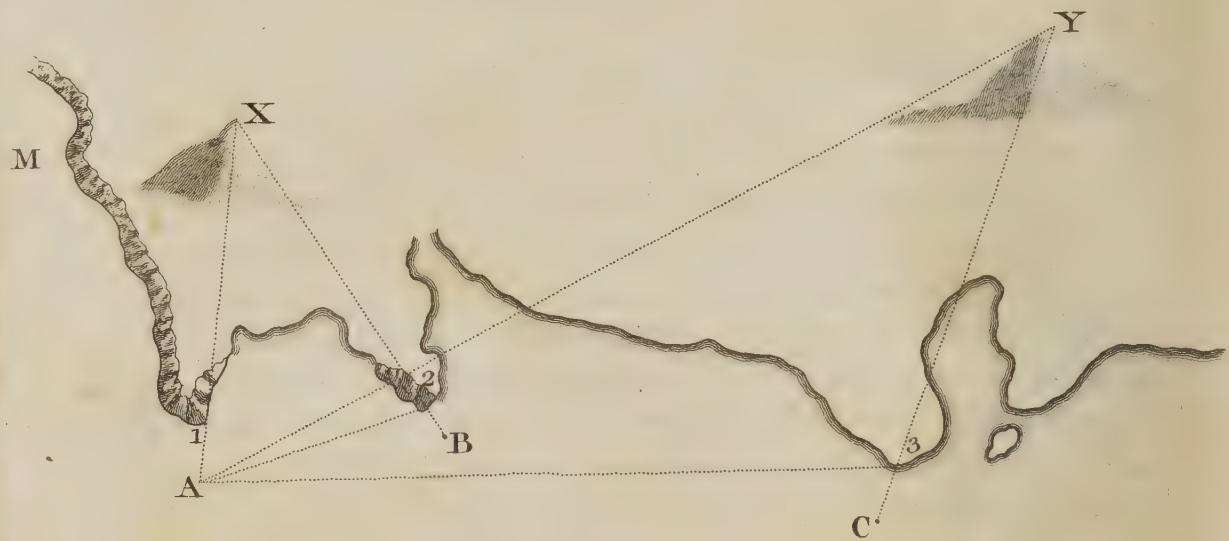
As the Assistant goes along the Shore in the Boat, he may at the same Time sound the Depths of the Water, and mark them on an Eye-sketch of the Coast; or insert them by their Bearings from the Stations, and other remarkable Parts of the Coast.

For this Manner of Surveying, it is necessary that the Assistant should be tolerably acquainted with the Business, that he may be capable of making a proper Choice of Objects, and Stations for determining their Distances.

This Method of surveying a Coast is sometimes necessary; and though not so exact as when the Angles can be taken with a Theodolite, or *Hadley's Quadrant*; yet when the stasimetric Objects are advantageously situated, and the Coast so surveyed is not very extensive, it will be found sufficiently exact for the Purposes of Navigation. If any particular Part of such a Coast requires great Exactness, then three Assistants, and three Boats, must be employed at the same Time; each must have the same Sort of Signals, and they must take the Surveyor's Signal bearing on three different stasimetric Objects.



A Coast extending Northward & Southward to illustrate Example III. Page 85.



X and Y are two tapering Hills (mentioned in Article 2) whose Bearing may be taken at M, and their Distance found by their Latitudes and Variation of the Compass.

1, is the remarkable Promontory at which the Delineation begins (mentioned Article 5) 2 and 3, the second and third Points of the Coast (mentioned Art. 6 and 7)

A, B, C, are the three Stations of the Boat near these Points.

The dotted Lines are Compass-bearings taken at the several Stations, by whose intersections the Places of the Boat and Distances of the Points of Land are determined.

Objects at once; by which Means he may take the Angles between them by a Theodolite, or by *Hadley's Quadrant*, which will be the same with the Angles between the three stasimetric Objects; and from thence may find the Place of Station precisely, by longim. Prob. 4.

If there is any Stream of Tide, or Breeze of Wind, where the Boat lies, it will not be kept steady by one Anchor only, but must be made to ride by two, in order to continue in the same Place while the Surveyor makes his Observations on Shore.

### E X A M P L E III.

#### *How to survey a Coast, without going ashore on it.*

There are some Coasts where it is impracticable to land for a continued Swell and Surf on the Shore; and others, where the Barbarity of the Inhabitants renders it unsafe; the Survey of such Places must be made altogether at Sea; and may be done in the following Manner.

#### *CASE I. If the Coast extends northward and southward.*

1. Find the Variation of the Compafs.
2. Chuse two remarkable tapering Hills, or other sharp Objects, 20 or 30 Miles asunder, and as near the same Meridian as they can be had; but so situated, that by sailing along the Coast or into a Bay, or River, one of them may be taken in a Line with the other, so that their mutual Bearing may be exactly taken with a good amplitude Compafs.
3. The Variation being found, bring a Ship to an Anchor right East, or West, from each Hill or Object, on different Days, and there observe their Latitudes with *Hadley's Quadrant*; which will be the Latitudes of the two Hills respectively.
4. From these Latitudes and their Bearing, find the Distance of the two Hills, by longim. Prob. 2. and protract a stasimetric Scheme, in which these two Points will be given.

5. In

5. In a rowing Boat of 8 or 10 Oars, and in moderate Weather, go as near to some remarkable projecting Point, or Promontory, as you can with Safety; there anchor your Boat, and with an amplitude Compafs, constructed and balanced so as to be as little affected by the Motion of the Boat as possible, take the Bearing of the two stasimetric Objects; draw these out from their correspnding Points on the Scheme, in the opposite Direction, and their Intersektion will be the Place of the Boat. Take the Bearing of some remarkable Part of the Point, or Promontory next you; judge its Distance as near as you can; draw that Bearing on the Paper from the Point found for the Boat's Place; and on it mark the estimated Distance of the observed Part of the Point or Promontory; which will be one Point in the Coast to be surveyed, and sufficiently exact for the Purposes of Navigation; the Error being only what may arise from estimating a Distance that is not above a quarter or half a Mile off.

6. In the Boat, take the Bearing of the next Point along the Coast, from the Point or Promontory determined; draw it on the Paper; mark its estmated Distance on that Line; and with the Point of your Compasses trace on the Paper the Figure of the Coast between these two Points of Land.

7. Go to the Point last observed, as near as you can; take its Bearing from one of the stasimetric Objects, which will intersect the Line of Direction from the first Place of Anchorage in the corrected Distance of the second Point along the Coast; correct also the Figure of the Coast, and pencil it; and lastly insert remarkable Objects on the Land, and Diversities of the Coast between the Points, together with the Rocks observed, and Soundings taken by the Way. Take the Bearing of the third Point, mark its estmated Distance, sketch the Curve of the Coast, and correct both when you are at that Point as before; and go on in this Manner, from Point to Point, till the whole is finished.

Though this Method of Surveying is less perfect than any of the other Methods, yet if the Card of the Compass can be made to stand at Rest in the Boat, either by Art, or the Smoothness of the Sea, a Draught of a Coast may be made in this Manner, or of Bays or Harbours in it, which

will give a distinct Notion of them, and prove a pretty good Direction to Shipping, either along the Coast, or into these Bays or Harbours.

CASE 2. If the Coast extends eastward and westward, then two stasimetric Objects must be chosen, which lie about N E or S W from each other, and also their Bearing and Latitudes found, in order to find their Distance, and form a stasimetric Scheme for carrying on the Survey in the same Manner as before.

#### E X A M P L E IV.

*From a Vessel at Anchor in the Mouth of a Harbour, how to sketch the Harbour speedily, as far as can be seen.*

Suppose Fig. 6. in Plate iii. is the Harbour, and the Vessel at V: take the Bearing of any remarkable Part D, right up the Harbour; guess as near as you can to its Distance; draw out the Bearing on Paper, and by a convenient Scale mark on it that judged Distance. In the same Manner, mark a Point for the Bearing and judged Distance of C: then draw between C and D on the Paper, the Curvature, or Windings of the Coast, as they appear to the Eye. Next take the Bearings and judged Distance of the Point B, and draw the Irregularities of the Coast from B to C, and to F. Do the other Side of the Harbour from A to D and to G in the same Manner, and the Figure on the Paper will have an apparent Resemblance to the Harbour, exhibit its Dimensions something near the Truth, and shew the true Course into it, if there are no unperceived Rocks or Shoals in the Way. Next, insert in the Draught the most remarkable Appearances of the Coast, the Houses, Hills, Trees, Forts, &c. and such Rocks or Shoals as can be seen while you remain there. In sailing to or from the Harbour, the Lay of the Coast, on one, or both Sides of the Entry; and within the Harbour from B to C, and from A to E, may be corrected by the Bearings.

This Method of sketching a Bay, or Harbour, though far from being exact, will give such an Idea of it as may be of Service on some Occasions; and therefore should not be neglected when there is no Opportunity of doing it more exactly.

E X A M-

## EXAMPLE V.

*How to make a serviceable Draught of a Harbour by a Compass, or magnetic Needle, only.*

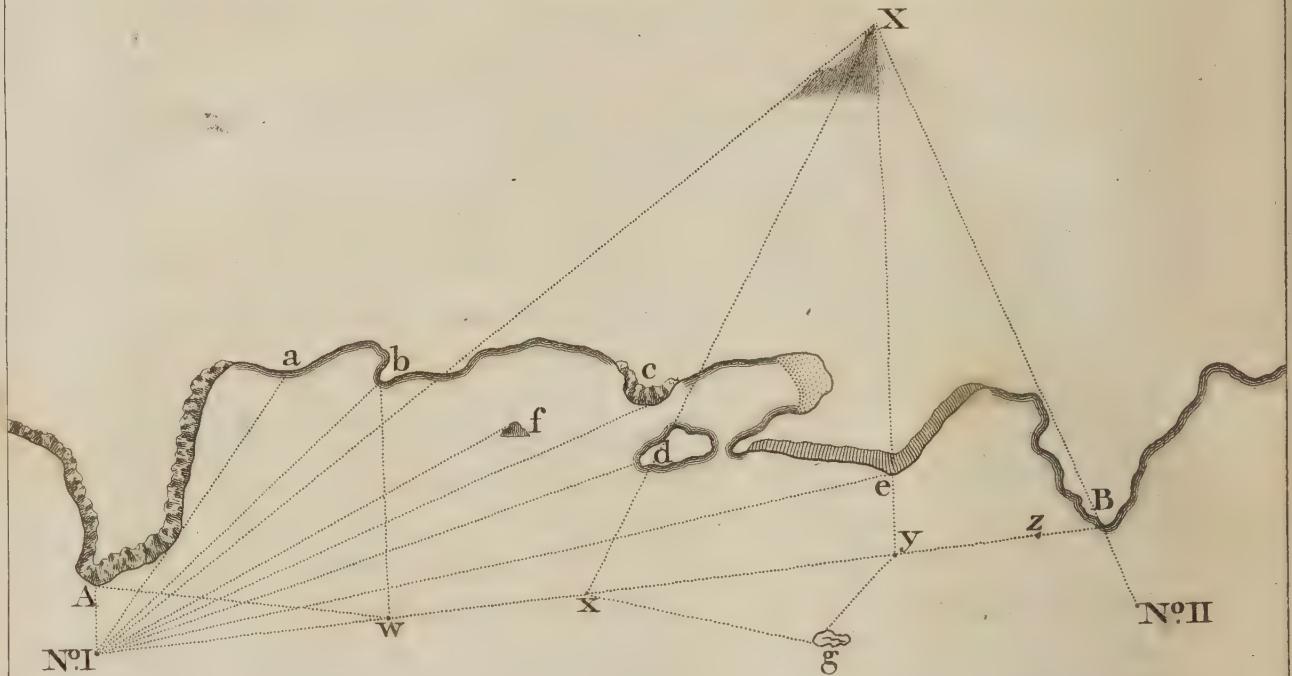
## PLATE II. FIG. I.

Go to some Eminence where the whole, or a great Part, of the Harbour may be seen, and select two conspicuous sharp Objects, as tapering Hills, Steeples, Chimney-heads, &c. as A and C, at a competent Distance from each other, and conveniently situated for a base Line; (if no such sharp Objects are to be found, put up Signals on the Hills, or Eminences;) take their Bearing in Degrees, and Parts of a Degree by the Needle, or Compass, which should be provided with Sights, and graduated for that Purpose; and estimate by the Eye the Distance of C from A, as near as you can judge. At A and C take the Bearing of B, and of any other remarkable Object D, on the other Side of the Harbour; and from these Data protract a stasimetric Scheme on Paper, wherein the Points A, B, C, D, will have the same magnetic Directions and proportional Distances as the Object they represent.

With this Scheme go to any adjacent Point of Land in the Harbour, as X; there set down your Compass, or Needle; take the magnetic Directions of two of the stasimetric Objects that are most advantageously situated; draw them out from their respective Points on the Scheme, and their Intersection will be the Point of Station X. At X take the Bearing of the next Point Y; guess its Distance from you; take that Distance from the Scale; mark it on the Bearing-line, and trace on the Paper with the Point of your Compasses the Curve of the Coast between X and Y; take also the Bearing of a and b, and of such Rocks and Shoals as may be seen along the Coast, and draw them on the Paper. Then go to Y; there take the Bearing of A, or of any other of the stasimetric Objects that may make a more direct Intersection; draw out its Direction, and its Intersection with X Y will be the corrected Place of Y. By the Bearing of b and of a, correct likewise their Positions; then the Curve of the Coast between X and Y may be traced with Black-lead more exactly, and remarkable Houses, Trees, &c. inserted, and



A Coast to illustrate Example VI. Page 89.



From A to B is the Extent of Sketch 1<sup>st</sup> X, an inland Hill determined by its Bearing at N<sup>o</sup>.I, and N<sup>o</sup>.II, or more advantageously at y.

N<sup>o</sup>.I, is the first Station, or Anchorage, of the Ship; its Distance from A is by Estimation.

N<sup>o</sup>.II, the second Station, or Anchorage; its Distance from B Estimated, or determined by two Bearings of Objects advantageously situated.

IA, Ia, Ib, &c. the Bearings of the several Points, Rocks &c. whose Distances from N<sup>o</sup>.I, are all to be marked, at first, by Estimation, and the Coast faintly sketched between them.

IB, the Ships Course, or the Line in which she is supposed to sail toward the second Station, measured by the Log.

W, x, y, z, are Distances on that Line, measured by the Log; at each of which some intersecting Bearings are taken and drawn out while the Ship is under Sail, for correcting the former Estimated Distances of the Points, Rocks &c. and then sketching more exactly the Figure of the Coast between the several Places.

and the Diversities of the Coast between X and Y marked. Proceed in the same Manner from Point to Point till the whole Coast-line of the Harbour is delineated. If any more stasimetric Objects become necessary, let their Places be determined by magnetic Intersections drawn from two of the former Objects.

Next, by Intersections of the Needle taken on Land, or of the Compass on Sea, find the Places and Extent of the several Rocks, Shoals and Sand-banks in and near the Harbour, take Marks for avoiding them, and sound the Depth of the Water round them, and in other Parts of the Harbour; and mark the best Anchorage.

Such a Draught as this of a Bay or Harbour neither requires much Skill nor Time to execute; and though not exact in the Distances, exhibits a just Representation of the Place; and will be found sufficient for directing Vessels in and out without a Pilot, if the Channel is not *very* intricate.

#### E X A M P L E VI.

*How to sketch a Coast in sailing along it.*

1. Prepare a Sheet of large Paper, with parallel Black-lead Lines and Perpendiculars crossing them all over, as in a stasimetric Scheme. Prepare also, in a Book, an *Observation Table*, in which is to be wrote distinctly and regularly, the several celestial Observations, Bearings, Distances measured by the Log-line, Rocks, Shoals, Soundings, Overfalls, Races of Tide, and other Remarks that may be made along the Coast. The Table may consist of seven or eight Columns, disposed in the following Order.

N

OBSERVA-

## OBSERVATIONS IN NAVIGATING THE COAST OF ——

## SKETCH I.

From Cape —— to the Point ——, being —— Miles measured by the Log. The Course from Station I. to II. being S $\frac{1}{4}$  W.

Year, Mon. and Day.	Mer. Alt. and of the Sun.	Bearings at Station I.	Time and Distances sailed from St. I.	Bearings and Soundings taken at these Dis- tances.	Bearing of Rocks, Shoals, Overfalls, and their esti- mated Distances, when in a Line with Points or Heads on the Coast.	Remarks on the Tides, and on the Nature and Di- mensions of the Rocks, Shoals and Anchorage.
D. M.	H. M.	Mi.	Fath.	Points and Heads	Mi.	
1774 Jan.	A W. 17° N. a W. 22 S.	1 20   $\frac{3}{4}$   11 45   4	A N. 5° W. b W. 20 S.	{ 21 c, and Rock. W. 7° N. d, and Hill. S. 19° W.	2 $\frac{1}{2}$	This Rock dries at low Water, and seemed 100 Yards in Length from N E. to S W. a Leading-mark to it is,

2. Let four expert Persons be appointed; one to take the Bearings exactly with an Amplitude Compass; one to oversee the Running-out of the Log-line, and to keep an Account of the Ship's Way, so as to be able readily to tell the Distance she has run, when required; the third to attend the heaving of the Lead, to write down the Soundings, and the Bearings of one or two Heads, Points, or remarkable Parts of the Coast taken at each Depth: the fourth, a Draughtsman, to draw out the necessary Bearings and Distances, and to delineate the Figure and Windings of the Coast at the several Stations, and to correct their Forms and Dimensions while the Ship is sailing along the Land.

3. Begin to take the Sketch, if it can be done, off some remarkable Promontory, or Point, that extends farther out than the rest of the Coast, and may intercept and turn off the principal Stream of Tide, so that there may be

be little Stream where the Ship's Way is to be first measured by the Log. Bring the Ship as near the Point as can be done with Safety; there drop Anchor, if the Depth of the Water and the Ground are proper for it; if they are not, make the Ship *lie to* as steadily as possible.

4. On a convenient Part of the prepared Paper make a *Dot* to represent the Place of the Ship; draw a small Black-lead Circle round the Dot to make it more conspicuous, and mark it with N° I. for the first Station. Then take the Bearing of the Point, or Promontory you lie off; draw it out with a Protractor from the Dot; estimate its Distance by the Eye as exactly as you can; take that Distance from a proper Scale of equal Parts, mark it on the Bearing-line with a Dot, and with the capital Letter A. Next take the Bearing of the farthest Point, or Promontory, which is seen distinctly, and is intended for the Limit of the first Sketch; draw it out from N° I; estimate its Distance by the Eye; take it from the Scale, and mark it on the Bearing-line with a Dot and with the capital Letter B. Take also the Bearings, and estimate the Distances, of all the Points, Heads, Inlets of Rivers or Harbours, Rocks, Corners or Heads of Bays, that are seen between the two Points A and B; draw them out and mark them with the small Letters a, b, c, &c. as far as they are seen with any Distinctness.

5. Sketch faintly (so as it may be easily rubbed out when it is to be corrected afterward) the Figure of the Coast between the several Dots; as from A to a, from a to b, from b to c, &c. as far as two or three Miles; or only as far toward B as you have a distinct Sight of the Bays, Irregularities, and Dimensions of the several Parts of the Coast, and let the rest be sketched as you sail past them, after their Distances have been corrected by the Ship's Run measured by the Log.

6. Look for some conspicuous tapering Hill, Building, Steeple, or Grove of Trees, up the Country, and several Miles distant from the Shore, and situated as far along the Coast toward B, or beyond it, as can be found; take the Bearing of that Hill, or Object exactly, draw it on the Paper; and when another Bearing of it is taken at a proper known Distance from the Station I, its Position will then be determined on the Draught by the Intersection of these two Bearings, and will serve for correcting, or determining the Distances of Points, Promontories, Bays, &c. along the Coast.

7. While the Ship lies at Station I. off A, let the Latitude be observed, if there is an Opportunity, either by the Sun or Stars. Then get the Ship under Way, and sail in a straight Line for the second Station off B; measure the Ship's *Run* carefully by the Log, and keep an Account of the Time by a Watch as you sail; that when you are off any of the Heads, Points or Openings marked with the small alphabetical Letters, their Distance may be known more exactly and corrected: Correct also, by the Way, the Figures and Dimensions of those Parts of the Coast which were imperfectly seen and sketched at N° I. Also, sound the Depth of the Water by the Way, and at each Sounding take the Bearing of the first Point A, or of the last Point B, and of one or more of the Heads, or Points on the Coast; and likewise of the inland Hill, or other Object, in order to determine afterward the Places of the several Soundings, and to discover whether the Ship is carried out of the straight Course by Streams of Tide, Currents, or Lee-way, and to correct that Deviation. Let such Rocks, Shoals, Overfalls, or remarkable Streams of Tide as are met with, have two or three Bearings taken of them when they appear in a Line with any of the Points, Heads, or Islands in the Draught, and their Extent estimated by the Eye; and all wrote in a Book, to be inserted in the Draught more correctly afterward; and if there is an Opportunity, to be more particularly examined another Time.

8. When the Ship has got to the Head, or Point B, the Limit of the first Sketch, drop Anchor; or bring the Vessel to, as at N° I; write at the Top of the *Observation Table* the Distance between A and B, as measured by the Log; and write in the Table all the intermediate Distances so measured; and likewise the Observations, Bearings, Times of them, Distances, Soundings, and Remarks; all of which must be reserved for future Use, when the Sketch undergoes another Correction.

9. At Station II. take the Bearing of the inland Hill, or Object, one Direction of which was drawn out from the first Station, and these two Bearings will intersect in the Place of the Hill; which will then serve to correct the Position of some Parts between the first and second Stations, and to determine by Intersections the Position of Heads, Points, and other Parts of the Coast between the second and third Station; and verify, or correct, their Positions measured by the Log.

Take

I

Take also the Bearing of the Top of some other conspicuous inland Hill, or Object, situated farther along the Coast, the Position of which must be determined afterward by another Bearing taken at a proper Distance; and *that* Hill will then serve for the same Use between Station III. and IV. as the other Hill does between Station II. and III. Then take the Bearing of the Point, or Head, B, next you; estimate its Distance, and mark it on the Paper in the Bearing-line with a Dot, and N° II. for the second Station, or Beginning of the second Sketch; also the Bearing of a distant Point or Promontory, for the other End of it; estimate its Distance by the Eye, draw it out from N° II, and mark it with a Dot and the capital Letter C. Then take Bearings of all the Heads, Points, Inlets, &c. between B and C, estimate their Distances, and mark them on the Bearings, and sketch the Figure of the Coast between the Points, and proceed as was directed at Station I. till the whole is finished in the same Manner. This however is but a rough imperfect Sketch, and must be corrected at more Leisure, by drawing the whole anew on clean Paper from the Distances and Observations in the Observation-Table.

10. If the Coast lies northward and southward; observe the Latitudes of the Heads, Points, or Inlets, at every 20 or 30 Miles Distance, if it can be done; and by these let the principal Distances be a second Time corrected. If the Coast lies nearly East and West, the Distance cannot be corrected by Latitudes, and therefore frequent Observations of the Sun are unnecessary.

11. When you are obliged to leave off the Work for want of Day-light, or by bad Weather, do it as near some Head, or remarkable Part of the Coast as you can; take Bearings, or Marks, on the Place you left off at, that it may be easily found when you want to begin again.

12. If the Coast navigated is a *Strait*, or the Mouth of a River, where the Land on each Side can be seen; at each Station on one Side, take the Bearings of remarkable Heads, Points, Islands, Hills, &c. on the opposite Side, and let that be *one* Direction by which their Positions and Distances are determined.

A Coast sketched in the above Manner may contain many good Harbours or Bays, which are not to be perceived in sailing along it. If there is an Opportunity,

Opportunity, the Ship should go into some of them, until a Delineation was made of them, and the Entrance sounded in and out; which would render the Sketch of the Coast of much more Service:

### C H A P. IV.

Rocks and Shoals determined and described. Tides and Soundings.

Tides described. Harbours described. Copying and reducing Draughts. Necessary Instruments.

#### S E C T I O N I.

*How to determine the Distance and Extent of Rocks, Sand-banks, and Shoals, lying off a Coast that has been surveyed; and how to avoid them by the approximating Angle.*

**T**O determine their Distance; anchor a Vessel at one End of the Rock, Sand-bank, or Shoal; and in the Vessel, with *Hadley's Quadrant*, take the Angles between three Objects on Shore, whose Distances from one another are known, and (by longim. Prob. Page 18.) find the Distance of the Vessel from these Objects, which will give that End of the Rock or Shoal.

Or, anchor a Boat at one End of the Rock, or Shoal, and take the Angles on Shore from two Stations whose Distance is known, and (by longim. Prob. Page 10.) find the Distance of the Boat.

To find the Extent of the Rock, Sand-bank or Shoal; if its Extent is small, it may be estimated by the Eye sufficiently near: if large, the other Extremity must be determined in the same Manner, by the Angles taken on Shore; or rather, by one Angle taken on Shore, and that Direction intersected by a magnetic Bearing at the other End before determined.

When a Coast has been first carefully surveyed, the Distance of Rocks, Shoals, and Sand-banks may be found with sufficient Exactness by two Bearings of some distinct Part of them taken on Shore with a good magnetic

Needle ; provided the Stations are not at such a Distance, or Position, from each other as to make too oblique an Intersection.

If the Shoal lies so far from the Coast eastward or westward that the Land cannot be seen at it ; in moderate Weather, anchor a Vessel at the Shoal, and two Boats with their Sails loose at a proper Distance and Position from each other, between the Shoal and Land, and so as the Boats may be seen from the Shoal, and the Land seen from the Boats. Let an Observer in each Boat take the Angles between three known Objects on Land with *Hadley's Quadrant* ; and another Observer in the Vessel at the Shoal, take the Bearing of the two Boats as exactly as possible with a Compass. By the first Angles the Distance of the Boats from the Land, and from each other, will be found ; and by the Bearings, the Distance of the Shoal from the Boats, and of Consequence from the Land, will likewise be determined.

The Distance of the Shoal from any one of the Boats may likewise be measured by the Velocity of Sound, by appointing two or three Guns to be fired from the Vessel when a Signal is made, as directed in Long. Oper. 3.

If Boats, or small Vessels, cannot anchor within Sight of the Land, and, at the same Time, be seen at the Shoal ; then sail from the Shoal towards the Land with a leading Wind, measure the Ship's Run carefully with a Log-line, and when the Land is seen distinctly, there drop anchor ; and find your Distance from it by *Hadley's Quadrant* ; and the Distance run from the Shoal, added to the Distance of the Land (and reduced to one Direction, if necessary) will give the Distance and Position of the Shoal from the Land nearly. If the Shoal lies North or South from the Coast, its Distance may be found by the Latitude, and the Ship's Course in sailing towards the Land.

At every Rock, Shoal and Sand-bank, Marks ought to be taken on the Land, for avoiding them. In doing this, chuse such Objects only as are remarkable and striking, which a stranger to the Coast may readily distinguish ; and not such as are small, or faint, which none but a Person particularly acquainted in the Place can find out. If no such distinct Land-marks can be found for directing Vessels past them ; then they may be avoided by the *Approximating Angle*, taken in the following Manner. On the Out-side of such Rock or Shoal, with *Hadley's Sextant* observe carefully the Angle which two remarkable

remarkable Objects on the Land, at a proper Distance and Situation, subtend; take also their Bearings by the Compass; and write all down for future Use. And while at any other Time, in sailing towards that Shoal, the Angle between these two Objects is found *less* than the approximating Angle, you are certainly without the Shoal; and also without all Rocks or Shoals lying within a Circle that may be supposed to pass through the two Objects on Land, and the Point where the approximating Angle was taken. For all Angles *in* the Periphery of that Circle, which the two Objects subtend, are equal; the Angles *within* the Periphery, greater than these; and the Angles *without* it, less. Euc. iii. 21. The Bearings will shew when you are a moderate Distance from it on either Side.

*To find out Shoals that may be along any Coast:* in blowing Weather, go to the Top of a Hill, Cliff, or Eminence, and observe where the Sea appears rougher, or the Waves break more, than in other Parts, for there the Water is shallowest; except the Roughness arises from a Stream of Tide going against the Wind; or two Streams meeting each other. Take Land-marks, or the Bearing of the Breakers, from one, or two, remarkable Places, and that will direct you to the Shoal in a Boat afterwards, when the Weather is moderate.

*Or, in a Calm, observe just after Slack-water, where a grey Stream, or Ripple on the Surface, first begins; and at the smooth Edge of that Ripple the shallowest Water will be found.*

Breakers are often seen in calm Weather, before a Gale of Wind, where the Water is shallower than in the neighbouring Parts; though there may be a sufficient Depth where the Breakers rise.

## S E C T I O N II.

### *Circumstances to be taken Notice of in describing Rocks, Shoals and Sand-banks.*

It may not be unnecessary first to caution a Surveyor against taking his Accounts of Rocks, Shoals, Tides, or Harbours, solely from the Information of others, without examining them himself: for there are few to be met with who have been careful and circumstantial enough in their Observations; and

and some who will affirm for Fact, as if grounded on actual Observation, what is only a vague Opinion long entertained on uncertain Authority. Even professed Pilots are often satisfied with knowing how to avoid Dangers by keeping at a sufficient Distance from them, or in the common Channel, without troubling themselves about farther Particulars. Information should be made Use of, but only as a Direction for farther Inquiry, and to facilitate your own Search and Examination. In describing Rocks, Banks, and Shoals,

1. Take Notice how far they lie from some remarkable Head, or Part, of the nearest Land ; and how they bear from that and from some other conspicuous Part of the Coast.
2. How far they extend in Length, and in what Direction.
3. What Land-marks lead directly on them.
4. At what Time of the Tide they begin to dry, or are covered.
5. The least Water on the Shoal.
6. How they are to be avoided on each Side.
7. The approximating Angle, if distinct Land-marks are not to be found.

### S E C T I O N . III.

#### *Affections of Tides relative to the Soundings.*

As the perpendicular Rise of spring and neap Tides varies sensibly according to the Distance and Position of the Moon with Respect to the Earth; and of the Earth and Moon with Respect to the Sun, and also by Winds and Weather; the Depths of the Water proper to be inserted in Draughts of the Sea-coast, seem to be such only as are taken at low Water in ordinary Spring-tides ; and not such as are found when the Tides are affected by extraordinary Causes. Ordinary Spring-tides may be reckoned, such as happen in moderate Weather, on the third Day after new or full Moon, when she is at her mean Distance from the Earth, and the Sun not near the Equinoxes.

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It would be extremely tedious to wait for low Water Spring-tide, to take all the necessary Soundings. Therefore, that the Depths proper to be marked may be known, whatever Time they are taken, let the perpendicular Rise, or Fall, of the Water, in moderate Weather, on the third, or fourth Day after Full or Change, and after Quarter Moon, be first found by Experiment; and then the Depth at low Water ordinary Spring-tide may be known sufficiently near, by making the following Allowances at the Time each Depth is taken.

If the Depth is taken at High-water Spring-tide, then deduct the full Rise of ordinary Spring-tide from the Depth found, and the Remainder will be the Depth to be inserted in the Draught.

Spring-tide, at the

1st Hour before and after High-water, deduct $\frac{11}{12}$	of the full Rise of ordinary Spring-tide.
2d Hour before and after High-water, deduct $\frac{3}{4}$	
3d Hour before and after High-water, deduct $\frac{1}{2}$	
4th Hour before and after High-water, deduct $\frac{1}{4}$	
5th Hour before and after High-water, deduct $\frac{1}{2}$	
6th Hour, or about Low-water, deduct 0	

If the Depth is taken at High-water Neap-tide, from each Depth deduct  $\frac{4}{5}$ ths of the full Rise of Spring-tide, or the whole Rise except one fifth Part.

Neap-tide, at the

1st Hour before and after High-water, deduct $\frac{4}{5}$	of the full Rise of Spring-tide.
2d Hour before and after High-water, deduct $\frac{3}{4}$	
3d Hour before and after High-water, deduct $\frac{1}{2}$	
4th Hour before and after High-water, deduct $\frac{1}{4}$	
5th Hour before and after High-water, deduct $\frac{1}{5}$	
6th Hour, or about Low-water, deduct 0	

These Allowances are not given as quite precise, for they will vary a little from several Causes; but as a general Rule, which will be found sufficiently exact on most Occasions. Shoals and Channels that have less than four Fathoms, and are much in the Way of Shipping, require to be sounded at low Water Spring-tide, that the Depth of the Water there may be more precise.

That the Depths to be marked may be known with more Exactness, let it be observed, that when the Moon is in the Perigea, or least Distance from the Earth, then the perpendicular Rise of spring and neap Tides are about

one eighth more than ordinary; and that about the Equinoxes, the full Rise is about one tenth more, from that Cause also: that when the Moon is in the Apogea, or at her greatest Distance from the Earth, then the Tides are proportionally less than ordinary: so that sometimes the greatest neap Tides may rise as high as the least spring Tides. But these and other Affections of the Tides have not yet been so minutely inquired into as they deserve. Hard Gales of Wind in any Place, especially if it blows in the Direction of the Flood, swells the Tide to an uncommon Height: Storms at a Distance, are found sometimes to occasion an extraordinary Rise of the Water in Places where the Gale is not felt: and two Streams of Tide running round an Island, make the Tide rise higher near their Meeting than it does in other neighbouring Parts.

Where the Depth at low Water is above five Fathoms, there a Foot or two, more or less, is of no great Consequence; and therefore it is sufficient if the Deduction from the Rise of spring Tide is in Proportion to the Time of the Tide: that is, at high Water, deduct the full Rise; at quarter Ebb, three Quarters; at half Ebb, or Flood, half the Rise; and at three Quarters Ebb, or one Quarter Flood, deduct one Quarter of the full Rise from each Depth found.

The best Way to find the exact Time of high Water on the full and change Days is, in moderate Weather, when the Sea is smooth; then, about three Hours before high Water, mark the Place on the Shore to which the Tide reaches, and write down the Hour and Minute you observed it: on the Ebb, watch till the Tide falls to the same Mark on the Shore, and see what Hour and Minute it is then; and to the Time between the two Observations add two Minutes for each Hour, on Account of the Acceleration of the Tide, and half that Sum, added to the Time of the first Observation, will be the true Time of high Water that Day.

#### S E C T I O N IV.

##### *Circumstances to be taken Notice of in describing Tides.*

1. Observe the Time of high water on the full or change Day of the Moon. Let it be expressed in Hours and Minutes, not as is commonly

O 2 done

done, by the Bearing of the Moon, which serves no useful Purpose whatever, but always occasions an unnecessary Reduction of Time, and Points of the Compas, into one another.

2. Observe the perpendicular Rise of the Tide on the third, or fourth Day, after new or full Moon; and at different Seasons of the Year; and when the Moon is in the Perigea and Apogea.

3. Observe the Direction of the Streams of Flood and Ebb, near the Coast, and at a Distance from it; the Irregularities, or Variations of their Direction; and at what Time of the Tide on the Shore, the Stream in the Offing begins to turn.

4. Observe the greatest Velocity of the Stream at spring and neap Tides, off the Coast, and in Channels; measured by a Log-line heaved from a Boat or Ship at Anchor; or by driving with the Stream, in a Calm, a known Distance.

5. In speaking of irregular Tides, do not say, it flows 9, or 8 Hours, and ebbs 3, or 4 (as is often done) when you mean only, that the Stream runs 9, or 8 Hours one Way, and 3, or 4 the other; for that creates Confusion, and sometimes Mistakes. While the Tide rises, and then only, it is Flood; and only while it falls, is it Ebb; in whatever Direction these Streams run. Therefore always express such an Irregularity in the Stream, by the Time it *runs* in one Direction, and the Time in the other.

## S E C T I O N V.

### *Circumstances to be taken Notice of in describing Harbours.*

1. Observe if they are well sheltered, or, on what Quarter they are exposed.
2. The Depth of the Water, and Nature of the Ground in the Anchorage.
3. What Number of Ships they are capable of containing easily; and of what Size, or Draught of Water.
4. In what Part to anchor; and the Land-marks, or Bearings for it.
5. How

5. How the Entry may be distinguished at a Distance, or how to fall in with it coming from Sea.

6. How to avoid Rocks and Shoals in the Way to a Harbour; or sailing to the Anchorage.

7. If there is a Sufficiency of fresh Water, and where it is; and what Provisions the Place affords.

8. In unknown Places, observe what Forts, or Batteries are in them; what their Strength is; and how near a large Ship may get to them: or, in what Parts of the Harbour, Forts, or Batteries, may be advantageously erected, either for Defence, or Attack.

#### S E C T I O N . VI.

##### *How to copy a Draught exactly.*

Pin clean Paper over the Draught to be copied, so as they may not shift: hang, or hold, them up before a Window, or the Glass of a Chariot, and the Drawing will appear through the clean Paper, and may be traced on it with a black-lead Pencil, and afterwards with Ink.

Or, rub the Back of the Draught, or as much of it as may be necessary, over with Charcoal; lay the rubbed Side on clean Paper, and pin them together, and lay both on a Table, the Draught uppermost, and with a smooth-pointed Tracer, of Brass, Ivory or Steel, not very sharp, trace exactly all the Coast-line first, and then other Parts, and the Under-paper will retain the Impression distinctly. If the Charcoal stains other Parts of the Under-paper, it will rub out by the Crumb of Loaf-bread, after the Tracing is inked.

If, in the Copy, you want the true Meridian right up and down on the Paper, instead of the Magnetic: draw the true Meridian first on the Middle of the original Draught; then in a proper Part of the clean Paper (before it is pinned to the other) draw a Line right up and down; and pin them together so, that the Line on the clean Sheet may lie exactly under, or over.

over (according to the Manner of copying) the true Meridian drawn on the Original ; then trace the one from the other.

## SECTION VII.

*How to reduce a Draught to a smaller Scale.*

A Pantograph is by far the easiest and most expeditious Instrument for diminishing Draughts with Accuracy : but, when one of a sufficient Size for large Draughts cannot be got, either of the following Methods will serve.

With Black-lead draw the large Draught all over with Cross-lines, forming exact Squares : draw the clean Paper for the Copy also over with the same Number of Squares ; but their Sides smaller, in Proportion to the intended Size of the Scale ; such as one half, one fourth, &c. of the Length of the other : distinguish, by a stronger Line, and mark with a Figure, every fifth or sixth Row of Squares in both, so that the several corresponding Squares may be more readily perceived : then, in each of the small Squares, draw, by the Eye, a Curve similiar to that which is in the corresponding larger Square, till the whole is copied.

Or ; draw a Black-lead Line from End to End of the large Draught, in any Direction, as A B (Plate III. Fig 7.) by its proper Scale divide that Line into any Number of Miles, or half Miles ; at each Division raise Perpendiculars, and draw them out to the Coast-line, and mark them 1, 2, 3, &c. at each End.

On a proper Part of the Paper for the small Draught, draw a Black-lead Line right up and down on the Paper, for the Meridian ; and another Line making the same Angle with it, as the divided Line in the large Draught does with its Meridian, and in the same Position as a b (Fig. 8.) make a b the half, the fourth, &c. of A B, according to the intended Diminution of the Scale ; divide it into the same Number of Miles, or half Miles, by its proper Scale, as the other is ; at each Division raise Perpendiculars, draw them out a sufficient Length, and mark them at each End, 1, 2, 3, &c. as in the other, that the corresponding Perpendiculars may be readily found.

Measure

Measure on its proper Scale the Length of the first Perpendicular on A B, in Miles and Parts of a Mile: take the same Number of Miles and Parts from the small Scale, and set it from 1 to 1, in the small Figure, and mark it with a Dot: do the same in every other Perpendicular, and so many Points of the Coast in the small Figure will be found.

Then, between every two extreme Dots in the small Figure, sketch by the Eye the corresponding Curve of the large Coast, and both Figures will be similar in every Respect. Hills, Rocks and other Objects may be drawn in the same Manner.

There are proportional Compasses and Compasses with three Legs made, by which small Draughts may be diminished; but they do not answer so conveniently for very large Draughts.

#### S E C T I O N. VIII.

##### *Instruments and other Necessaries for taking a Survey of the Sea-coast.*

A good astronomical Quadrant 14 Inches Radius of Bird's Construction; with a Pirn of fine Silver-wire for the Plummet.

A Hadley's Octant, 18 Inches Radius.

A Hadley's Sextant of Brass, 9 Inches Radius.

A Bell-metal Theodolite, with a spirit Level, Needle and two Telescopes: the Objects to be seen direct through the Telescope, not inverted.

A Case of good Pocket-instruments.

A Brass-scale 3 Feet long, with various Lines of equal Parts on it.

A Wooden-ruler 5 Feet long.

A beam-compass of Wood two Feet long and  $\frac{1}{4}$  broad, divided from the Center Point along a chamfered Edge into Inches and eights of an Inch.

A Gunter's Scale 2 Feet long; and one a Foot long.

A Brass

A Brass circular Protractor 12 Inches Diameter, with an Index and Pricker to mark the Degrees and Minutes on the Paper; the graduated Edge silvered. Another 5 Inches Diameter; and another semicircular, 3 Inches Diameter.

A measuring Pole 30 Feet long, divided into Feet; and consisting of two Pieces that may be joined together End-ways, and taken asunder at Pleasure.

A straight Mahogany-pole, precisely 6 Feet long, divided into Feet and Inches; the two extreme Feet into Decimals of an Inch.

Two Iron-chains, each 60 Feet long; each Link a Foot, and the Thickness of a Quill; with brass Marks at every tenth Link.

Four Iron-Reels, with Stakes and Lines, such as Gardeners use.

Three wooden Poles, each 10 Feet long, and  $1\frac{2}{3}$  Inches Diameter; 12 Poles, each 6 Feet long and one Inch Diameter: all pointed with Iron on a Socket, so as to pierce the Ground easily and stand firm in it.

Thirty cylindrical Iron-pins, sharpened at one End, for piercing the Ground at the End of each Chain, or Measuring-pole; each about 8 Inches long and half an Inch Diameter.

Two Canvas-pockets for holding these Pins, with Strings for tying them round the Waste of the Measurers.

A good amplitude Compas.

A magnetic Needle, with an agat Cape, Brass graduated Circle, and a Lever, in a square wooden Box, with jointed Sights to lie within the Box when shut.

A Pocket-compas, with an agat Cape.

A Ship's-compas, for taking Bearings in the Boat; the Depth of the Box equal to the Diameter of the Card, so as not to touch the Glass above, nor the Bottom below with any Motion; and the Angle of the Cape a Right-angle.

A good portable refracting Telescope, of Doland's Construction.

A Case

A Case of Leather, divided Length-ways by Pasteboard into two Apartments, 2 Feet 1 Inch long, and 18 Inches broad; for holding the original Draught, spare Paper, Scales and Protractors, and securing them from Rain and Salt-water: also a small Drawing-broad, the Size of the Case, with Leather Straps and Buckles, for carrying both together over a Man's Shoulder.

A Draught-cover, 2 Feet long, and 18 Inches broad, made of Pasteboard covered with Oil-cloth, like the Cover of a Book, for laying the original Draught in, that it may slip easily in and out of the Leather Case without Injury.

A good Stop-watch, that shews Seconds.

A Two-foot reflecting Telescope, for observing Eclipses of Jupiter's Satellites, or of the Moon, when necessary to find the Longitude.

A large Drawing-table, 7 Feet long, and 5 Feet broad, with a round Moulding along the Edge of one Side and one End, to keep the Paper smooth when drawn over the Table to reach the Top of the Draught. The Board, when not in Use, to come asunder in two Leaves, one of the Leaves  $1\frac{1}{2}$  Inch broader than the other: when in Use, to be fastened together on the under Side by two Hooks, one near each End; and to be supported by a Frame, or by a long three-legged Stool at each End; with a loose Post, or Pillar, under that Side where the Drawer commonly sits, that the Board may not warp, or bend, by his Weight when leaning on it: the Parts of the Frame, or Stools, to be made and marked so as to be easily taken to Pieces, or joined, occasionally.

A Folio Observation-book of 4 Quires medium Paper, ruled at the Top for a running Title; and for a local Title; a two-inch Margin on the left Side of each Page; and two Columns on the Right: the first half an Inch, for Degrees and Minutes, the other one Inch, for magnetic Bearings.

A Folio Journal-book, of 3 Quires medium Paper, ruled at the Top; with a two-inch Margin, and a small Column for the Days of the Month on the left Side of each Page.

A View-book, of half a Quire imperial Paper, opening Length-ways to the Extent of two Sheets, and twelve Inches broad; the Leaves pasted and guarded in the Middle.

Memorandum-books, each of 6 Sheets Demy-paper, with red Leather Covers.

East India, or such other large Paper, for the original Draughts; that may bear to be often folded without cutting.

Imperial Paper for the Draughts.

Transparent Paper, for copying Parts of a Draught readily for occasional Service.

Writing Paper.

Cartridge Paper, for covering up a clean Draught while it is making out.

Quills, Crow-quills; Cake-ink, Indian-ink; Hair-pencils, some of the smallest, and some of a middling Size; Black-lead Pencils, some of which flat; Charcoal, for copying a Draught by tracing it on a Table; Penknives.

A Survey-boat, or six-oar'd Cutter, 24 Feet from Stem to Stern, 7 Feet 3 Inches broad, and 2 Feet 11 Inches deep, with Wash-boards, Masts, Sails, &c. and two Graplings, a larger and a less; the larger to have a Block made fast to the Ring for running 40 or 50 Fathoms of Rope through, in order to moor the Cutter off the Shore or Rocks, and haul her in or out at Pleasure. Also an Awning of painted Canvas to sleep under where it is not safe landing. Such a Boat will carry, of King's Allowance, one Week's Beer, and two Weeks Provisions (Water and Firing excepted) of all other Kinds for eight Persons; together with a Tent, Bed-cloths and cooking Vessels; and is fit for rowing as well as sailing, if made light.

A Captain's Tent and Marki, with a Bed and Bedding to serve two Persons.

A Vessel about 120 Tons Burden, pretty broad in the Beam, and full in the Bows; and of such a Mould as to draw little Water, and to take the Ground

easily; with a Cabbin 12 Feet in the Floor, well lighted, and fit to hold the above-mentioned Drawing-table. Her Complement of Men (besides the Surveyor, his Assistant and two Servants) to be a Master, a Purser, a Mate, a Midshipman, a Carpenter, a Sail-maker, a Boatswain, a Cook, fourteen able Seamen before the Mast, and a Pilot for the Coast. A larger Vessel cannot go into small Creeks, which will often retard the Service, and be found very inconvenient on a Survey: a Vessel much less, cannot keep Sea well, nor carry Provisions and other Stores sufficient for the Season; which is likewise a Disadvantage. In hot sickly Climates, and Countries where Seamen cannot be got easily, a greater Number of Hands will be necessary, and likewise a Surgeon. If the Surveyor has two Assistants under his Direction, that both may be employed at the same Time, another six-oar'd Cutter, with six Hands more, will be necessary: and, if the Coast has many Shoals, Sand-banks, or Bar-harbours along it, or few Harbours that are accessible at all Times of the Tide, a Pilot for each Boat, who is minutely acquainted with the Coast, will also be necessary, to guard against Misfortunes, and for finding out the Shoals and Channels more readily.

## C H A P. V.

SEC. I. To find the Longitude by an Eclipse of one of Jupiter's Satellites. 2. To continue a Meridian through a Kingdom. 3. To trace a Parallel of Latitude through a Kingdom.

## SECTION I.

*How to observe an Eclipse of Jupiter's Satellites, in order to find the Longitude of the Place of Observation.*

THE most exact Survey of a Coast cannot be reckoned complete, till its Longitude from some other noted Place is known: and the most careful Surveys, if they extend far eastward, or westward, will require to be corrected by an Observation of the Longitude. It is therefore a necessary Part of a Surveyor's Business to understand how this Problem is to be performed.

For determining the Longitude of one Place from another on Land, Eclipses of Jupiter's Satellites, especially of the first, are found to be the most advantageous celestial Appearance. If the Hour, Minute and Second when any such Eclipse happens is observed in two Places, the Difference of these Times, converted into Degrees, Minutes and Seconds of a Circle, will give the Distance of their Meridians, or the Longitude of one Place from the other. Or, if the precise Time of an Immersion, or Emersion at one Place is known by Calculation, and found by Observation at any other, their Difference of Longitude will then be known. There are Tables published annually, in the nautical Almanack, of the Time at Greenwich, when Jupiter's Satellites will be eclipsed: if therefore the Time in another Place, when such Eclipse happens, especially of the first Satellite, is found by Observation, their Difference of Longitude will be then known.

To find the precise Time of the Day at which an Immersion, or Emersion of a Satellite is observed, four principal Operations are necessary. 1. *To find an exact Meridian-line*, so that by it a Clock, or Watch, may be set either to apparent or to mean Time at the Place of Observation. 2. *To regulate the Clock or Watch*, so as they may keep equal Time: or, to find the Error of their Going, that proper Allowance may be made for it. 3. *To set the Clock, or Watch, to mean or equal Time by the Meridian-line*. 4. *To observe, and note, the precise Moment when the Eclipse begins, or ends*.

1. *To find an exact Meridian Line.* For doing this, see several Ways by the Stars (Part 1. Beginning at Page 47.) which are more to be depended on than any Observations of the Sun, if the Observer is not provided with the best Instruments and a convenient Observatory.

2. *To regulate a Clock, or Watch, so that they may keep equal Time*; or to find how much they go fast or slow, that Allowance may be made for it. The most simple and certain Test of equal Time is, the diurnal Rotation of the Earth on its Axis: or, which results from it, the apparent daily Revolution of the fixt Stars round the Axis of the Earth. By that, therefore, let your Clock, or Watch, be regulated, as directed in Merid. Prob. Page 56.

3. When

3. When the Clock, or Watch, is regulated to go equally, or the Error of their Going is found; then let them be set to the mean Time \* of the Day by the Meridian-line, in the following Manner.

In a Room with a large Door, Window, or Opening towards the North, from whence your Meridian-line northward may be seen (if the Meridian has been marked toward the North) and another Door, Window, or Opening, towards the South, right opposite to the other, through which the Sun may be seen at Mid-day; hang up two Plumb-lines exactly in the Meridian, as far asunder as conveniently may be; let the Plummets, or Weights, be as heavy as the Lines will easily bear; and make them swing in Water, to prevent shaking by the Motion of the Air. Stop the Clock, or Watch, precisely so many Minutes and Seconds before, or after twelve, as are equal to the Sun's Equation that Day, which may be found in Equation Tables; and with a smoked or tinged Glass before your Eye, observe when the Center of the Sun comes exactly in a Line with the two Plumb-lines, and that Moment call out, Now; that your Assistant may set the Clock, or Watch, agoing immediately. Or, which perhaps is more accurate, stop the Clock as many Minutes and Seconds before the Time pointed out by the Equation, as the Center of the Sun takes to move the Length of his Semi-diameter that Day (which may be found in the nautical Almanack) and when the western Edge of the Sun just touches the Plumb-line, set the Clock, or Watch, agoing, and they will then be set to mean or true Time. In most Places it will be necessary to have a small temporary Room, for an Observatory, built on purpose; with an Opening toward the North, and another toward the South, that may be shut at Pleasure, to keep out any Draught of Air which would disturb the Plumb-lines.

\* Mean, or equal Time is, the Time shewn by a Clock that is set right, and goes exactly equal. Or, it is the Time shewn by the Sun, corrected by the Equation; or when Allowance is made for the Inequality of the Sun's Motion.

Apparent Time is, the Time shewn by the Sun (whether in the Meridian or not) without making any Allowance for the Inequality of the Sun's Motion: this is the Time shewn by a good Sun-dial; or found by an Azimuth or Amplitude of the Sun.

Equation of Time is, the Acceleration, or Retardation, of the Sun's Motion; or how much the Sun is too fast, or too slow, of equal Time.

4. To observe, and note, the precise Time of the Immersion, or Emerson, of a Satellite: Two or three Days preceding the Night of the Eclipse, set your Clock, or Watch, to mean Time by the Sun and the Meridian; and, if it can be done, examine its Accuracy, the Day of the Eclipse also; but without altering the Clock, though it has not gone quite exact; for Clocks and Watches often go irregular a little after they have been set back or forward: only, note how much it is less, or more than the mean Time each Day, that proper Allowance may be made in settling the true Time of the Eclipse. Guess as near as you can to your Longitude from Greenwich, and from thence find at what Time the Eclipse may be expected with you: if you are eastward of Greenwich, the nominal Time it will happen with you will be so much later than at Greenwich; if you are westward of Greenwich, the Time with you will be so much sooner, according to the Longitude. An Hour before it is expected, adjust the Telescope \* to your Sight, so as to see Jupiter's Belts distinctly. A Quarter of an Hour before the Eclipse is expected (or more, if your Longitude is very uncertain, or if it is any Satellite but the first) sit down to your Telescope; and if it is an Emerson, keep looking towards that Part of Jupiter's Disk on which you find, by the nautical Almanack, it is to happen. But observe, that if your Telescope inverts Objects, the Eclipse will happen on that Part of Jupiter's Disk which is diametrically opposite to that in the Almanack. While you are watching the Appearance of the Satellite, let your Assistant count every Second audibly, by the Beats of the Clock, or by the Motion of the Second-hand of the Watch; and the Instant the Satellite appears, if an Emerson; or the Instant it disappears, if an Immersion, call out, *Stop*: and immediately, the Assistant should repeat the last Second he counted, which the Observer should note down directly, while the Assistant looks to the Minute, and notes it down. Write down likewise what Sort of Telescope you observed with, and its length and magnifying Power: for Eclipses will be perceived somewhat sooner by a good Telescope than by one of a worse kind.

\* The Telescopes proper for observing the Eclipses of Jupiter's Satellites are, common refracting Telescopes, from 15 to 20 Feet; reflecting Telescopes of 18 Inches or two Feet; and Telescopes of Mr. Dollona's Construction, with two Object-glasses, from 5 to 10 Feet: or, which are still more convenient, those of  $3\frac{1}{2}$  Feet, which he has constructed with three Object-glasses.

An experienced Observer may count the Seconds by the Beats of the Clock, without an Assistant; but one unaccustomed to these Observations, by attending to the Time and Satellite together, will be liable to a Mistake in one, or to Inaccuracy in the other.

If it is the apparent Time of an Immersion, or Emerson, that is observed, (which is found by setting the Clock to Twelve, when the Sun's Center is on the Meridian) it must be corrected by the Equation answering to the Time of the observed Eclipse, to reduce it to mean Time, to which Eclipses are calculated in the Almanack; and the Difference between the calculated Time at Greenwich and the equated observed Time of the Eclipse, converted into Degrees and Minutes, will be the Longitude from Greenwich. To convert Time into Degrees and Minutes of a Circle: for every Hour allow 15 Degrees; for every Minute of Time over the Hours, 15 Minutes of a Degree; and for every Second of Time over the Minutes, 15 Seconds of a Degree.

Emersions are visible only from the Time of Jupiter's Opposition to the Sun, to the Time of his Conjunction: during which Time, he rises after the Sun sets. Immersions are visible only from his Conjunction to his Opposition; during which Time, he rises before the Sun sets.

To find when Jupiter is above the Horizon, and to distinguish him amongst the Stars; the easiest and readiest Way is by a celestial Globe, as follows. Find his Place in the Ecliptic from an Ephemeris for the Day proposed; also his Latitude, or Distance South or North of the Ecliptic; mark that Distance on the Globe on a Circle of celestial Latitude passing through his Place in the Ecliptic. Rectify the Globe for the Latitude of the Place you are in, and for the Sun's Place at Noon, and set the Hour Index to 12; turn the Globe till the Hour Index points out the Hour proposed; lay a Quadrant of Altitude from the Zenith over the Point marked on the Globe for his Place, and it will shew his Height above the Horizon, and likewise his Azimuth at that Hour: allow for the Variation of the magnetic Needle, and you will have his Bearing by the Compas. By the Altitude and Bearing Jupiter will be easily found among the Stars, for he is the largest to Appearance of the Stars and Planets, except Venus, who is likewise a little brighter than he.

Thought

Though an Eclipse of Jupiter's first Satellite is much more to be relied on than the rest, because the Irregularities of its Motion are better known; yet Eclipses of the other Satellites ought not to be neglected when there is an Opportunity of observing them: because such Eclipses may be observed in some other Place; by which the Difference of Longitude of these two Places will be as exactly known as by an Eclipse of the first Satellite; and if the Observations and Telescopes are good, is more to be depended on than any Calculation.

If the Part of Jupiter's Disk, and the Distance from it, at which the Emersion of a Satellite will happen, are known, that seems preferable to an Immersion; because the Appearance of any luminous body is more instantaneous than its Disappearance.

To find the Longitude of a Place by an Eclipse of the Moon; a Meridian Line must be first found; then a Clock, or Watch, regulated to keep equal Time, then set to mean or true Time by the Sun in the Meridian, as before directed. Then the precise Time noted by a Clock, or Watch, when the Moon's eastern Limb touches the Earth's Shadow; also the Time when the Shadow touches some remarkable Spots in the Moon's Disc; and when the Moon's western Edge touches the Shadow at its going off, when the Eclipse is just over: these several Times, compared with corresponding Observations in other Places, or with a Calculation, will give the Longitude between these Places. But Eclipses of the Moon are not to be relied on nearer than to three or four Minutes of Time; chiefly from the Difficulty of distinguishing the Edge of the Shadow precisely: and therefore Eclipses of Jupiter's Satellites are to be preferred to lunar Eclipses, for determining the Longitude of any Place.

## SECTION II.

*How to continue a Meridian Line through a Kingdom, or large Country.*

### PLATE III. FIG. IX.

First chuse some remarkable Place from whence to begin the Meridian, as a noted Observatory, Church, Castle, Hill, &c: if it is near the Middle of the

the Kingdom, so that the Meridian may be continued northward and southward from it, it is the better. Then provide a skilful and faithful Assistant, with two Sets of Instruments, one Set for yourself, and one for him; and six Men, at least, to attend each. Provide also two convenient Tents, one for lodging you, and your Attendants and Instruments, the other for him, his Attendants and Instruments. The Tents will serve likewise for Signals, and for making Observations in, and should therefore be so contrived, that a Breadth, or two, of the Cloth at each End, from Top to Bottom, may be laid open occasionally for observing the Stars. The necessary Instruments are, two Sextants of two Feet Radius, graduated so as to be capable of taking horizontal Angles to Parts of a Minute; with a spirit Level, and two good Telescopes to each, with Cross-wires in the common Focus of the Glasses, and a steady Stand to support the Instruments horizontally: two good magnetic Needles with Sights; four Plumb-lines, each six Feet long, and two Stands for them, one to support two together, about four Feet asunder; six straight spiked Poles, each twenty Feet long, with black and white Flags for their Tops, to render them conspicuous, as they happen to bear on the Sky, or on dark Land; and a Line drawn on one square Side of each, right up and down, for a Plumb-line to be applied to, in order to set them perpendicular: if the Plumb-line is hung in a Box, to screen it from the Wind when applied, it would be an Advantage. A Borer, or some such Instrument, for readily making a Hole in hard Ground for the Poles to stand in firmly: in Wind, these Poles will need to be supported by Ropes from the Top, like a Tent-pole, or by three spreading Legs like a Theodolite: twenty spiked Poles, each six Feet long; a measuring Pole, or Chain, with Reels and Lines on them, for measuring a straight Line on the Ground.

At the Place where the Meridian is to begin, find an exact Meridian Line, by one of the most accurate Methods explained in Part 1. (beginning Page 47.) mark its Extremity by a long Pole set perpendicular in the Ground, and a Flag at Top. Continue the Direction of that Part of the Meridian farther, if necessary, by taking that Pole, and the Pole or Object you began at in a Line; or by setting other upright Poles in the same Direction, till you come to the Summit of a Hill, or Eminence, where there is an extensive View beyond it; there mark a Point in the continued Meridian, by a Pin driven in the Ground, and pitch your Tent over it, so that the two

Tent-poles may stand parallel to the Meridian, at about the Distance of a Foot from it, and not hinder your seeing any Objects that are in the Continuation of it. Within the Tent (that your Observation may not be disturbed by the Wind) hang two Plumb-lines on their Stand; and when at Rest, make them hang exactly in the Meridian backward, or toward the Beginning of the Meridian, by cutting the Object you began at, and one or more of the Poles which continued the Meridian. Stand next behind the other Plumb-line, and, when both are at Rest, observe if they cut any small remarkable Part of a distant Object which may be precisely known when you go to the Object; if so, then is that Part of the Object in the Meridian; which may be continued from it as before, by taking it and the Middle of the Tent in a Line, till you come to some neighbouring Hill, or Eminence, convenient for pitching the Tent on, and making Observations. But if the Plumb-lines do not cut such a small distinguishable Part of an Object (which indeed will very seldom happen) then proceed in the following Manner.

Find out some Hill, or Eminence, on one Side of the Direction of the Plumb-lines (the nearer to it the better) and so far off as that your Assistant, at that Eminence, may see your Tent, and you may see his, when it is pitched on it: if they can just be perceived by the naked Eye, it is sufficient; for then they will be distinctly seen through the Telescopes on the Sextants. Let your Assistant take the Bearing of the Hill, and of the Direction of the Plumb-lines with the magnetic Needle, and note them down, and these will serve to direct him nearly to the Hill and Meridian: send your Assistant to the Hill proposed, with his Tent, Servants and Instruments, and let him pitch his Tent on a convenient Part of the Hill, or the neighbourhood of it, where he can see your Tent, and you his; and, at the Door of his Tent, set up a long Pole, with a conspicuous Flag at the Top of it, for a Signal to be observed through your Telescope: set you up such another Signal in the Direction of the Meridian, near the Door of your Tent, to be observed by him. As soon as his Signal is up, set your Sextant level, with the Center of it at the Place of the Signal, and take the Angle which his Signal makes with the Meridian, or Direction of the Plumb-lines: write it down, and send a Man with it to the Assistant; when you have done,

done, set up your Signal in its Place, with another Flag above the former, to signify that you have finished your Observation; and let them remain so till your Assistant's Tent is taken down.

When the Assistant has set up his Signal, then he must go as near the Direction of the Meridian as he can, by its magnetic Bearing which was taken at the Principal's Tent, and, in that Neighbourhood, find out some Eminence, or convenient Place, for the next Station, where his own Tent and the Principal's can be seen: there let him set up a large Pole and Flag, and two, or more, smaller, in a straight Line between it and the Signal at his Tent. Then, at his Tent-signal, let him take the Angle with the Sextant, formed by the Line of Poles crossing the Meridian and the Direction of the Principal's Signal, and write that down. This Angle, and that sent by the Principal, added together, and subtracted from  $180^{\circ}$  will leave the third Angle of the Triangle. Let him keep in the Direction of the Line of Poles, till he finds by the magnetic Bearing that he is in, or near the Meridian; then let him set up the Sextant, and observe the Angle between the two Tent-signals; if it is equal to the forementioned Remainder, the Center of the Sextant is in the continued Meridian: if the Angle observed is greater than that Remainder, the continued Meridian is farther from the Assistant's Tent than the Center of the Instrument; if the observed Angle is less, the Continuation of the Meridian is between the Instrument and the Assistant's Tent, and the Sextant must be shifted either one Way or the other, but always in the Direction of the Line of Poles, till this observed Angle, and the two formerly observed at the Tents, make exactly two Right-angles; then the Point directly below the Center of the Sextant is to be marked by a Pin driven into the Ground, and that will be another Point in the Meridian continued from the first Station. If this Place is not convenient for a new Station, he must continue the Direction of the Meridian by Poles till he comes to an Eminence, or Place more convenient; there he is to mark a Point with a Pole, and setting up another perpendicular Pole at a Distance, in a Line with the Principal's Signal, let him order his Tent to be taken down, as a Signal for the Principal to take down his, and to come away with it and with all his other Things. Then, over the Direction of the Meridian found by the Assistant the Principal must pitch his Tent, hang two Plumb-lines within it on the Meridian, and proceed as before to continue it farther.

If a Town, Wood, Building, high Wall, or other Obstacle, shall happen in the Way, so as to hinder the Continuation of the meridian Line, as the Wood X in Plate III. Fig. 9. then, before you reach such Obstacle, on a smooth Plane mark off with the Sextant A B a Perpendicular to the Meridian, and measure in that Line as great a Length as will extend beyond the Obstacle, so that the Perpendicular B C raised on the End of the Measurement may fall on one Side of the Obstacle; and continue B C, if necessary, with upright Poles, till you find a smooth Plane extending perpendicularly cross the Direction of the Meridian; on that Plane make C D perpendicular to B C, and measure from C to D a Distance equal to A B, and D will be a Point in the Meridian continued from A. Then a Perpendicular D E, on the Line C D, will be a Continuation of the original Meridian: by upright Poles placed in these two Points, and other Poles in their Direction, it may be continued farther at Pleasure.

If A B, the Perpendicular, cannot be measured exactly, because of the Unevenness of the Ground, mark a Line in another Direction, where the Ground is level, as A F, or A G, and take the Angle M A F, or M A G; at F, make the Angle A F C = M A F; and F C will be parallel to M A; or at G, if that is the Direction chosen, make the Angle A G C = M A G, and G C will be parallel to M A. Then in the right-angled Triangle A B F, or A B G, having the Hypotenuse A F, or A G; and, in the first, the Angle A F B, equal to the Supplement of A F C, or, in the other, the Angle A G B = M A G, thence, by Protraction, or Calculation, find the Perpendicular A B, and note it down. At C mark C D a Perpendicular to C B; on it measure the Length of A B, and it will terminate in D, a Point in the Meridian M A continued.

If a Perpendicular C D cannot be measured exactly, then mark out a Line on either Side of it where the Ground is level, as C H, or C I; take the Angle B C H, or B C I, according to the Direction assumed, and from thence, having the Length of the Perpendicular C D = A B, and the Angle D C H = B C H - 90°, or the Angle D C I = 90 - B C I, the Hypotenuse C H, or C I, may be found: measure that on the Ground, and it will terminate in H, or I, a Point in the Meridian M A continued. At H, on C H, make the Angle C H E = M A F; or, at the Point I, the Angle C I E = M A G, and E will be another Point in the Meridian M A continued.

nued. Place perpendicular Poles in these two Points, and by them continue the Meridian to some Eminence where a distant View foreward may be had; there pitch your Tent over the meridian Line, hang up the Plumb-lines in it, and examine your Meridian by the Stars before you proceed farther, in Case the measuring should have occasioned any Inaccuracy; correct the Error, if there is any, and then proceed as before, by taking the magnetic Bearing of the Meridian, and of some distant Eminence on one Side of it, and sending your Assistant to pitch his Tent on that Eminence, in order to make the necessary Observations there, and find out the Direction of the Meridian; and continue the same Process till you reach the End of the Territory proposed.

The Foundation of the foregoing Operations, in getting over Obstructions, is so obvious from the Properties of parallel Lines, that it needs no Explanation to one who is acquainted with the first Book of Euclid's Elements. It may however be proper to observe, that, if the Measurements are not made on smooth Planes, with great Care, and the Angles taken and set off on the Ground with great Exactness, it may make a sensible Deviation in the Meridian: and that therefore, after such an Operation, it will generally be expedient to find the Meridian anew by the Stars, and proceed to continue the Direction of it as before. But whether such Obstructions are met with or not, it will be prudent, every 80, or 100 Miles, to examine the Meridian by the Stars, that if any small Inaccuracy has arisen, it may not affect the subsequent Parts of it.

### SECTION III.

#### *How to trace a Parallel of Latitude through a Kingdom.*

Provide a good portable Instrument for taking the meridian Altitude of the Sun, or a Star, to Parts of a Minute. A Sextant of three Feet Radius, with a convenient Stand to support it, may be sufficient.

At the Place whose Parallel is to be traced, find the Latitude, and an exact meridian Line: set off, by Poles placed upright in the Ground, a Perpendicular to the Meridian, and continue that Line about one Degree of Longitude,

Longitude, (suppose 30 or 40 Miles) till you come to a level Plane extending about one Mile northward, or toward the elevated Pole; there find the Latitude and a meridian Line: on that meridian Line continued through the Plane, measure a Distance equal to the Difference of the first and last Latitude (which will be less than one Mile) and it will terminate in a Point of the Parallel required. Another Point in it may be marked about the Middle of the Perpendicular to the first Meridian: for the Tangent of half a Degree falls very little without the Arc.

At the first-mentioned Point in the Parallel, mark off with Poles a Perpendicular to the Meridian as before; continue it 30 or 40 Miles forward till you find another level Plane; there observe the Latitude and find a meridian Line: on that Meridian measure the Difference of the first and last Latitudes, and that Measurement will terminate in another Point of the Parallel: and the Middle of the Perpendicular to the former Meridian will be a fourth Point. Proceed in the same Manner to find other Points as far eastward, or westward as is necessary, and straight Lines on the Ground joining these several Points will mark out on the Ground the Parallel of Latitude of the Place required; or more strictly a Polygon not sensibly differing from it.

The Polygon will be still nearer a Circle, if a Perpendicular is continued East and West from each of the Meridians, and two Points of the Parallel are marked in each of these Perpendiculars, one on one Side of the crossing Meridian, the other on the other Side, and about one third of a degree of Longitude from it.

If it shall happen that a smooth Plane is not to be had for measuring the Difference of the two Latitudes along the Meridian, measure a straight Line in any other Direction making an Angle with the Meridian; find that Angle by the Sextant, and you will then have two Sides of a Triangle (the Difference of Latitude and the measured Line) and the Angle comprehended between them, to find another Angle, which second Angle observed and marked off by a Pole in the Ground will determine the Direction of the third Side; and where it intersects the Meridian will be a Point in the Parallel of Latitude sought.

The

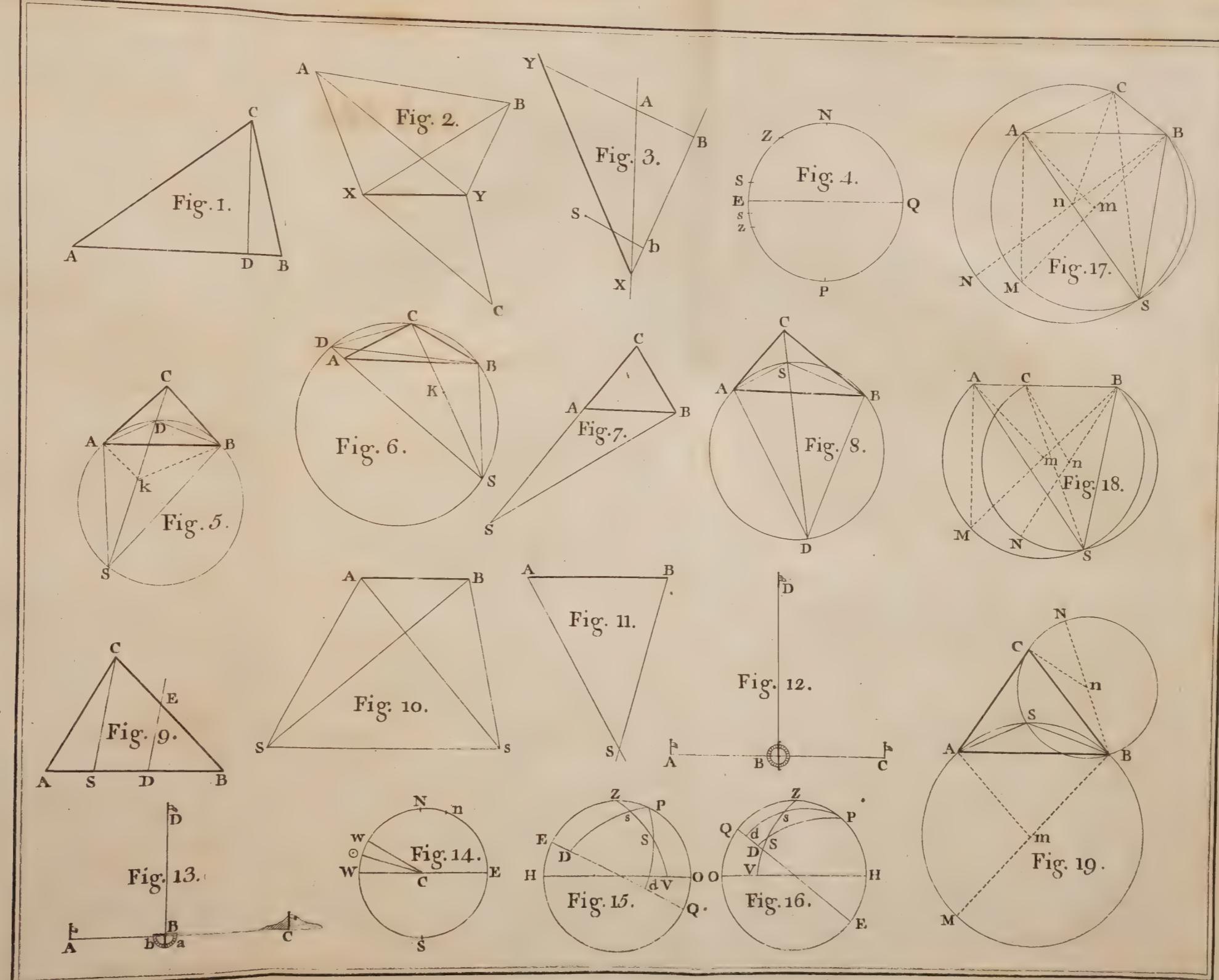
The Exactness of this Operation depends chiefly on the Accuracy of the Instrument with which the several Latitudes are taken: the Sextant, therefore should be carefully adjusted immediately before each Observation.

THE END.

## E R R A T A.

Page 19. Line 22, *after towards S, read as at k.*—p. 23. l. 8, *after A S B, read* (Fig. 17 and 18.)—p. 27. l. 1, *for Pales, read Poles.*—p. 34. l. 16, *for Horizon, read Horizons.*—p. 36. l. 16, *after true Angle, read To prevent it, the Index should be of a Substance not apt to bend, and as broad as the Sextant can conveniently admit of; it should move quite easy round the Center; but by means of a spring and screw should compress the Arch of the Instrument with requisite Firmness, and so as to slide smoothly along it when moved.*—p. 55. l. *last but two, for d' Q, read D E.*—p. 59. l. 22, *for wuond, read wound.*—p. 88. l. 16, *for Objed, read Objects.*

PLATE. I.



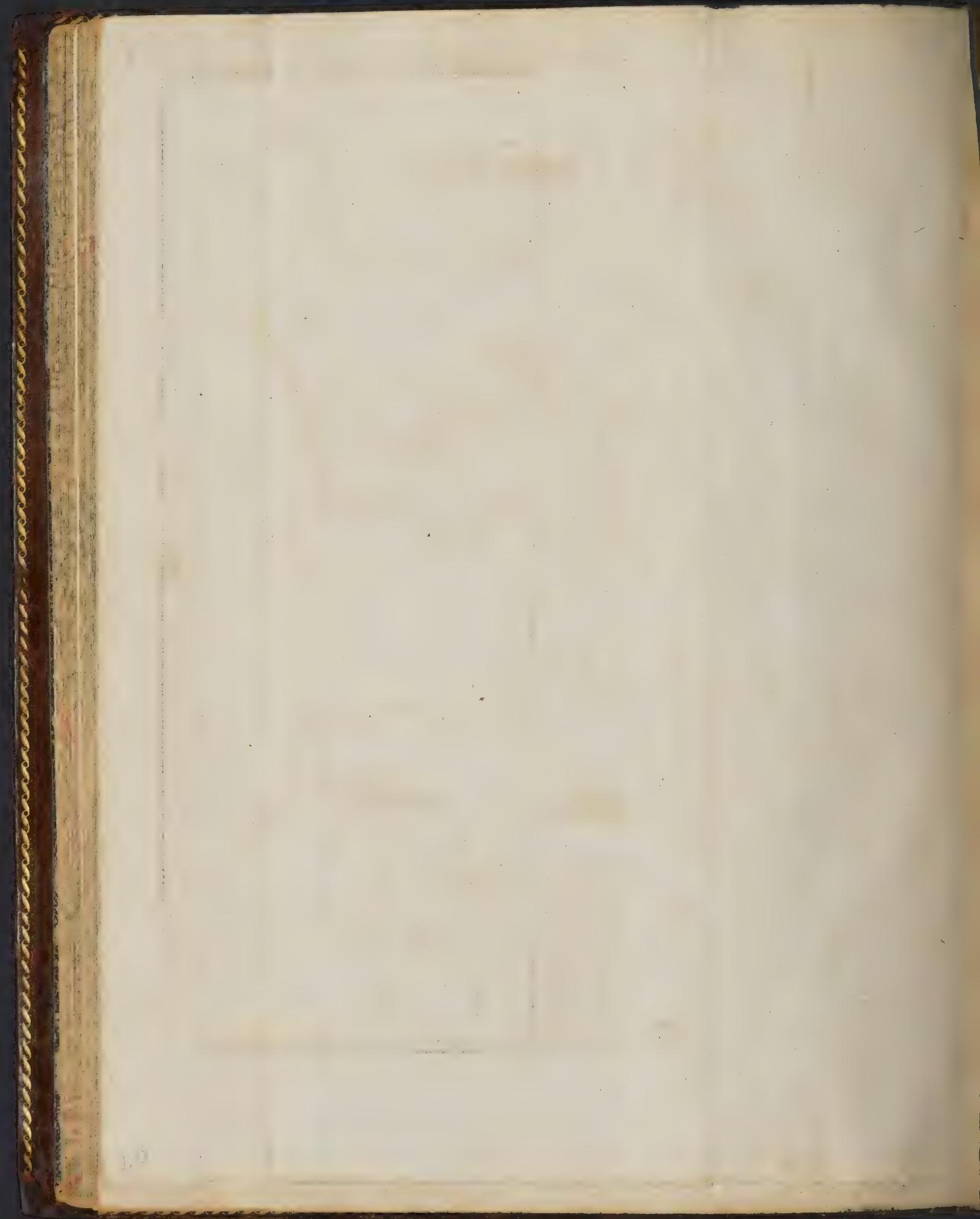


PLATE II.

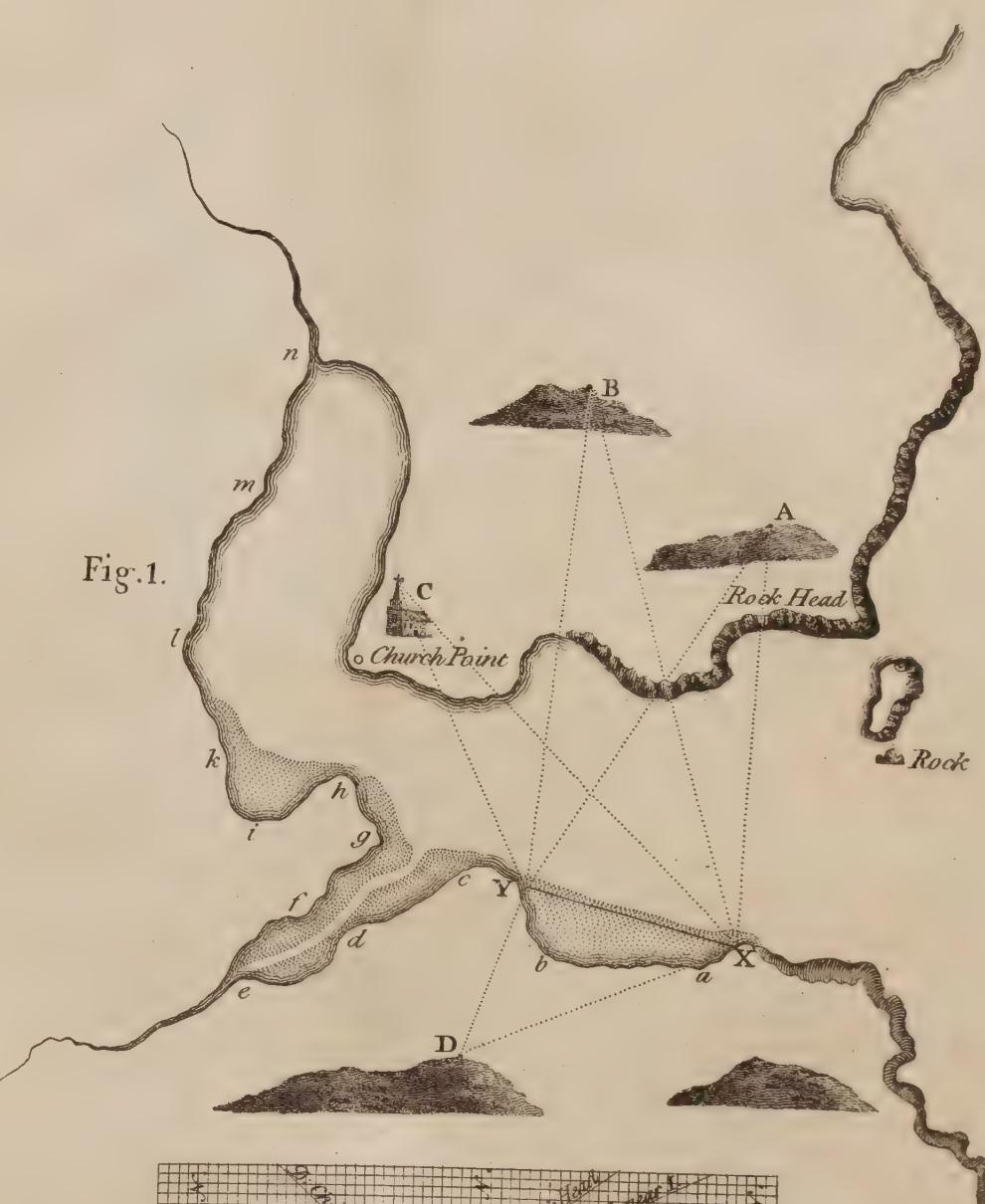


Fig.

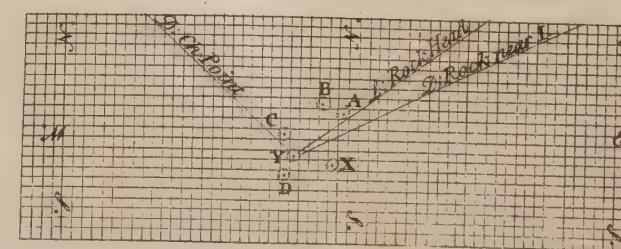


Fig. 2

St. Louis, Mo.

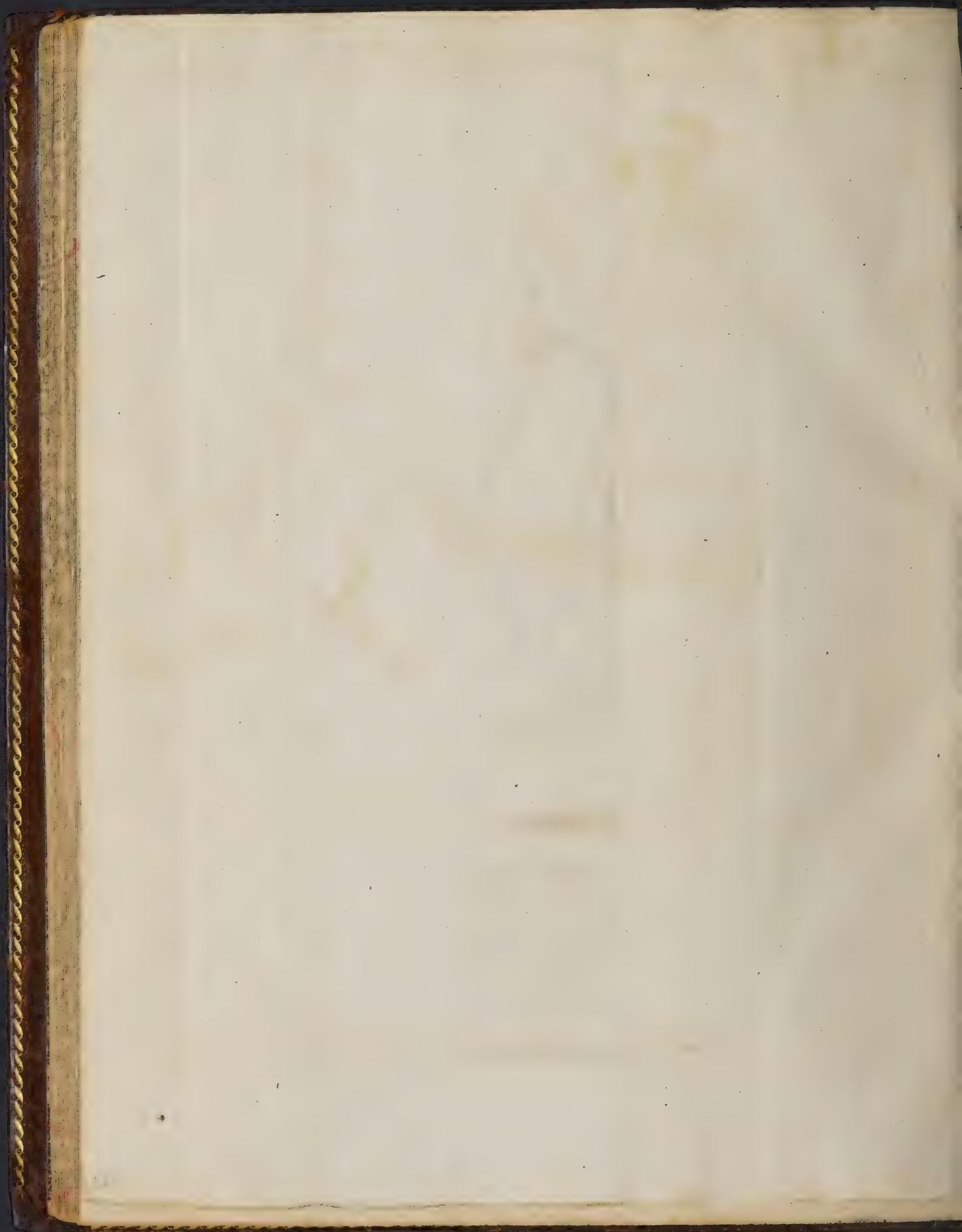
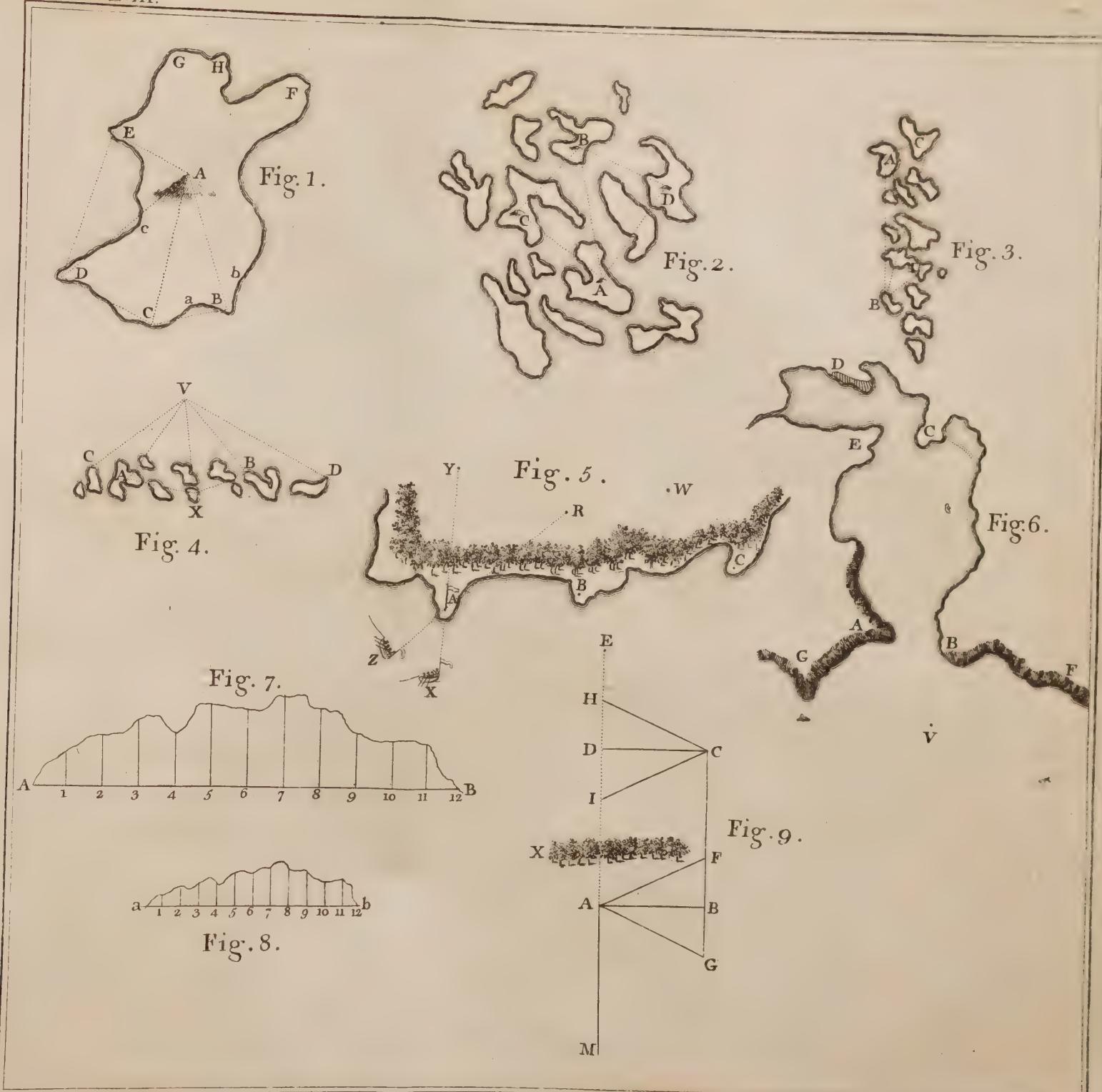


PLATE III.



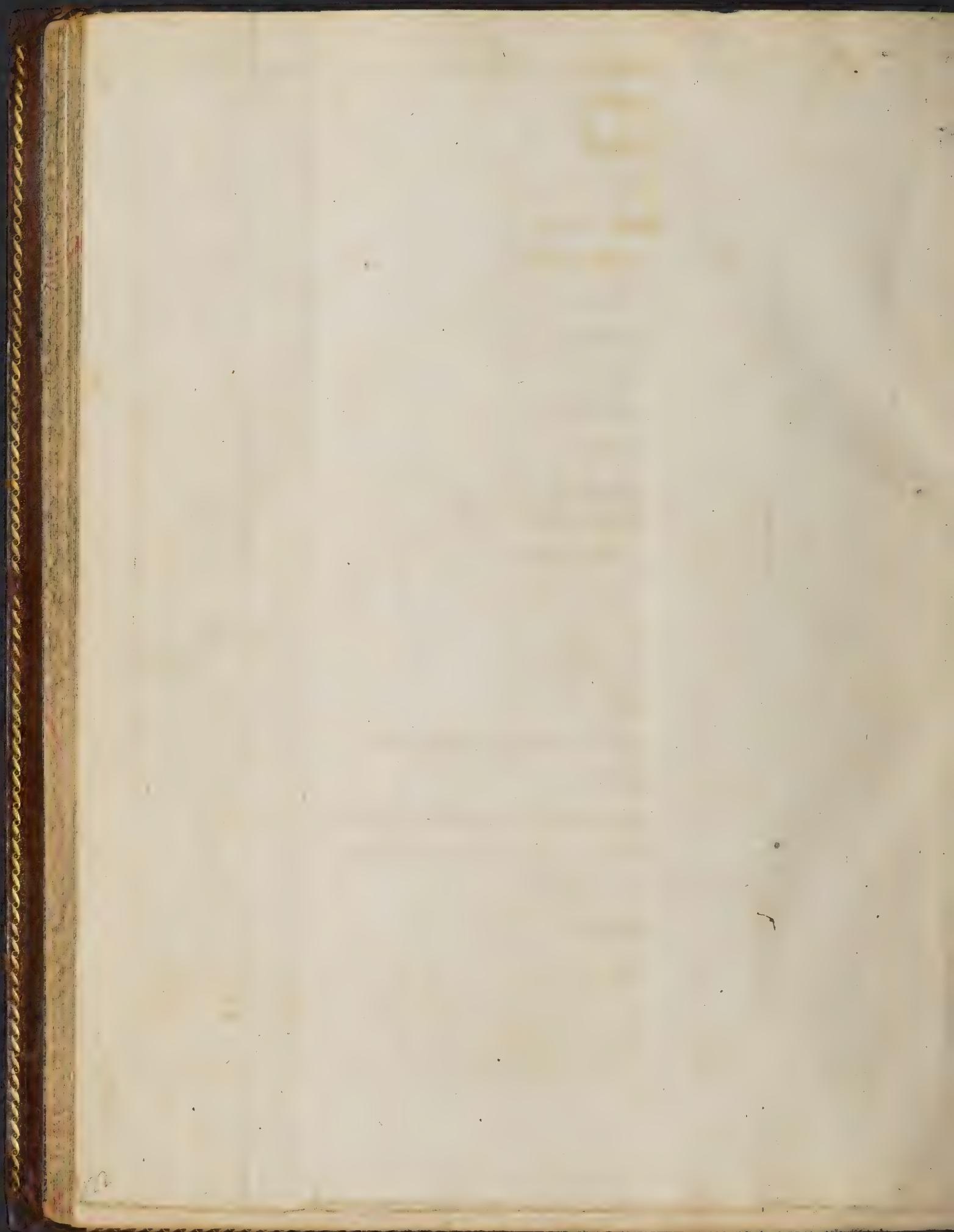


PLATE. IV.

*Marks and Diversities to be represented in Draughts.*

-  Rocky Cliffs.
-  Grassy, or Clay Cliffs.
-  Low Rocky Shores.
-  Low Sandy Shores.
-  Sandy Soil near the Shore.
-  Rocks always above Water.
-  Rocks that dry and are covered daily.
-  Rocks that dry with Spring-tide only.
-  Rocks always below Water.
-  Sand-banks that dry and are covered daily.
-  Sand-banks dry with Spring-tide only.
-  Sand-banks always below Water.
-  The safest Anchorage.
-  Stopping-Places.
-  The best Channel.
-  An Eddy.
-  Overfalls; or rough, breaking Seas.
-  Whirlpools.
-  Leading-marks, for avoiding Rocks and Shoals.
-  The Direction of the Stream of Flood.
-  Beacons, or Perches.
-  A Can and Cask Buoy.

Below the Soundings, (r) signifies rocky Ground; (c) Clay; (m) Mud, or Ooze; (g) Gravel; (s) Sand; (sh) Shells, or Shell Sand; (st) large Stones; (co) small Corral; (sg) Sand and Gravel; (ssh) Sand and Shells. The numeral Letters shew the Time of High-water on the full & change Days of the Moon. Perpendicular Rise of ordinary Spring-tide      Feet; of Neap-tide      Feet.

